

# Introduction to Quantum Field Theory (Fall, 2016)

Download Lecture Notes (homeworks) in the instructor's English homepage:  
<http://physics.whu.edu.cn/en/?q=node/112>

## 1 Schedule, Instructor and Teaching Assistant

*Time and Place: Tuesday, Class 6-8, I-3-404*

Instructor: Yong Zhang.

- Office: 5-516, at the new building of the School of Physics and Technology;
- Email: [yong\\_zhang@whu.edu.cn](mailto:yong_zhang@whu.edu.cn).

Teaching Assistant: (None).

## 2 Lecture Notes

\* Yong Zhang (Wuhan): the fourth version, online lecture notes on QFT.

*The teaching plan of this course is made in accordance with the main reference.*

## 3 Main Reference

\* Michael Luke (Toronto): online lecture notes on QFT.

## 4 References

- \* Peskin & Schroeder (Stanford): An Introduction to Quantum Field Theory;
- \* David Tong (Cambridge): online lecture notes on QFT;
- \* Sidney Coleman (Harvard): online lecture notes;
- \* Mandl & Shaw: Quantum Field Theory (Second Edition).

## 5 Homeworks

Peskin & Schroeder's Problem Sets or Luke's Problem Sets or Tong's Problem Sets.

## 6 Research Projects

Ask the instructor for it.

## 7 Evaluation

Homeworks(50%) + Final Exam(50%).

Final Exam      ↗ Open test.  
                    ↗ Open to Classnotes and Homeworks.

*Remark: do homework=do research*

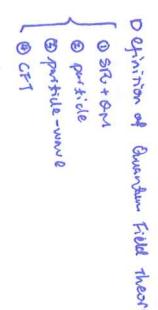
$$\left\{ \begin{array}{l} \text{Perspective: Focus on a small problem;} \\ \text{Persistence: Solve your problem completely;} \\ \text{Patience: Write down your solution step by step;} \\ \text{Power: Quote your references carefully.} \end{array} \right.$$



lec 1. Introduction to QFT  
 lec 2. Introduction to Feynman diagram (key point of this course)  
 lec 3. Classical Field Theory

What about the final exam in Quantum Information & Computation?

Open fest.  
Results are terrible.



I.1. About the Course (4.00-QFT - course - Introduction - 2016.2014)

II A Class meeting Tues, class 6-8, 1-3-404

II B Instructor Yong Zhang; office: 5-516 New building; Email address: Yong-Zhang@whu.edu.cn

No teaching assistant; the answer key notes on QFT;

III C Lecture notes, the Fourier version rec. notes on QFT;  
 III D Main reference Michael Luke (Curran), online notes on QFT;  
 III E Performance Peskin & Schroeder (Standard), Introduction to QFT;  
 III F Homeworks 3 homeworks; taken from Luke's problem sets; Tony's problem sets; Peskin & Schroeder's;

Do homework = Do research  
 perspective: focus on small problem  
 persistence; solve your problem completely  
 patience; write down your solution step by step (check your solution's last way)  
 power; self-confidence, guess your reference carefully

II 1.9 Evaluation: 3 homeworks ( $3 \times 3 = 9$  problems)  $\rightarrow 50\%$   
 final exam  $\rightarrow 50\%$   
 open fest

Remark: Homeworks & final exams will be much simpler than Homeworks & final exam in QIC.

Remark: All the solutions to final exam will be discussed in detail in class.  
 Taking classnotes and doing homeworks are crucial for doing final exam  
 III Research projects: Ask instructor for it.

1.2. Overview of modern theoretical physics  
 what's the place of QFT in MTP. Historical conceptual development of QFT.

particle mechanics (Newtonian mechanics)  $\rightarrow$  classical mechanics  $\rightarrow$  quantum mechanics

revision  $\rightarrow$  Electromagnetic field

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P1-P2

1. Modern Theoretical Physics, branches

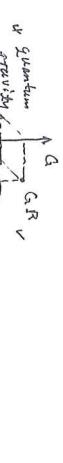
2. QFT  $\xrightarrow{\text{QFT}}$  Continuum limit

Quantization

Quantum mechanics } quantum field theory } (string theory)  
 special relativity } general relativity } unification theory

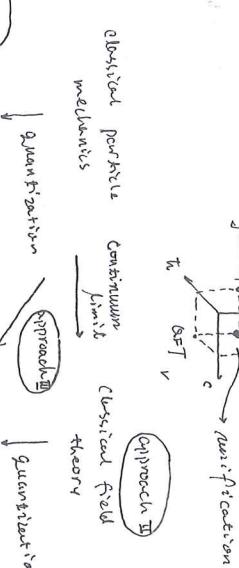
(gravity)  $\rightarrow$  gravitational constant  
 $\rightarrow$  Planck Constant  
 (quantum mechanics)

< (light speed)  
 (special relativity)



Approach I: Classical particle continuum limit  
 Approach II: Quantum mechanics  $\xrightarrow{\text{Quantization}}$  Classical field theory  
 Approach III: Quantum field theory

what is quantization procedure



Approach I: Classical particle continuum limit  
 Approach II: Quantum mechanics  $\xrightarrow{\text{Quantization}}$  Classical field theory  
 Approach III: Quantum field theory

what is quantization procedure

Quantization / Canonical Quantization (we know its name but don't know its name)  
 Path quantization (we know its name but don't know its name)

Def 1. QFT = Special Relativity + Quantum Mechanics  
 (use models)      (dilutes)  
 Note: QFT means Relativistic Quantum Field Theory  
 Relativistic Quantum Mechanics (wrong name) = QFT

Def 2. QFT = Quantum many-particle system with unfixed particle number  
 (unfixed mass-energy relation)  
 (about particles)

Remark: At high energies, with special relativity. Particle can be created and annihilated,  
 so the particle number is not conserved, which is described by relativistic QFT;  
 (destructive)

Remark: At low energies, for example, condensed matter physics. Particle number is a  
 conserved quantity, which is described by non-relativistic QFT.

Non-relativistic QFT  $\rightsquigarrow$   
 = Quantum many-particle theory with fixed particle number.  
 (about field)

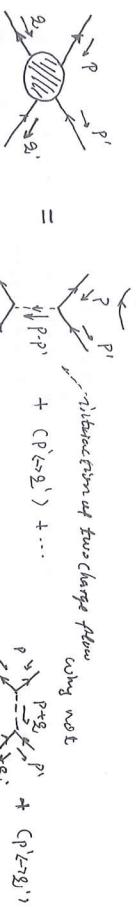
Def 3. QFT = Quantization of classical field theory = quantization from special relativity  
 satisfying the constraints of non-relativistic QFT in  
 e.g. Quantum Electrodynamics (QED) = Relativistic Quantization of Electrodynamics.

Def 4. QFT  $\Rightarrow$  is a natural language of describing the particle-wave duality;  
 (use don't understand the duality, therefore, the root)  
 but use QFT, we can describe the duality  
 (classical? field)

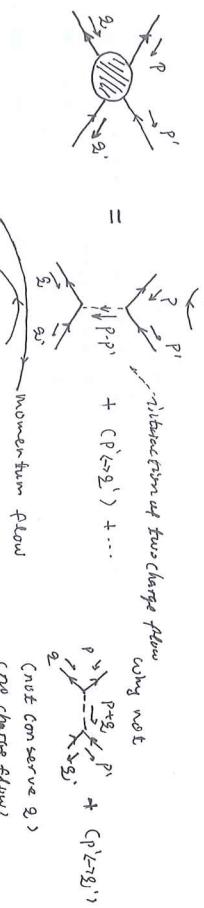
each fundamental particle is a derived object of quantization of the associated field.



e.g. 3. Two Dirac- "Nucleon" scattering  
 $\bar{n}^*(p) + \bar{n}^*(q) \rightarrow \bar{n}^*(p') + \bar{n}^*(q')$ ,  $p+q = p'+q'$



e.g. 4. "Nucleon" - Anti- "Nucleon" scattering :  $n^*(p) + \bar{n}^*(q) \rightarrow n^*(p') + \bar{n}^*(q')$ ,  $p+q = p'+q'$



charge flow	momentum flow
Same direction "Nucleon"	diff. direction / opposite "Anti- Nucleon"

In this sense, QFT is very simple; we just draw the FDs without understanding the QFT.

e.g. 4. "Nucleon" - Anti- "Nucleon" Scattering :  $n^*(p) + \bar{n}^*(q) \rightarrow n^*(p') + \bar{n}^*(q')$ ,  $p+q = p'+q'$

Review: Models  
 Lec. III Introduction to FDs  
 Lec. IV Introduction to Classical Field Theory (later)

Review: Particle (Fields)  
 The pseudo-particle interaction

$p = (p^\mu) = (p^0, \vec{p}) = (\epsilon, \vec{p})$ ,  $\tau_3$ : spin  $\left\{ \begin{array}{l} = \frac{1}{2}, \uparrow \\ = \frac{1}{2}, \downarrow \end{array} \right.$   
 $E$ : polarization vector (4-dim)  $E = (E^\mu) = (e^0, \vec{e})$ ,  $(e^0, \vec{e}_1, \vec{e}_2, \vec{e}_3)$

pseudo-nucleon anti-pseudo nucleon meson  
 $"n"(\eta)$   $"\bar{n}"(c_2)$   
 $\phi(q)$

Complex scalar field  $\psi(q)$ ,  $\bar{\psi}(q)$ , meson  
 nucleon anti-nucleon meson  
 $N(p, \tau)$   $\bar{N}(q, \tau)$   
 $\phi(q)$ , spin

Dirac field :  $\psi(q)$ , Dirac conjugation :  $\bar{\psi}(q)$   
 $(\bar{\psi}\psi)^* = \bar{\psi}\psi$  (just complex conjugation)

Quantum Electrodynamics (QED)

electron position  $e^-(p, \tau)$ , photon

$\gamma(k, \epsilon)$ ,

Dirac field  $\psi(q)$ ,  $\bar{\psi}(q)$

gauge potential  $A_\mu(x)$

Review:

$\overrightarrow{p+q}$  ( $\Leftrightarrow \overrightarrow{\phi(p+q)}$ )

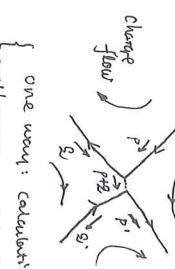
→ momentum arrow (flow)  
 → charge arrow  $\Rightarrow$  same direction  $\Leftrightarrow$  "n( $\eta$ )"

$\overrightarrow{q}$  ( $\Leftrightarrow \overrightarrow{\phi(q)}$ )  
 $\overrightarrow{p}$  ( $\Leftrightarrow \overrightarrow{\phi(p)}$ )

### III. 1. The pseudo-particle Interaction

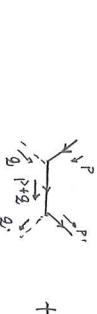
e.g. 4. "n( $\eta$ )" and  $\phi(q)$  scattering :

$$n^*(p) + \bar{n}^*(q) \rightarrow n^*(p') + \bar{n}^*(q'), \quad p+q = p'+q'$$

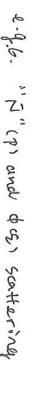


e.g. 5. "n( $\eta$ )" and  $\phi(q)$  scattering :  $n^*(p) + \phi(q) \rightarrow n^*(p') + \phi(q')$ , another way: derivation

one way: calculation derive FDs  
 another way: intuition

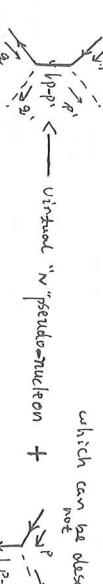


e.g. 6. "n( $\eta$ )" and  $\phi(q)$  scattering :  $n^*(p) + \phi(q) \rightarrow n^*(p') + \phi(q')$ , incoming  $\phi(q)$ , why not:

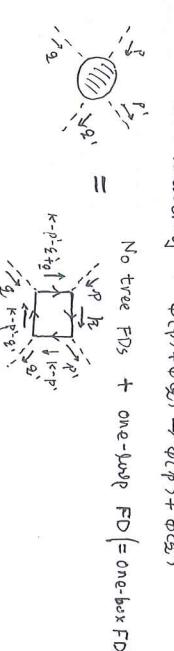


e.g. 7. "n( $\eta$ )" and  $\bar{n}^*(q)$  annihilation :  $n^*(p) + \bar{n}^*(q) \rightarrow \phi(p') + \phi(q')$ , particles are changed!

That's QFT, we can understand number of particles which can be described using QM.



e.g. 8. meson-meson scattering :  $\phi(p) + \phi(q) \rightarrow \phi(p') + \phi(q')$ , No tree FDs + one-loop FD (= one-box FD)



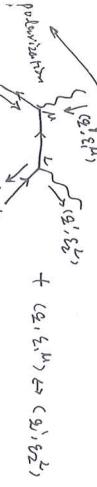
Virtual pseudo-nucleon ("nucleon")

flow

$\Updownarrow$  for loop

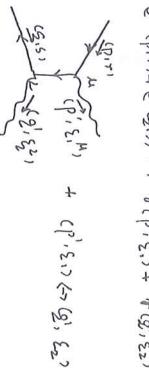


e.g. 6 Position - Photon Scattering (Compton Scattering)  $e^+(p_1, \epsilon_1) + \gamma(q_2, \epsilon_2) \rightarrow e^+(p'_1, \epsilon'_1) + \gamma(q'_2, \epsilon'_2)$



e.g. 7. Pair Annihilation

$$\bar{e}^-(p_1, \epsilon_1) + e^+(p_2, \epsilon_2) \rightarrow \tau(\rho_1, \xi_1) + \tau(\rho_2, \xi_2)$$



Lec IV Classical Field Theory  
Lec V Symmetries & Conservation Laws

Remark: Lec II & Lec III Feynman diagrams

- II & III. perturbative QFT
- 2. R-S model - Yukawa interaction
- 3. Yukawa interaction

#### 4. QED (Quantum electrodynamics)

e.g. 3. Decay of Boson into Fermion and Anti-Fermion

e.g. 2. Fermion - Fermion scattering

e.g. 4. Fermion - Boson scattering

e.g. 5. Anti-Fermion - Boson scattering

e.g. 6. Anti-Fermion - Fermion scattering

e.g. 7. Fermion - Fermion Annihilation

e.g. 8. Boson and Boson scattering

Remark 1: Charge  $\partial$ , distinguishing particle and anti-particle

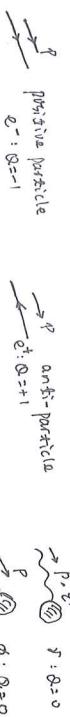
$$e^- : \partial = -1 \quad \bar{\nu} : \partial = +1$$

(some spin/mass)

$$proton : \partial = 1 \quad \bar{n} : \partial = -1$$

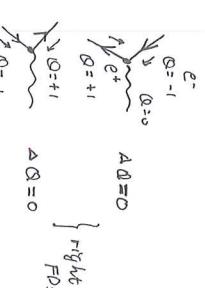
$$\bar{p} : \partial = 0 \quad \bar{\pi} : \partial = 0$$

(Special example, how to distinguish them?)



Feynman diagram are able to represent particle and anti-particle

Remark 2: How to draw FDS  $\rightarrow$  Charge conservation in (and) FDS  
At every interaction vertex, total charge  $\Delta \partial = 0$  (guide)



Remark 3: energy-momentum conservation



$\Phi(p_1, q_1) \rightarrow N(p_2, r) + \bar{N}(q_2, s)$  is allowed, but  $\gamma \rightarrow e^+ + e^-$  isn't allowed

$k = p + q$

$\Delta p = 0$

$k = p + q$

$\Delta p \neq 0$

Remark 4: Total angular-momentum conservation

$$\vec{J} = \vec{L} + \vec{S}$$

$$e^-(p_1, \epsilon_1) + e^-(q_2, \epsilon_2) \rightarrow e^-(p'_1, \epsilon'_1) + e^-(q'_2, \epsilon'_2)$$

$$p_1 q_2 = p'_1 q'_2$$

$$(L_1 + S_1) + (L_2 + S_2) = (L'_1 + S'_1) + (L'_2 + S'_2)$$

Remark 5: How to draw FDS

$$\bar{e}(p_1, \epsilon_1) \rightarrow \nu(p_2, \epsilon_2) + \bar{\nu}(q_2, \epsilon_2)$$

Interchange two Fermions: " - "

Interchange two Bosons: " + "

Step 1. draw incoming lines and outgoing lines

Step 2. draw interaction vertex

One vertex two vertices

Step 3. connect incoming/outgoing lines with vertices

$\rightarrow$  cone (lowest tree diagram)

Remark 6: Spin-statistics theorem and FDS

$$e^-(p_1, \epsilon_1) + e^-(q_2, \epsilon_2) \rightarrow e^-(p'_1, \epsilon'_1) + e^-(q'_2, \epsilon'_2)$$

Interchange two outgoing particles

" + " : FDS

" - " : Spin-statistics theorem

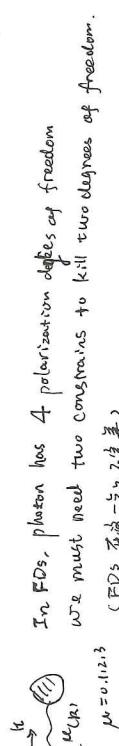
Feynman integrals

## (6) Virtual particles and Causality

Virtual particles can cause violation of causality, but who knows! (Virtual particles cannot be observed.)



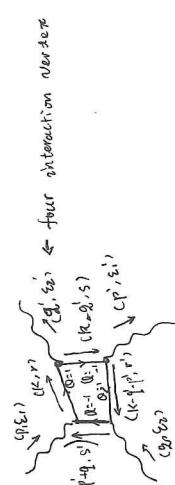
Remark 8: Photon = quantita of EM field, photon only allows two polarization degrees of freedom.



$\mu = 0, 1, 2, 3$  (FDs  $\vec{E} \vec{B} - \vec{B} \vec{E}$ )

Remark 9: Photo - photo scattering

$$\sigma(\rho, \varepsilon) + \tau(g, g_z) \rightarrow \tau(p, \varepsilon') + \tau(g, \varepsilon')$$



Remark 10: CPT invariance and FDs (need more knowledge)

## (7) Fourier-vector Calculations

① Three-vector Calculation  $\vec{r} = (x_1, y_1, z_1)$ ,  $\vec{p} = (p_1, p_1, p_3)$ ,  $\vec{D} = (D_1, D_1, D_3)$

$$\vec{r}^2 \equiv \vec{r} \cdot \vec{r} = x_1^2 + y_1^2 + z_1^2$$

$$\vec{p}^2 = p_1^2 + p_1^2 + p_3^2$$

$$D^2 = D_1^2 + D_1^2 + D_3^2 = \Delta$$

② Contravariant 4-vector

$$u_\mu = (ct, \vec{r}) \quad c: \text{light speed} \quad p^\mu = \frac{\partial}{\partial x^\mu} = \left( \frac{\partial}{\partial ct}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right)$$

$$= (ct, x, y, z)$$

③ Covariant 4-vector

$$x_\mu = (ct, -\vec{r}) \quad p_\mu = \frac{\partial}{\partial x^\mu} = \left( \frac{\partial}{\partial ct}, -\frac{\partial}{\partial x}, -\frac{\partial}{\partial y}, -\frac{\partial}{\partial z} \right)$$

(contravariant  $\eta_{\mu\nu}$   $\eta^{11} = -1$ )

④ Minkowski metric  $\eta = (\eta_{\mu\nu})$  : associates contravariant and covariant

Note: ( $\eta_{\mu\nu}$ ) has physical meaning, it means flat space-time without gravity

$$(\eta_{\mu\nu}) = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix}, \quad x_\mu = \eta_{\mu\nu} x^\nu$$

Note:  $\eta^{-1} = (\eta^{\mu\nu}) = \eta = (\eta_{\mu\nu})$  ( $\eta^{\mu\nu} \eta^{-1} = \delta_{\mu\nu}$ )

Note: Notation and convention makes GFT difficult. (different people use different notation)

Remark: Notation and convention makes GFT difficult. (different people use different notation)

$$\eta^{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad \eta_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\eta^{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \quad \eta_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

$$\eta^{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \quad \eta_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

But spacetime interval is unchanged  $c^2t^2 - x^2 - y^2 - z^2$ , and mass-energy relation holds.

Note:  $\rho_\mu = \eta_{\mu\nu} \rho^\nu$ ,  $\rho_{\mu\nu} = \eta_{\mu\nu} \rho^\nu$ ,  $\rho_\mu = \eta_{\mu\nu} \rho^\nu$ ,  $\partial_\mu = \eta_{\mu\nu} \partial^\nu$  (the same as  $x^\mu, x_\nu$ )

⑤ Inner product (scalar product) and invariant in special relativity

space time interval  $x^\mu x_\nu = \eta_{\mu\nu} x^\mu x^\nu = c^2 t^2 - \frac{x^2}{c^2} - \frac{y^2}{c^2} - \frac{z^2}{c^2} = c^2 t^2 - \frac{x^2}{c^2}$  (the definition of  $x^\mu$  is flexible, but all should make  $\rho^2$  invariant.)

mass-energy relation  $\rho^2 = \rho_\mu \rho^\mu = \eta_{\mu\nu} \rho^\mu \rho^\nu = (p_0)^2 - \vec{p}^2 c^2 = E^2 - \vec{p}^2 c^2 = (mc^2)^2$

Simple but abstract  $E^2 = (mc^2)^2 + \vec{p}^2 c^2$  (the second type of invariant.)

complex but physical  $E = \sqrt{(mc^2)^2 + \vec{p}^2 c^2} = mc^2 \sqrt{1 + \frac{\vec{p}^2}{m^2 c^2}}$  (practical energy)

$$\Delta E = \Delta m \cdot c^2$$

Remarks. In high-energy physics, particles can be created or annihilated, number is prefixed  $\otimes F T$

Remark 2.  $\rho^2 = (mc^2)^2 + \vec{p}^2 c^2 \rightarrow e^{-ict} + \vec{p} = p$   $\Rightarrow k^2 = p_1^2 + p_2^2$

Complex but physical  $\theta = 2mc^2 + \vec{p}^2 c^2$

Remark 3. Fourier transform

$Rx = k_0 x^0 - \vec{k} \cdot \vec{x}$

def:  $f(x) = \int \frac{d^4 k}{(2\pi)^4} f(k) e^{ikx}$ ,  $f(k) = \int d^4 x f(x) e^{-ikx}$  are (inverse) transform. in 4-vector language.

Dirac-delta function (def using integral)

1-dim  $\delta(x) = \begin{cases} +\infty, & x=0 \\ 0, & x \neq 0 \end{cases} \int_0^\infty \delta(x) dx = 1$

4-dim  $\delta^4(x) = \delta(x_0) \delta^3(x_1) \delta(x_2) = \int \frac{d^4 k}{(2\pi)^4} e^{ikx}$

Step function there is natural relationship between  $\delta(t)$  and  $\delta(x)$

$\theta(t) = \begin{cases} +1, & t > 0 \\ -1, & t < 0 \end{cases}$

Dirac-delta - delta function

$\delta_{\mu\nu} = \begin{cases} 1, & \mu = \nu \\ 0, & \mu \neq \nu \end{cases}$

Note: not preserved  $\delta_{\mu\nu}$  but preserved  $\delta^4_{\mu\nu}$ .

Summary: Lec I IV Remark on FDs Fourier-vector calculation

Lec II V Remark on FDs Fourier-vector calculation

Lec III VI Classical field theory

Momentum transformation, important for

momentum group and representation

Lec IV VII Symmetries and Conservation Laws

Space-time translation invariance

Space-time rotation invariance

CPT invariance and angular-momentum conservation

Recall what we have done

Topic 1. Conceptual development of SFT,  $\mathcal{SFT} = \mathcal{QFT} + \mathcal{CSR}$

Topic 2.  $\mathcal{FDS} \text{ of } \mathcal{SFT} = \text{Summarization of } \mathcal{CFT} = \text{Quantum many-particle with particle number conserved}$

Topic 3.  $\mathcal{PQFT} = \text{set of FDs}$

Topic 4. Example 1. pseudo-Yukawa: "N", "N",  $\phi$

Topic 5. Example 2. Yukawa: N, N,  $\phi$

Topic 6. Example 3. QED:  $e^-, e^+, \gamma$

Topic 7. How to understand and draw FDs?

energy-momentum conservation

charge conservation

angular momentum conservation

conservation, different particles

spin-statistics theorem

conservation, different particles

How to draw Feynman diagrams:  $\mathcal{FD} = \text{external lines} \setminus \text{Incoming particles} \setminus \text{Outgoing particles} \setminus \text{Vertices: Interaction}$

Remark (about the teaching schedule plan): only 3 constants of motion (lec notes will be taught this semester).

How to draw Feynman diagrams:

+ internal lines:

propagation of a virtual particle(s)

called Feynman propagator

e.g. QED

External lines

$(p_\mu, \epsilon_\mu)$  electron

$(p_\mu, \epsilon_\mu)$  photon

$(p_\mu, \epsilon_\mu)$  incoming

$(p_\mu, \epsilon_\mu)$  outgoing

$(p_\mu, \epsilon_\mu)$  electron

$(p_\mu, \epsilon_\mu)$  photon

$(p_\mu, \epsilon_\mu)$  incoming

$(p_\mu, \epsilon_\mu)$  outgoing

$(p_\mu, \epsilon_\mu)$  electron

$(p_\mu, \epsilon_\mu)$  photon

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$(p_\mu, \epsilon_\mu)$  photon

$(p_\mu, \epsilon_\mu)$  incoming

$(p_\mu, \epsilon_\mu)$  outgoing

$(p_\mu, \epsilon_\mu)$  electron

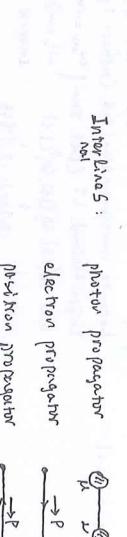
$(p_\mu, \epsilon_\mu)$  photon

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$(p_\mu, \epsilon_\mu)$  electron

$(p_\mu, \epsilon_\mu)$  photon



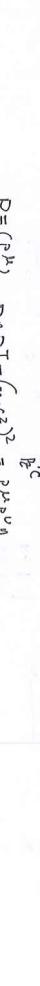
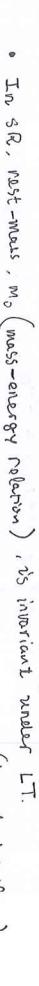
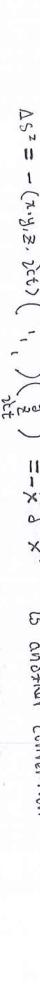
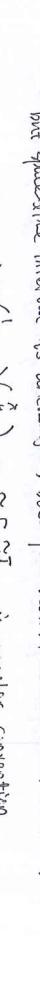
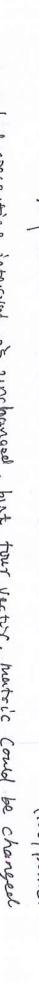
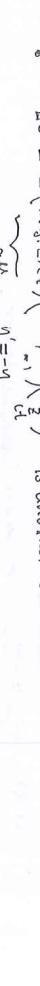
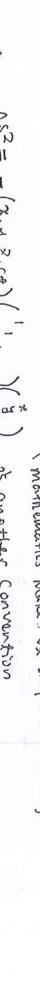
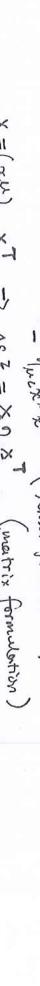
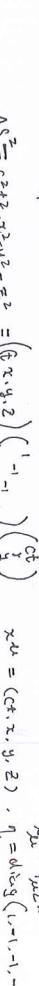
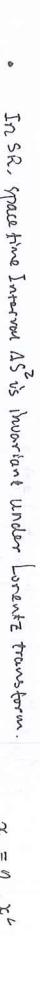
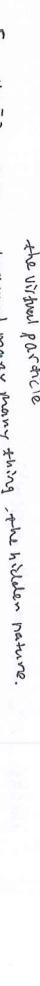
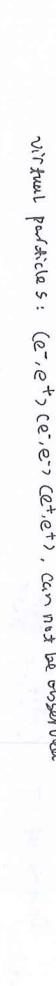
e.g. one-loop FPs for photon-polarization

the same loop, but three stories



In one loop, infrared (lines) momentum k is unconstrained, -> black & red,

virtual particles:  $(e^-, e^+)$ ,  $(\bar{e}, e^-)$ ,  $(e^+, e^+)$ , can not be observed



(top point)  $\gamma_{\mu\nu} = \eta_{\mu\nu} \gamma^\nu$   
 $\gamma_{\mu\nu} = \text{diag}(1, -1, -1, -1)$

$\Delta S^2 = c^2 t^2 - \vec{r}^2 - m^2 - \varepsilon^2 = (c^2, \vec{r}, m, \varepsilon)$

$\Delta S^2 = -(\vec{r} \cdot \vec{y}, \vec{z}, \vec{v}) \begin{pmatrix} 1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix} = -\vec{r} \cdot \vec{y} - \vec{z} \cdot \vec{v}$

In SR, rest-mass, no (mass-energy relation), is invariant under LT.

$$E^2 = p^2 c^2 + m^2 c^4$$

$$\vec{p}^2 = (p^2)^2$$

$$(p^2)^2 = (p^\mu p_\mu)^2$$

$$(p^\mu p_\mu)^2 = p^\mu p^\nu \eta_{\mu\nu}$$

$$(8) \quad x^2, p^2 = (mc^2)^2, \partial_\mu \partial^\mu = \frac{\partial^2}{c^2 \partial t^2} - \nabla^2 := \square \quad \text{is invariant}$$

物理上，洛伦兹变换（Def. Constant-Type）是不变的

Lorentz transformation (Def. Constant-Type) rotation 3-space.

Infinite-dimensional LT (Def. Then  $\begin{cases} w_\mu = -w_\nu \\ w_\mu = w_\nu \end{cases}$ )  $\rightarrow$   $\begin{pmatrix} \gamma & -\beta \\ -\beta & \gamma \end{pmatrix} = \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix}$

$w_\mu = w_\nu \rightarrow \gamma = 1$  (which is 3D)

Def:  $\Lambda = (\lambda^{\mu}_{\nu})$  is defined as 4x4 matrix making  $x^2, p^2, \partial_\mu \partial^\mu$  invariant under  $\lambda^{\mu}_{\nu}$ .  
proper orthogonal  $\lambda^{\mu}_{\nu} \lambda^{\nu}_{\mu} = \lambda^{\mu}_{\nu} \lambda^{\mu}_{\nu}$   
details:  $\det(\lambda^{\mu}_{\nu}) = 1$

$x^2 = x^2 = \eta_{\mu\nu} x^\mu x^\nu = \eta'_{\mu\nu} x'^\mu x'^\nu = \eta_{\mu\nu} \gamma^\mu \gamma^\nu = \eta'_{\mu\nu} \gamma^\mu \gamma^\nu$  (①)

$\gamma^\mu = \gamma_\mu$  assumption

By comparing the ① and ②

•  $\gamma^2 = \gamma_0^2 + \gamma_1^2 + \gamma_2^2 + \gamma_3^2 = 1$  (which is 3D)

•  $\gamma^2 = \gamma_0^2 + \gamma_1^2 + \gamma_2^2 + \gamma_3^2 = 1$  (which is 3D)

•  $\gamma^2 = \gamma_0^2 + \gamma_1^2 + \gamma_2^2 + \gamma_3^2 = 1$  (which is 3D)

•  $\gamma^2 = \gamma_0^2 + \gamma_1^2 + \gamma_2^2 + \gamma_3^2 = 1$  (which is 3D)

Inverse of LT:  $(\lambda^{-1})^\mu_\nu = \frac{\partial x^\mu}{\partial x'^\nu}$

check:  $\eta^{\mu\nu} \eta_{\nu\lambda} \lambda^{\lambda\sigma} = \frac{\partial x^\mu}{\partial x'^\nu} \frac{\partial x'^\nu}{\partial x^\sigma} = \delta^{\mu\sigma}$

Indeed, we can define the Lorentz transformation

Expansion: Infinitesimal transform of  $\lambda^{\mu}_{\nu}$  and  $(\lambda^{-1})^{\mu}_{\nu}$

infinitesimal parameter  $\alpha$  along  $\alpha$

$\lambda^{\mu}_{\nu} = \delta^{\mu}_{\nu} + \omega^{\mu}_{\nu} + O(\omega^2)$

infinitesimal parameter  $\alpha$  along  $\alpha$

$(\lambda^{-1})^{\mu}_{\nu} = \delta^{\mu}_{\nu} - \omega^{\mu}_{\nu} + O(\omega^2)$

check:  $\eta_{\mu\nu} (\lambda^{-1})^{\mu}_{\nu} = \delta^{\mu}_{\nu} + O(\omega^2)$

check2:  $\eta_{\mu\nu} = \eta_{\nu\mu}$  Anti-symmetric tensor (very strong constraint on LT)

•  $\eta_{\mu\nu} (\delta^{\mu}_{\nu} + \omega^{\mu}_{\nu}) (\delta^{\nu}_{\lambda} + \omega^{\nu}_{\lambda}) = \eta_{\mu\lambda}$

"

$\eta_{\mu\nu} \delta^{\mu}_{\nu} \delta^{\nu}_{\lambda} + \eta_{\mu\nu} \omega^{\mu}_{\nu} \delta^{\nu}_{\lambda} + \eta_{\mu\nu} \delta^{\nu}_{\lambda} \omega^{\mu}_{\nu} = \eta_{\mu\lambda}$

"

$\eta_{\mu\lambda} + \eta_{\mu\nu} \omega^{\nu}_{\lambda} + \eta_{\mu\nu} \omega^{\nu}_{\lambda} \delta^{\nu}_{\lambda} = \eta_{\mu\lambda}$

"

$\omega^{\mu}_{\nu} = -\omega^{\nu}_{\mu}, \omega^{\mu}_{\nu} = \eta^{\mu\lambda} \omega^{\nu}_{\lambda} \Rightarrow \omega^{\mu}_{\nu} = \eta^{\mu\lambda} \omega^{\nu}_{\lambda} = -\omega^{\nu}_{\mu} \Rightarrow \omega^{\mu}_{\nu} = -\omega^{\nu}_{\mu}$

$\Rightarrow \omega^{\mu}_{\nu} + \omega^{\nu}_{\mu} = 0$

check3:  $\omega_{\mu\nu} = -\omega_{\nu\mu}$  rank-2 covariant tensor

$\omega^0_i = \omega^0_j, \omega^i_j = \omega^j_i \leftarrow \text{rank}(0,1)$  tensor

"

$\omega_{0j} = -\omega_{j0}, \omega^0_i = \eta^{00} \omega_{i0} \Rightarrow \omega^0_i = \eta^{ii} \omega_{i0} = -\omega_{i0} \Rightarrow \omega^0_i = -\omega_{i0}$

$= -\omega_{i0}$

$\omega^i_j = -\omega^j_i, \omega^i_j = \eta^{ii} \omega_{ij} = -\omega^j_i > \omega^i_j = -\omega^j_i$

$\omega^i_j = \eta^{ij} \omega_{ij} = -\omega^j_i$

e.g. Lorentz boost along  $x$ -direction

$\text{ct} \uparrow \quad \text{ct}' \quad \text{ct}'$

$$\text{Pf-Pf} \quad \text{Lorentz transformation (Def. Constant-Type)} \quad \text{rotation 3-space.}$$

$$\text{Infinite-dimensional LT (Def. Then } \begin{cases} w_\mu = -w_\nu \\ w_\mu = w_\nu \end{cases} \text{) } \rightarrow \begin{pmatrix} \gamma & -\beta \\ -\beta & \gamma \end{pmatrix} = \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix}$$

$$\text{Inverse: } \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & -\beta \\ -\beta & \gamma \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & 0 \\ 0 & \gamma \end{pmatrix}$$

$$\text{Invariance under } \lambda^{\mu}_{\nu} \text{ and } \lambda^{\mu}_{\nu} \text{ is invariant}$$

$$\text{proper-orthogonal } \lambda^{\mu}_{\nu} \lambda^{\nu}_{\mu} = \lambda^{\mu}_{\nu} \lambda^{\mu}_{\nu} \text{ and } \det(\lambda^{\mu}_{\nu}) = 1$$

$$\text{e.g. 2: Lorentz rotation along } z\text{-axis about } 0$$

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$$\text{IV \& VI. VII. Lorentz transformation (LT)}$$

$$\text{Def: } \Lambda = (\lambda^{\mu}_{\nu})$$

$$\text{is defined as 4x4 matrix making } x^2, p^2, \partial_\mu \partial^\mu$$

$$\text{proper-orthogonal } \lambda^{\mu}_{\nu} \lambda^{\nu}_{\mu} = \lambda^{\mu}_{\nu} \lambda^{\mu}_{\nu}$$

$$\text{det}(\lambda^{\mu}_{\nu}) = 1$$

$$\text{and } \det(\lambda^{\mu}_{\nu}) = 1$$

$$\text{and } \det(\lambda^{\mu$$



## 10) VII. Classical Field Theory

### 1. Definition of Classical Field (and example)

Classification of Classical Field {  
scalar  
vector  
Dive + bipolar  
Bipolar + longitudinal LG  
Anti-symmetric}

(Def 1) A classical field is a physical object described by a function of space-time.  
e.g.  $\vec{E}(t, \vec{x})$ ,  $B(t, \vec{x})$

Fields {  
Independent DOF  
L. S (World invariant)}

Def 2). A classical field describes a physical object with infinite number of DOF (degrees of freedom).

where space one just labels instead of observables in QM.

e.g.  $\vec{E}(x^0) = \vec{E}(t, \vec{x})$ ,  $\vec{B}(x^0)$ ,  $\vec{x} \in \mathbb{R}^3$ ; but in QM,  $\vec{x} \rightarrow \frac{\vec{p}}{m}$

\* right formulation  
here  $\vec{x}$  just labels, not observables  
\* QFT and LFT

e.g. 3.  $\vec{q}(t, \vec{x})$ ,  $q(t) = \phi(t)$  → particle mechanics (finite number of freedom)

(Def 3) A classical local field is a physical quantity whose dynamics at a space-time point is determined by field around such space-time point, namely, is described by PDE.

(Def 4) A classical local field represents a non-instantaneous interaction between particles at different space-time points.

e.g.  $\vec{E}(\vec{x}) = \frac{q}{4\pi\epsilon_0 r^2} \hat{r}$   
source charge test charge

Definition of Classical (local) Field theory ↑

### 7. Classification of classical fields

Classical fields	Representation of Lc	Transformation law	Example → Feynman (feynman)
real scalar fields $\phi(x)$	$(0,0)$ ↑ spin-less	$x \rightarrow x' = \gamma x$ $\phi(x) \rightarrow \phi(x') = \phi(x)$ $q(x) \rightarrow q'(x) = q(x)$	mesons pseudo-nucleons # Higgs
Complex scalar fields: $q(x)$	$\begin{cases} (0,0) & \text{if } q(x) \\ \text{spin } \frac{1}{2}, 0 & \text{if } \bar{q}(x) \end{cases}$	$\gamma \rightarrow \gamma' = \gamma \gamma'$ $q(x) \rightarrow q'(x) = m^{-1} \bar{q}(x)$ $q(x) \rightarrow q''(x) = g(\gamma - \gamma') \gamma$	photons electrons and nucleons
vector potentials	$A_\mu(x)$ (gauge potential)	$x \rightarrow x' = \gamma x$ $A_\mu(x) \rightarrow A'_\mu(x') = \gamma_{\mu\nu} A_\nu(x)$	Field theory

Dirac bi-spinor fields	$\psi(x) = \begin{pmatrix} \psi_1(x) \\ \psi_2(x) \\ \psi_3(x) \\ \psi_4(x) \end{pmatrix}$ (just column matrix)	$(\frac{1}{2}, 0) \oplus (0, \frac{1}{2})$ spin $\frac{1}{2}$ & 0 if $\psi$ .	$x \rightarrow x'$ $\psi_D(x) \rightarrow \psi_D(x') = \gamma_{\mu\nu} \psi_D(x)$ $S(x) \neq 0$	electrons and nucleons $S(x) \neq 0$
------------------------	--	--	---	--

2. Definitions	What about space-time translation group	$x \rightarrow x' = x + \epsilon$ infinite-dimensional translation
(Def 1) A classical field is a physical object described by a function of space-time.	$\phi(x) \rightarrow \phi(x') = \phi(x)$ $\psi(x) \rightarrow \psi(x') = \psi(x)$ $A_\mu(x) \rightarrow A'_\mu(x') = A_\mu(x)$ $\psi_D(x) \rightarrow \psi_D(x') = \gamma_{\mu\nu} \psi_D(x)$	• Hamiltonian unbroken. No gauge unbroken. No gauge unbroken. No gauge unbroken. No gauge
2. Lagrangian Formulation	Lagrangian	• How to derive FOF is not as easy as draw them.
3. Hamiltonian Formulation	Lagrange's eq.	• Space-point translation $\leftrightarrow$ Energy-momentum tensor. Lorentz boost $\leftrightarrow$ angular-momentum tensor
4. Canonical Form	Lagrange and Hamilton formulations	• $\vec{p}_i \rightarrow \vec{p}'_i$ : $\dot{x}_i \rightarrow \dot{x}'_i$ $\vec{p}_i \rightarrow \vec{p}'_i$ : $\dot{x}'_i \rightarrow \dot{x}_i$ $\leftrightarrow$ Hamiltonian and Hamiltonian $\vec{p}'_i$
5. Particle Brackets Form	* fields: (general case)	$\phi^\alpha(x, \vec{x}) = \phi_\alpha(x)$ : $a = 1, 2, \dots, n$ are components N-component field $\left\{ \begin{array}{l} \text{4-dim component} \\ \text{4-dim 1-component} \\ \text{4-dim 1-component} \\ \text{4-dim 1-component} \end{array} \right.$
6. Lorentz Group	* fields: (general case)	$\phi^\alpha(x, \vec{x}) = \phi_\alpha(x)$ : $a = 1, 2, \dots, n$ are components N-component field $\left\{ \begin{array}{l} \text{4-dim 4-component} \\ \text{4-dim 1-component} \\ \text{4-dim 1-component} \\ \text{4-dim 1-component} \end{array} \right.$
7. Fields	* fields: (general case)	$\phi^\alpha(x, \vec{x}) = \phi_\alpha(x)$ : $a = 1, 2, \dots, n$ are components N-component field $\left\{ \begin{array}{l} \text{4-dim 4-component} \\ \text{4-dim 1-component} \\ \text{4-dim 1-component} \\ \text{4-dim 1-component} \end{array} \right.$
8. Remarks	• Remark: particle mechanics = No spatial label $g_{\alpha\beta} = \delta_{\alpha\beta}$	• field $\rightarrow$ general object particle $\rightarrow$ special object
9. States	* states (independent Dof)	* states (independent Dof) $\rightarrow$ why not $\phi^\alpha(t, \vec{x})$
10. Action principle	* action principle (Lagrangian) $\rightarrow$ L = $\int d^3x \mathcal{L}(c^\alpha, \partial_\mu c^\alpha)$	$\phi(x, \vec{x}), \partial_\mu \phi(x, \vec{x})$ (like $g_{\alpha\beta}$ )
11. Lagrangian	* Lagrangian (Lagrangian): $\mathcal{L} = \mathcal{L}(a, \partial_\mu a)$	* Lagrangian (Lagrangian): $\mathcal{L} = \mathcal{L}(a, \partial_\mu a)$
12. Action principle	* Action principle: $S = \int d^4x \mathcal{L}(c^\alpha, \partial_\mu c^\alpha)$	* Action principle: $S = \int d^4x \mathcal{L}(c^\alpha, \partial_\mu c^\alpha) \rightarrow$ non-local theories
13. Field theory	Note: $L = \int d^4x \delta(c^\alpha, \partial_\mu c^\alpha)$	• $\mathcal{QFT} \rightarrow \text{QM+SR}$ Action is LT invariant. because $d^4x' = d^4x$ , $\mathcal{L}(c^\alpha, \partial_\mu c^\alpha) = \mathcal{L}(c^\alpha, \partial_\mu c^\alpha)$ Action is translation-invariant. $d^4x' = d^4x$
14. Action principle:	* Action principle: particle mechanics $\rightarrow$ field theory	Action is everything.
15. Time boundary is fixed	Time boundary is fixed	• $\phi_\alpha^{(0)} \rightarrow \phi_\alpha^{(0)} + \delta\phi_\alpha^{(0)}$ Arbitrary infinitesimal transform. Time boundary is fixed
16. Stationary Action $S_{\text{S}}$	Stationary Action $S_{\text{S}} = 0$	• $\phi_\alpha^{(0)} \rightarrow \phi_\alpha^{(0)} + \delta\phi_\alpha^{(0)}$ Arbitrary infinitesimal transform. Time boundary is fixed



(2)

$$\phi(x, \vec{r}) \in \mathbb{R} \quad (\text{real}), \quad \phi(x) = \phi(x_1, \dots, x_n) \quad (\text{scalar})$$

$$\phi'(x+\xi) = \phi(x) \quad , \quad \vec{x}' = \vec{x}(x) \quad (\text{scalar})$$

$$x = x_0 + x_{\text{int}} \quad , \quad x_0 \xrightarrow{\text{mass of meson, real physics}}$$

$$x_0 = \frac{1}{2} \left( \partial_\mu \phi \partial^\mu \phi - m^2 \phi^2 \right) \quad (\text{intrinsic transformation, and Action Principle})$$

$$V(\phi) \geq 0 \quad (\text{potential energy})$$

$$\text{e.g. } V(\phi) = \frac{1}{4!} \phi^4 \quad (\lambda : \text{coupling constant, } \phi \text{ to } \phi \text{ model})$$

$$S = \int d^4x \mathcal{L} \quad , \quad \phi \rightarrow \phi + \delta \phi \quad (\text{intrinsic transformation, and Action Principle})$$

$$\frac{\partial \mathcal{L}}{\partial \phi} = \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)}$$

$$\frac{\partial \mathcal{L}}{\partial \phi} = -m^2 \phi - \frac{\partial V(\phi)}{\partial \phi} = -m^2 \phi - V'(\phi) \quad (\cancel{\text{is 2nd term, not 1st term}})$$

$$\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = \partial_\mu \phi \quad , \quad \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = \square \phi = \partial_\mu \partial^\mu \phi$$

$$\text{Eq: } (\square + m^2) \phi = -V'(\phi)$$

$$\text{where } V(\phi) = 0 \Rightarrow (\square + m^2) \phi = 0$$

real Klein-Gordon equation  
(with no interaction)  
Exact solution: known

$$V(\phi) = \frac{1}{4!} \phi^4$$

$$(\square + m^2) \phi = -\frac{1}{3!} \phi^3 \quad \text{Nonlinear equation}$$

Eq:

Exact Solution: unknown up to now. (That's why we use perturbative QFT, Feynman diagram)

we are weak in solving the EOM.

Feynman diagram / we know the exact solution of Real Klein-Gordon

we don't know... at 4D interaction

Pseudo-nucleon "n" "

Note: Quantum Complex Scalar Field is used for Higgs particle

and spin is nature even in classical field

•  $\psi(x) \in \mathbb{C}$  (complex field)

$\psi(x) = \psi(x), \quad \psi = \psi(x) \quad (\text{scalar field})$

$\psi(x+\xi) = \psi(x)$

$\psi(x) \in \mathbb{C}$  (complex field)

$\psi(x) = \psi(x), \quad \psi = \psi(x) \quad (\text{scalar field})$

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$\psi(x) \in \mathbb{C}$  (complex field)

$\psi(x) = \psi(x), \quad \psi = \psi(x) \quad (\text{scalar field})$

$\psi(x+\xi) = \psi(x)$

\* Eq 1. should derive from  $\delta S = 0$

Action:  $S = \int d^4x (\mathcal{L}_0 + \mathcal{L}_{\text{int}}) \quad , \quad u \text{ and } v \text{ are independent DOF}$

(o whq? )

$$\phi'(x+\xi) = \phi(x) \quad \frac{\partial \mathcal{L}}{\partial \phi} = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \quad \frac{\partial \mathcal{L}}{\partial \phi} = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \psi)}$$

$$\frac{\partial \mathcal{L}}{\partial \phi} = -m^2 \psi^\dagger - \frac{\partial V(\phi)}{\partial \phi} \quad \left. \begin{array}{l} \text{use Hamilton Formulation to derive} \\ \text{the same Eqs as Lagrange Formalism} \end{array} \right\} \quad \psi^\dagger = \partial_\mu (\partial_\mu \psi) = \square \psi^\dagger \quad \downarrow$$

$$\frac{\partial \mathcal{L}}{\partial \psi} = -m^2 \psi + \frac{\partial V(\phi)}{\partial \psi} \quad \left. \begin{array}{l} \text{Hamiltonian and Canonical conjugate} \\ \text{momentum} \end{array} \right\} \quad \left. \begin{array}{l} (1+m^2) \psi = -\frac{\partial V(\phi)}{\partial \psi} \\ (1+m^2) \psi = \pi_\psi \end{array} \right.$$

\* Hamiltonian and Canonical momentum

$$\Pi_\psi = \frac{\partial \mathcal{L}}{\partial \dot{\psi}} \quad , \quad \Pi_\psi^\dagger = \frac{\partial \mathcal{L}}{\partial \dot{\psi}^\dagger} \quad (\text{two conjugate momentum})$$

=  $\dot{\psi}^\dagger$

=  $\dot{\psi}$

=  $\psi^\dagger$

=  $\psi$

=  $\psi^\dagger$

### • Gamma Matrices

Gamma matrix  $\gamma^\mu$ ,  $\mu = 0, 1, 2, 3$ ,  $4 \times 4$  matrices, satisfying

$$\textcircled{1} \quad \{\gamma^\mu, \gamma^\nu\} = \gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2\eta^{\mu\nu}$$

$$\textcircled{2} \quad (\gamma^\mu)^\dagger = \gamma^\mu, \quad (\gamma^i)^\dagger = -\gamma^i, \quad i = 1, 2, 3$$

( $i\gamma_1 \gamma_2 \gamma_3 \gamma_4$ )

Dirac Conjugation  $\bar{\psi}_D(x) \equiv \psi_D^\dagger(x) \gamma^0$

Hermitian Conjugation

$$\bar{\psi}_D \equiv \gamma^0(\gamma^0)^\dagger \bar{\psi}_D = \gamma^\mu$$

(The Dirac Conjugation  $\bar{\psi}_D$ ,  $\bar{\psi}_D \neq \bar{\psi}_D$  (Because  $\bar{\psi}_D$  is Dirac conjugate))

$$\bar{\psi}_D = \bar{\psi}_{D,\alpha} \pi_\alpha + \bar{\psi}_{D,\bar{\alpha}} \bar{\pi}_{\bar{\alpha}}$$

summation with a

Lagrangian density for Dirac field

$$\mathcal{L}(\psi_D) = \bar{\psi}_D(x) i^\mu \partial_\mu \psi_D(x) - m_e \bar{\psi}_D(x) \psi_D(x)$$

$\hookrightarrow$  electron mass

$$= \bar{\psi}_D(x) (i^\mu \partial_\mu - m_e) \psi_D(x)$$

$\textcircled{1}$

$$= \bar{\psi}_D(x) (i^\mu \partial_\mu - m_e \delta_{ab}) \psi_b^D(x)$$

$$= \psi_{D,a}^\dagger i^\mu \partial^\mu + i^\mu \bar{\psi}_D^\dagger (\gamma^0 \delta_{ab}) \partial_b \psi_b^D(x)$$

$\textcircled{2}$

$$= \psi_{D,a}^\dagger i^\mu \partial^\mu + i^\mu \bar{\psi}_D^\dagger (\gamma^0 \delta_{ab}) \partial_b \psi_b^D(x)$$

$\textcircled{3}$

From:  $\psi_D^\dagger$  and  $\bar{\psi}_D$  are independent D.o.F

(matrix  $\psi_D$ ,  $\bar{\psi}_D$ , we usually write down component  $\psi_D^\alpha(x)$ ,  $\bar{\psi}_D^\alpha(x)$ )

FORM:

$$\frac{\partial \mathcal{L}}{\partial \psi_D^\mu} = 0 \quad (\Rightarrow) \quad \frac{\partial \mathcal{L}}{\partial \bar{\psi}_D^\mu} = 0$$

$$\begin{array}{l} \text{Left side: } \\ \frac{\partial \mathcal{L}}{\partial \bar{\psi}_D^\mu} = 0 \quad (\Rightarrow) \quad \boxed{\frac{\partial \mathcal{L}}{\partial \psi_D^\mu} = 0} \\ \text{Right side: } \\ \frac{\partial \mathcal{L}}{\partial \bar{\psi}_D^\mu} = 0 \end{array}$$

Eqs should be (cancel Dirac equation)

$$\frac{\partial \mathcal{L}}{\partial \psi_D^\mu} = (i^\mu \partial_\mu - m_e) \psi_D = 0 \quad (\text{matrix})$$

$$(i^\mu \partial_\mu - m_e) \psi_D = 0 \quad \rightsquigarrow \text{Right Dirac Conjugate, } \bar{\psi}_D^\mu - i^\mu \bar{\psi}_D^\mu = 0$$

Trivial Hermitian Hamiltonian Density

Two diff conventional

$$\mathcal{T}_{D,\mu} = \frac{\partial \mathcal{L}}{\partial \dot{\psi}_D^\mu(x)}, \quad \overline{\mathcal{T}}_{D,\mu} = \frac{\partial \mathcal{L}}{\partial \bar{\psi}_D^\mu(x)} = 0$$

$$\textcircled{2} \quad i \bar{\psi}_D^\mu(x)$$

In component formulation  $\bar{\psi}_D^\mu <$  complex field.

$$H = \int d^3x = \int d^3x \cdot i \bar{\psi}_D^\mu \frac{\partial}{\partial t} \psi_D^\mu = i \bar{\psi}_D^\mu \frac{\partial}{\partial t} \psi_D^\mu \quad (\text{classical Hamiltonian})$$

$$H = \int d^3x = \int d^3x \cdot i \bar{\psi}_D^\mu \frac{\partial}{\partial t} \psi_D^\mu = i \bar{\psi}_D^\mu \frac{\partial}{\partial t} \psi_D^\mu \quad (\text{classical Hamiltonian})$$

$$\text{Remark: In Quantum Mechanics, } i \frac{\partial}{\partial t} \psi_D^\mu = H \text{ (} \hbar = c = 1 \text{)}$$

$$\text{In Field (classical) theory, } i \bar{\psi}_D^\mu \frac{\partial}{\partial t} \psi_D^\mu = m_e \bar{\psi}_D^\mu$$

$$\Rightarrow i \frac{\partial}{\partial t} \psi_D^\mu + i \bar{\psi}_D^\mu \frac{\partial}{\partial t} \psi_D^\mu = m_e \bar{\psi}_D^\mu \quad (\text{no } \beta)$$

$$\Rightarrow i \frac{\partial}{\partial t} \psi_D^\mu = (-i \vec{\alpha} \cdot \vec{p} + m_e \beta) \psi_D^\mu$$

$$= (\vec{\alpha} \cdot \vec{p} + m_e \beta) \psi_D^\mu \quad \left\{ \begin{array}{l} \beta = \gamma^0 \\ \alpha^i = \gamma^i \gamma^0 \end{array} \right.$$

$$H = -i \vec{\alpha} \cdot \vec{p} + m_e \beta$$

But use  $\gamma^\mu$ , Lorentz transform is more simple

$$= \vec{\alpha} \cdot \vec{p} + m_e \beta$$

• Detail of Electro-Dynamics. See the solution of the first homework

### ▽ III. 4 Maxwell Vector Field Theory

Note: Quantum version of ... is for photon

$\vec{E}$  in  $\vec{B}$

- We gauge potential field  $A_\mu(x)$ .  $A_\mu(x) = A_{\mu(\nu)} = \begin{pmatrix} A_{0,00} \\ A_{0,01} \\ A_{0,02} \\ A_{0,03} \end{pmatrix}$

$$\text{Remark 3. } \vec{E} = -\nabla \phi - \frac{\partial \vec{A}}{\partial t}, \quad \vec{B} = \nabla \times \vec{A} \quad E, B \text{ strength}$$

$$A_{\mu(\nu)} = (A_{\mu(00)}, \vec{A}_{\mu(01)})$$

Remark 2: two maxwell equations  $\nabla \cdot \vec{E} = -\frac{\partial \vec{B}}{\partial t}$  can be derived use Remark 1's def of  $E, B$

$$\nabla \cdot \vec{B} = 0$$

Remark 3: Use  $A_{\mu(00)}$ , more convenient for Lorentz transform.

Field strength (tensor)  $F_{\mu\nu} = \partial_\mu A_{\nu(00)} - \partial_\nu A_{\mu(00)}$ ,  $E^i = F^{0i}$ ,  $B_i = \frac{1}{2} \epsilon_{ijk} F^{jk}$

Change Transformation:  $A_{\mu(00)} \rightarrow A_{\mu(00)} + \partial_\mu \tilde{A}_{\mu(00)}$ ,  $F_{\mu\nu}$  is unchanged! Totally anti-symmetric tensor

Arbitrary smooth function







$$g = \exp \left( -i (\vec{\beta}^{(\alpha)} + \vec{\gamma}^{(\alpha)}) \cdot \vec{\theta} \right) \exp \left( -(\vec{\beta}^{(\alpha)} - \vec{\gamma}^{(\alpha)}) \cdot \vec{\theta} \right)$$

$$\begin{pmatrix} \psi_+ \\ \psi_- \end{pmatrix}_{(1,1)} \approx \begin{pmatrix} \psi_+ \\ \psi_- \end{pmatrix}_{(1,2)} \otimes \begin{pmatrix} \psi_+ \\ \psi_- \end{pmatrix}_{(2,2)}$$

$$\begin{array}{c} \downarrow \\ \{\psi_{\mu}\} \\ \downarrow \\ |\bar{\psi}_m\rangle \otimes |\bar{\psi}_{2m}\rangle \\ (\bar{\psi}, \bar{\psi}) \end{array}$$

- spinless particle (scalar)  $(0,0) \approx |00\rangle \otimes |00\rangle$

$$\begin{array}{c} \downarrow \\ \{\psi_{\mu}\} \\ \downarrow \\ (\psi_{\mu}, \psi_{\mu}) \end{array}$$

$$(0, \frac{1}{2}) \rightarrow \text{Weyl / Majorana Spinor (Neutrino)} \quad |00\rangle \otimes |1\frac{1}{2}, \pm\frac{1}{2}\rangle$$

$$(\frac{1}{2}, 0) \oplus (0, \frac{1}{2}) \rightarrow \text{Dirac bispinor (electron)} \quad |\frac{1}{2}, \pm\frac{1}{2}\rangle \otimes |00\rangle \oplus |00\rangle \otimes |\frac{1}{2}, \pm\frac{1}{2}\rangle$$

↓ perhaps  $\varrho$  is not fundamental particle.

spin 1

$$\begin{array}{c} \downarrow \\ \{\psi_{\mu}\} \\ \downarrow \\ (\psi_{\mu}, \psi_{\mu}) \end{array}$$

$$\text{massive spin-1} \quad \text{Three polarization: } \pm, 0$$

$$(\frac{1}{2}, \frac{1}{2}) \text{ massless spin-1 (photon) Two polarization: } \pm$$

## Symmetries & Conservation Laws

### IX.1. Transformation of fields

#### Ix.1.a Space-time translation

- Translation group  $\mathcal{G} \subset G$ ,  $\varphi = \exp(-i p_{\mu} x^{\mu})$  similar operator?

- Lie Algebra  $[P_{\mu}, P_{\nu}] = 0$

- Space-time translation  $x^{\mu} \mapsto x'^{\mu} = x^{\mu} + \xi^{\mu}$

$$\delta x^{\mu} = \xi^{\mu} = -i \epsilon^{\mu}_{\rho} p^{\rho}$$

Coordinate change  
Notation different  
Just notation

- $\psi(x) \rightarrow \psi(x') = \varphi^{-1} \psi(x) \varphi$

$\varphi(x) \rightarrow \varphi(x') = \varphi(x) \varphi^{-1}(x')$

$$Q \rightarrow \Phi(x) = \varphi(\tau^{\mu} x) = \varphi(x - \xi)$$

$\varphi(x) = \varphi(x' - \xi)$

### IX.2 Lorentz transformation of fields

• (Real or Complex) Scalar fields

$$\pi \xrightarrow{\Delta} \pi' = \pi \pi$$

$$\phi(x) \xrightarrow{\square} \phi'(x) = \phi(x) - \phi(x)$$

$$\partial_{\mu} \phi = \partial'_{\mu} \phi, -\partial_{\mu} \phi = \phi_{,\mu} - \phi_{,\mu}$$

$$\pi^{\mu} = \pi^{\mu} \pi_{\nu} = (\delta^{\mu}_{\nu} + w^{\mu}_{\nu}) \pi^{\nu} = \pi^{\mu} + w^{\mu}_{\nu} \pi^{\nu}$$

$$\delta \pi^{\mu} = w^{\mu}_{\nu} \pi_{\nu} = \frac{1}{2} w^{\mu}_{\rho} w^{\rho}_{\sigma} \pi^{\sigma} \quad M_{\rho\sigma} = i(\pi_{\rho} \partial_{\sigma} - \pi_{\sigma} \partial_{\rho})$$

$$\text{use group language, generator of LC} \quad \frac{\delta \phi(x)}{\delta w^{\mu\nu}} = 0 \Leftrightarrow \delta \phi(x) = 0$$

$$\frac{\delta \phi(x)}{\delta w^{\mu\nu}} = \frac{i}{2} M_{\rho\sigma} \pi^{\mu} = \frac{1}{2} (\eta^{\mu}_{\rho} \pi_{\sigma} - \eta^{\mu}_{\sigma} \pi_{\rho}) \quad \text{Rep of scalar field}$$

$$\frac{\delta \phi(x)}{\delta w^{\mu\nu}} = 0 \Leftrightarrow \delta \phi(x) = 0$$

$$\delta \phi(x) = \phi(x - \pi) - \phi(x) = \phi(x - w^{\mu} \pi_{\mu}) - \phi(x) = -w^{\mu} \pi_{\mu} \delta \phi(x)$$

• (Dirac) bispinor field

$$\pi \xrightarrow{\Delta} \pi' = \pi \pi \quad \psi_{\rho}(x) \xrightarrow{\square} \psi'_{\rho}(x') = D(x) \gamma_{\rho} \psi_{\rho}(x)$$

$$D(x) = \exp \left( -\frac{i}{2} w^{\mu\nu} \gamma_{\mu\nu} \right) \quad \text{generator Spin-}\frac{1}{2}$$

group parameter

$$\delta_{\mu\nu} = \frac{1}{2} \delta_{\mu\nu} = \frac{i}{4} [\gamma_{\mu}, \gamma_{\nu}]$$

$$\delta \psi_{\rho}(x) = \psi'_{\rho}(x') - \psi_{\rho}(x) = -\frac{i}{2} w^{\mu\nu} \sigma_{\mu\nu} \gamma_{\rho} \psi_{\rho}(x)$$

$$\frac{\delta \psi_{\rho}(x)}{\delta w^{\mu\nu}} = -\frac{i}{2} \sigma_{\mu\nu} \psi_{\rho}(x) \quad \text{if equivalent}$$

→ transformation

$$\delta \psi_{\rho}(x) = \psi'_{\rho}(x) - \psi_{\rho}(x) = D(x) \gamma_{\rho} (\pi - \pi') - \pi'_{\rho}(x) = -\frac{i}{2} w^{\mu\nu} \gamma_{\mu\nu} \psi_{\rho}(x)$$

$$M_{\mu\nu} = \gamma_{\mu\nu} + \gamma_{\nu\mu}, \quad L_{\mu\nu} = i(\gamma_{\mu} \partial_{\nu} - \gamma_{\nu} \partial_{\mu})$$

• Maxwell Vector field

$$\pi \xrightarrow{\Delta} \pi' = \pi \pi \quad A^{\mu}(x) \xrightarrow{\square} A'^{\mu}(x') = \gamma^{\mu}_{\nu} A_{\nu}(x)$$

$$\Lambda = \exp \left( -\frac{i}{2} w^{\mu\nu} \gamma_{\mu\nu} \right) \quad \text{spin-1}$$

$$(\delta_{\mu\nu})^{\nu}_{\mu} = i(\gamma_{\mu} \gamma^{\nu} - \gamma_{\nu} \gamma^{\mu}) \quad (\text{check for consistency})$$

$$\delta A_{\mu}^{\nu}(x) = A'_{\mu}^{\nu}(x) - A_{\mu}^{\nu}(x) = -\frac{i}{2} w^{\rho\sigma} (\delta_{\mu}^{\nu} \gamma_{\rho} - \gamma_{\mu} \delta^{\nu}_{\rho}) A_{\nu}(x)$$

$$\frac{\delta A_{\mu}^{\nu}(x)}{\delta w^{\rho\sigma}} = -\frac{i}{2} (\delta_{\mu}^{\nu} \gamma_{\rho} - \gamma_{\mu} \delta^{\nu}_{\rho}) A_{\nu}(x)$$

by hand

You can derive the solutions of 1st home work in this way.

$$\delta A_{\mu}^{\nu}(x) = A'_{\mu}^{\nu}(x) - A_{\mu}^{\nu}(x) = \gamma_{\mu}^{\nu} A_{\nu}(x - \pi) - A_{\mu}^{\nu}(x) = -\frac{i}{2} w^{\rho\sigma} (\delta_{\mu}^{\nu} \gamma_{\rho} - \gamma_{\mu} \delta^{\nu}_{\rho}) A_{\nu}(x)$$

$$(\delta_{\mu\nu})^{\nu}_{\mu} = i(\gamma_{\mu} \gamma^{\nu} - \gamma_{\nu} \gamma^{\mu})$$



\* space-time translation group

$$\pi \xrightarrow{\tau} \pi' = \pi + \varepsilon = \tau \pi$$

the coordinate transformation

$$\Phi_a(x) \xrightarrow{U} \Phi'_a(x') = U \circ \Phi_a \circ U^{-1}(x') = \Phi_a(\tau^{-1}x')$$

infinitesimal transformation

$$U_t = \exp(-i\beta_\mu \varepsilon_\mu)$$

$\xrightarrow{b}$   
uniform transformation

$$\pi \xrightarrow{b} \pi' = \pi \varepsilon$$

\* Lorentz transformation  
+ the coordinates

$$\pi \xrightarrow{L} \pi' = \pi \varepsilon$$

momentum, can be expressed as

$$p_\mu = i\gamma_\mu \text{ for coordinate transform}$$

$$p_\mu = -i\gamma_\mu \text{ for field transform}$$

$$\left. \begin{array}{l} \text{spin} \\ \text{representation} \\ \text{including gravity} \\ \text{electron} \rightarrow \frac{1}{2} \text{ representation} \\ \text{graviton} \rightarrow 2 \text{ representation} \\ \dots \end{array} \right\}$$

infinitesimal transformation

$$D(\lambda) = \exp\left(-\frac{i}{2}\omega_{\mu\nu} S^{\mu\nu}\right)$$

$\xrightarrow{c}$   
spin momentum  $\left\{ \begin{array}{l} \gamma_2 \\ \vdots \\ 1 \end{array} \right.$

$$\Phi_a(x') = D(\lambda) \circ \Phi_a(x)$$

$$U(\lambda) = \exp\left(-i\frac{1}{2}w_{\mu\nu} N^{\mu\nu}\right)$$

$$\left. \begin{array}{l} \text{general field} \\ \text{Scalar} \\ \text{vector} \\ \dots \end{array} \right\}$$

$$\text{of Lorentz group}$$

The space-time Poincaré's group algebra is generated by  $\{P_\mu\}$  and  $\{M_{\mu\nu}\}$

$$[P_\mu, P_\nu] = 0$$

$$[P_\mu, M_{\rho\sigma}] = \eta_{\mu\rho} P_\sigma - \eta_{\mu\sigma} P_\rho$$

$$[M_{\mu\nu}, M_{\rho\sigma}] = \eta_{\mu\rho} M_{\nu\sigma} - \eta_{\mu\sigma} M_{\nu\rho} + \eta_{\nu\rho} M_{\mu\sigma} - \eta_{\nu\sigma} M_{\mu\rho}$$

$$M_{\mu\nu} = i(\gamma_1 \gamma_2 - \gamma_2 \gamma_1) + S_{\mu\nu}$$

$$\text{of Lorentz group}$$

$$\text{Two Casimir Operator}$$

$$[C, P_\mu] = [C, M_{\mu\nu}] = 0$$

$$(Conserved)$$

$$P^2 = P_\mu P^\mu = m_0^2$$

$$\uparrow \text{just number.}$$

$$W^2 = W^\mu W_\mu$$

$$W_\mu = -\frac{1}{2} \sum \epsilon^{\mu\nu\rho\sigma} P_\nu M_{\rho\sigma}$$

$$W^2 = -\frac{1}{2} S^{\mu\nu} P^2 + P_\mu S^{\mu\nu} S_\nu P^\mu$$

\* very good symmetry. has representation of mass, spin angular momentum  
(intrinsic quantity, not include  $C$ )

if,  $m_0 \neq 0$  choose rest frame  
(it's possible but not available for photon, neutrino)

$$P^\mu = (m, 0, 0, 0)$$

$$W^2 = -m^2 g^{ab} S^a S^b, \quad k=1,2,3$$

$$S^K = \frac{1}{2} \epsilon_{ijk} S^i j^K$$

only spin-angular momentum survived

$$\left. \begin{array}{l} \xrightarrow{d} \\ \text{particle in rest, the representation is simpler} \end{array} \right\} S^2 |S_{\pm 2}\rangle = S(S_{\pm 1}) |S_{\pm 2}\rangle$$

$|m, s, \pm_2\rangle$

$$z) \quad \gamma_5 m_0 = 0, \quad P^2 = 0, \quad W_\mu P^\mu = 0$$

Casimir operators is zero

Something must be changed

$$\text{define } \lambda = \frac{w^0}{p^0} = \frac{\tilde{S} \cdot \vec{p}}{|\vec{p}|}, \quad \text{it is Helicity operator, its representation is}$$

$$\frac{\tilde{S} \cdot \vec{p}}{|\vec{p}|} |\pm\rangle = \lambda |\pm\rangle, \quad \lambda = \pm \frac{1}{2}, \pm 1, \pm 2, 0$$

$$\left. \begin{array}{l} \text{massless} \\ \left\{ \begin{array}{l} |0, \pm \frac{1}{2}\rangle \\ |0, \pm 1\rangle \\ |0, \pm 2\rangle \end{array} \right. \end{array} \right\} \text{neutrino}$$

$$\left. \begin{array}{l} \text{massless} \\ \left\{ \begin{array}{l} |0, \pm \frac{1}{2}\rangle \\ |0, \pm 1\rangle \end{array} \right. \end{array} \right\} \text{photon}$$

$$\left. \begin{array}{l} \text{massless} \\ \left\{ \begin{array}{l} |0, \pm \frac{1}{2}\rangle \\ |0, \pm 1\rangle \end{array} \right. \end{array} \right\} \text{graviton}$$

$$\left. \begin{array}{l} \text{massless} \\ \left\{ \begin{array}{l} |0, \pm \frac{1}{2}\rangle \\ |0, \pm 1\rangle \end{array} \right. \end{array} \right\} \text{gluon}$$

$$\left. \begin{array}{l} \text{massless} \\ \left\{ \begin{array}{l} |0, \pm \frac{1}{2}\rangle \\ |0, \pm 1\rangle \end{array} \right. \end{array} \right\} \text{gravitino}$$

$$\left. \begin{array}{l} \text{massless} \\ \left\{ \begin{array}{l} |0, \pm \frac{1}{2}\rangle \\ |0, \pm 1\rangle \end{array} \right. \end{array} \right\} \text{Higgs}$$

## X. Symmetries & Conservation Laws

### X.1 Transformation of Fields

Space-time transf.	space-time coordinate	dynamic fields	examples
changed	changed	changed	translation, Lorentz transf.
fixed	fixed	changed	internal transf., Lorentz transf.

$\xrightarrow{e}$   
Poincaré group's representation  
of transl. 1  
(Very Very Important)

$\xrightarrow{f}$   
is the entire particle physics

Classification

Continuous  
transf.

Continuous  
transf.

No parameter

Parity, time-reversal  
Lorentz transf.

Transformation parameter  
transf.

continuous  
transf.

discrete  
transf.

charge conjugation  
(particle  $\leftrightarrow$  anti-particle)

big achievement of QFT  
it can predict anti-particle



### X.3.3 Internal Symmetries and Conserved Current.

Local fields  $\rightarrow$  unchanged

Coordinate of spacetime  $\rightarrow$  unchanged

$$\delta \varphi^\dagger = 0$$

$$\frac{\partial \varphi^\dagger}{\partial \omega^\alpha} = 0$$

$$\delta \varphi^\dagger = 0$$

So the conserved current

$$j^\mu_\alpha = \sum_a \frac{\partial \mathcal{L}}{\partial \dot{\varphi}_a^\dagger} \frac{\delta \dot{\varphi}_a}{\delta \omega^\alpha}$$

$$\text{just phase factor} \quad \tau_\theta(\lambda) = \left\{ e^{i\lambda} \mid \lambda \in \mathbb{R} \right\}$$

$$\text{parameter } e^{i\lambda} \rightarrow \lambda$$

$$\text{it is a Lie group}$$

The Lagrangian of Complex Scalar field

$$\mathcal{L} = \frac{1}{2} \partial_\mu \vec{\phi}^\dagger \partial^\mu \vec{\phi} - \frac{1}{2} m^2 \vec{\phi}^\dagger \cdot \vec{\phi} - V(\vec{\phi}^\dagger \vec{\phi})$$

The  $\text{U}(1)$  transformation

$$\psi(x) \xrightarrow{\text{U}(1)} \psi'(x) = e^{-i\lambda} \psi(x)$$

$$\psi(x) \xrightarrow{\text{U}(1)} \psi'(x) = e^{i\lambda} \psi(x)$$

$$\rightarrow \text{we can check } \psi + \psi \rightarrow \psi + \psi \text{ unchanged}$$

$$\partial_\mu \psi + \partial^\mu \psi \rightarrow \partial_\mu \psi + \partial^\mu \psi$$

$\mathcal{L} \rightarrow \mathcal{L}$  unchanged (4-d divergence)

The conserved current given by Noether's theorem

$$\frac{\delta \mathcal{L}(\psi)}{\delta \omega^\alpha} = \frac{\delta \mathcal{L}}{\delta \dot{\psi}} = -i \dot{\psi}^\dagger \psi - i \dot{\psi} \psi^\dagger$$

$$\frac{\delta \mathcal{L}}{\delta \omega^\alpha} = \frac{\delta \mathcal{L}}{\delta \dot{\psi}} = -i \dot{\psi}^\dagger \psi - i \dot{\psi} \psi^\dagger$$

$$\text{variation equations} \\ \delta \psi^\dagger = i \lambda \psi^\dagger \quad (\text{Lie group def.})$$

$$\begin{aligned} j^\mu &= \frac{\partial \mathcal{L}}{\partial \dot{\psi}} + \frac{\partial \mathcal{L}}{\partial \omega^\mu}, \frac{\delta \psi^\dagger}{\delta \lambda} = \partial_\mu \psi^\dagger (-i \dot{\psi}^\dagger) + \partial_\mu \psi (i \dot{\psi}) \\ &= i (\partial_\mu \psi^\dagger) \psi + (\partial_\mu \psi) \psi^\dagger \end{aligned}$$

$$= (i \phi_1^\dagger \phi_2 - \phi_1^\dagger \phi_2)$$

expressed by

corresponding

real scalar field

### Example 2. $\text{SO}(2)$ invariant Real Scalar Field $\rightarrow$ $\text{U}(1)$ invariant Complex Scalar field

$\text{SO}(2) = \{ R_{2 \times 2} \mid R^T R = \text{Id}_2, \det R = 1 \}$

$$R = \exp(i\sigma_y \cdot \lambda), \quad i\sigma_y = (-1, 0)$$

$$= \begin{pmatrix} \cos \lambda & \sin \lambda \\ \sin \lambda & -\cos \lambda \end{pmatrix}$$

$$\text{Define the field vector - spinor} \\ \vec{\phi} = (\phi_1, \phi_2) \quad \vec{\phi}^T = (\phi_1, \phi_2)$$

+1st term is cancelled  
it is simplified

But the transform of  $\delta \phi^\alpha$  is complicated

The corresponding Lagrangian Density

$$\mathcal{L} = \frac{1}{2} \partial_\mu \vec{\phi}^\dagger \partial^\mu \vec{\phi} - \frac{1}{2} m^2 \vec{\phi}^\dagger \cdot \vec{\phi} - V(\vec{\phi}^\dagger \vec{\phi})$$

check the  $\text{SO}(2)$  invariant

$$\vec{\phi} \xrightarrow{R} R \vec{\phi}$$

$$\vec{\phi}^\dagger \xrightarrow{R} R^\dagger \vec{\phi}^\dagger$$

$$\vec{\phi}^\dagger \vec{\phi} \xrightarrow{R} R^\dagger \vec{\phi}^\dagger \vec{\phi}$$

$$\partial_\mu \vec{\phi}^\dagger \partial^\mu \vec{\phi} \xrightarrow{R} \partial_\mu R^\dagger \partial^\mu \vec{\phi}$$

$$\partial_\mu \vec{\phi}^\dagger \partial^\mu \vec{\phi} \xrightarrow{R} \partial_\mu \vec{\phi}^\dagger \partial^\mu \vec{\phi}$$

use the  $\vec{\phi}' = R \vec{\phi} = \exp(i\sigma_y \lambda) \vec{\phi}$  to give the variation equations

$$\delta \vec{\phi} = \vec{\phi}' - \vec{\phi} = \lambda (i\sigma_y) \vec{\phi} \quad \delta \vec{\phi} = (i\sigma_y) \vec{\phi}$$

We take the Noether's current

$$j^\mu = \frac{\partial \mathcal{L}}{\partial \dot{\phi}^\alpha} = \partial_\mu \vec{\phi}^\dagger \cdot (i\sigma_y) \vec{\phi} \quad \leftarrow \text{do everything by component}$$

$$= \partial^\mu \phi_1 \cdot \phi_2 - \partial^\mu \phi_2 \cdot \phi_1$$

Note:  $j_\mu = i((\partial_\mu \psi) \psi^\dagger - (\partial_\mu \psi^\dagger) \psi)$  for complex scalar field,

we make the relationship  $\psi = \frac{1}{2} (\phi_1 + i\phi_2)$

$$\text{we can prove } j_\mu = (\partial^\mu \phi_1) \phi_2 - (\partial^\mu \phi_2) \cdot \phi_1$$

now, we have proved it since  
the same Lagrangian density

X.3.4 The Discrete Symmetries and CPT Theorem

1. Discrete Symmetries and CPT Theorem

No Noether's theorem

No conserved current

↳ quantum field theory

Def (of Discrete Symmetries) Position, time reversal, Charge Conjugation

CPT

CC

CPT theorem: Any Local Quantum Field are invariant under the Joint C.P.T  
↑ very very strong constraint on QFT



\* Homework 2nd. Due Nov, 24, 2016

PDF file for P & S's book, has the name of ... QFT - Non-Commercial ... pdf  
look for other books for solution

\* Homework 3rd. Due Dec, 20, 2016

Two problems, either problem 1 or problem 2 is encouraged to be chosen. Not necessary to choose both problems.

Problem 1: Research report or project on specific topic in QFT.

- 1) The topic has to be approved by instructor;
- 2) The file has to be written in PDF file (.tex, .doc), and E-mailed to Instructor.
- 3) Excellent research reports are encouraged to be talked by students <sup>+ audience</sup> in the last class

Recommended topics: ① Lorentz group and its representation.

- ② Poincaré group and its representation.

(why photon has two degrees of freedom?)

- ③ Spin - Statistics theorem. ✓

**Canonical Quantization**

- ④ Noether's theorem
- 
- (proof)

**Hamiltonian Formulation vs. Path Integral**

- ⑤ CPT theorem

- ⑥ Majorana fermion (wavy fermion)

**2. Steps of Canonical Quantization**

- (important in condensed physics and relativity) →

**Classical****Quantum****① to ⑥**

- ⑦ Vacuum energy and cosmological constant problem. ✓

- ⑧ Vacuum energy and Casimir force ✓

(refer to the text book in Quantum field)

- ⑨ Lorentz invariant Cross section and decay width.

- ⑩ Coulomb scattering in QFT.

- ⑪ Rutherford scattering in QFT. (Both in classical mechanics, QM, QFT)

$$[a, a^\dagger] = \delta \quad (\Rightarrow [a, \hat{a}] = i\delta)$$

$$\hat{H}, \hat{\vec{p}} = \hat{H}, \hat{\vec{p}}(a, a^\dagger) \Rightarrow \text{no SHO}$$

- ⑫ Dirac's formulae.
- 
- (derives it, refer to Dirac's original paper, and find simpler way)

- ⑬ Wick's theorem.

- ⑭ One-loop calculation in QED.

- ⑮ Regularization and Renormalization in QED.
- ⑯ Ward Identity in QED.  
(and gauge invariance of QED)
- ⑰ LSE reduction and S-matrix

### ③ Green function and Generating functions

$\langle \bar{\psi} \psi \rangle$  generating function for Green function and  
specific Green function

⑨ multi-loop Feynman diagrams.  
↑ it is a big subject

more topics, which are planned to be talked in the semester, but time is not enough.

Problem 2: The calculation of differential cross section in the Bhabha Scattering  $e^+ e^- \rightarrow e^+ e^-$   
(see text book of QFT)

(Luke's problem #6, problem 3)  
↑ it is simpler

(Or Peskin and Schroeder's book, problem 5.2, p.p. 170)

\* Question: why not talk about the entire contents of online lec. notes, like what has been done in the past four years?

Answer: Knowledge & Calculation is NOT that important for beginning beginners in QFT.

Basic concepts, self-study, and research experience are important for most student.

Note: In fact, for interested students, they are encouraged to check calculations in online lec. notes.  
(do it by yourself)  
(step by step)

### XI. 1. Part of Canonical Quantization

Def. Quantization: A systematic procedure of constructing quantum theory from classical theory.

→ Quantum theory is somehow depend on

(classical theory),  
it is impossible faster,

N.B.: Quantum Mechanics → Quantum Field theory

Quantum

Mechanics

Quantum

Field

Theory

→

Quantum

Field

Theory

↓

Note: Quantization is a procedure from the simplest to the complicated, or from the specific to the general, or from the boundary to the domain, hence, it is essentially a type of guesswork.

Instead of rigorous mathematical derivation.

→ QM, QFT is somehow very bad defined.

Quantization is very ugly thing!

Dof. Canonical Quantization : Construction of the Hamiltonian formulation of Quantum theory, from the Hamiltonian formulation of classical theory.

→ Because Hamiltonian is energy, but Lagrangian is not obvious.

Note: Why the Canonical Quantization

Heisenberg's Matrix QM is constructed by  
Canonical Quantization

Note: In principle, we are able to construct a number of Quantization procedures,

for example, path integral quantization, which selects the Lagrangian formulation and plays the essential roles in modern theoretical physics ( particles p. standard model condensed p. )

Note: In the Instructor's opinion, no quantization → perfect. best choices

or New Quantization → better

or Canonical / Path Integral → bad.

Questions: why choose Canonical Quantization, rather than Path Integral Quantization, for beginners in QFT?

ugly, but OK.  
essential in modern theoretical physics

best choice for beginner's  
Instructor's opinion

Answer: Comparison of Canonical Quantization and with Path Integral Quantization. ↪  
The understanding of The Action Principle  
(All the parts have contributions the Action principle is wrong)

Particle Interpretation  
Explicit (defined exactly)  
(we talk about particle physics, we must see particles)

Probability Amplitude  
Not first (particle we derived first)  
(Amplitude is not easy to understand)

Classical - quantum correspondence  
 $\rightarrow$   $t \rightarrow 0$   
Stationary phase  
Hesenberg Equations  
→ Canonical equations go back to

Formulation Hamiltonian  
(Energy → Explicit physical meaning)  
Lagrangean  
(Action → very important advantage)

Mathematics	Operator-Valued function (Everything is matrix)	ordinary function
Symmetry	Lorentz invariance is NOT explicit (Some mathematical Calculation needed)	Lorentz invariance is explicit
Causality	NOT obvious	NOT obvious
Vacuum energy	$\infty$	$\infty$
Physical Quantities (exponents, measurements)	$\infty$	<ul style="list-style-type: none"> <li>Need Extra Specification</li> <li>regularization of Canonical Quantization</li> <li>renormalization of Path Integral</li> <li>Computation is always in INFINITE situation</li> </ul>

- XI.2 Recan The Hamiltonian Formulation of CFT (Review)
- General Coordinates and Canonical Momentum  
 $\phi^a(t, \vec{x}) \rightarrow \phi$  a general field coordinate  
 $\pi^b(t, \vec{x}) \rightarrow \pi^b$  canonical field momentum  
 Satisfied the Poisson Brackets
  - $\{ \phi^a(t, \vec{x}), \phi^b(t, \vec{y}) \} = 0, \{ \pi^a(t, \vec{x}), \pi^b(t, \vec{y}) \} = 0$   
 $\{ \phi^a(t, \vec{x}), \pi^b(t, \vec{y}) \} = \delta^{ab} \delta^3(\vec{x} - \vec{y})$   
 Conserved energy  
 $H = \int d^3x \pi^a (\phi^a, \pi^a) \quad H \geq 0, H \geq 0$   
 Associated with space time transform.
  - The Canonical equations  
 $\frac{d\phi^a}{dt} = \{ \phi^a, H \} = \frac{\partial H}{\partial \pi^a}$   
 $\frac{d\pi^a}{dt} = \{ \pi^a, H \} = -\frac{\partial H}{\partial \phi^a}$   
 For Observables  $O(\phi^a, \pi^a, t)$   
 $\frac{dO}{dt} = \frac{\partial O}{\partial t} + \{ O, H \}$
  - Symmetry → translation invariance  
 $\rightarrow$  Lorentz invariance  
 $\rightarrow$  Energy is not obvious  
 Because Hamiltonian formulation select the special time!

$\{ f, g \} = \frac{\partial f}{\partial \phi^a} \frac{\partial g}{\partial \pi^a} - \frac{\partial f}{\partial \pi^a} \frac{\partial g}{\partial \phi^a}$  maybe a '-' in other courses

$\{ \phi^a, H \} = \frac{\partial H}{\partial \pi^a}$   
 $\{ \pi^a, H \} = -\frac{\partial H}{\partial \phi^a}$   
 For Observables  $O(\phi^a, \pi^a, t)$

$\frac{dO}{dt} = \frac{\partial O}{\partial t} + \{ O, H \}$

Lorentz invariants ( $\partial^\mu, \eta_{\mu\nu}$ )  
 $\rightarrow$  is not obvious  
 Because Hamiltonian formulation select the special time!

50) Causality (Special Relativity) obvious! in Maxwell equations!

50) Finite physical quantities : finite energy! Ok!

### X I 3. Canonical Quantization of Real Scalar Field Theory $\rightarrow$ one by one

- From ordinary coordinates and momenta  $\rightarrow$  field operator - valued field coordinates and momenta,
- Setting up

$$\begin{aligned} [\hat{\phi}(t, \vec{x}), \hat{\phi}(t, \vec{y})] &= 0, \quad [\hat{\pi}^a(t, \vec{x}), \hat{\pi}^b(t, \vec{y})] = 0 \\ [\hat{\phi}^a(t, \vec{x}), \hat{\pi}^b(t, \vec{y})] &= i\hbar \delta^{ab} \delta^3(\vec{x} - \vec{y}) \end{aligned}$$

- The Hamiltonian operator

$$\hat{H} \rightarrow \hat{\phi}_t^{\dagger}, \quad H \rightarrow \hat{H}$$

$$H|\psi\rangle, \langle \psi|H|\psi\rangle \geq 0, \quad \langle \psi|H|\psi\rangle \geq 0$$

Note: The operator product ordering ambiguity

$$\begin{aligned} \Phi_a \partial_a &\rightarrow \hat{\phi}^a \hat{\phi}^a & \partial_a \pi_b &\rightarrow \hat{\phi}^a \hat{\pi}_b \\ \partial_a \pi_b &\rightarrow \hat{\phi}^a \hat{\pi}_b & \pi_b \partial_a &\rightarrow \hat{\pi}_b \hat{\phi}^a \\ \text{commutative} &\quad \text{not commutative} & & \end{aligned}$$

This is unsolved and unavoidable problem in quantization.

$\leftarrow$  suffer from subtle problem, we have different choice

### 3. The Equation of motion (Heisenberg's Equation of motion)

$$\begin{aligned} i\hbar \dot{\hat{\phi}}^a &= [\hat{\phi}^a, \hat{H}] = \frac{\partial \hat{H}}{\partial \pi^a} \\ i\hbar \dot{\hat{\pi}}^a &= [\hat{\pi}^a, \hat{H}] = -\frac{\partial \hat{H}}{\partial \phi^a} \end{aligned}$$

for general observables  $\hat{O}(\hat{\phi}^a, \hat{\pi}^a, t)$

$$i\hbar \dot{\hat{O}} = i\hbar \frac{\partial}{\partial t} + [\hat{O}, \hat{H}]$$

### 4. Symmetry

Lorentz invariance NOT EXPLICIT  $\rightarrow$  maybe, we construct time-dependent one

- we talk about equal-time Commutators, and Heisenberg's form prefers the special choice of Time.
- But we can prove the Lorentz invariance.

Note: Classical Symmetries may be violated (in QFT), which is called Anomalies (check text book)

Quantization is very subtle

Oct. 22, 2016. teaching plan

### Canonical quantization (II)

Hamiltonian, Vacuum Energy, Fock space

micro causality ...

complicated.

$$\langle S'_1 | S_2 | \rangle = \int D\phi(x) \exp \frac{i}{\hbar} \int S[x]$$

$$\begin{aligned} &= \sum_{\text{all paths}} \exp i \frac{S[\text{path}]}{\hbar} \\ &\xrightarrow{\text{from } S[t] \rightarrow S'_1[t]} \end{aligned}$$

In QFT: all paths

In QFT: several paths

### Lec X I - X II Canonical Quantization

\* Answer questions on previous lec.

#### 1) Space time translation

$$\begin{aligned} &\text{space time coordinate translation} \quad x \xrightarrow{T} x' = x + \xi \\ &\text{field translation} \quad \phi(x) \xrightarrow{T} \phi(x') \equiv \tau_\xi \phi \tau_\xi^\dagger(x) = \phi(T^{-1}x) \end{aligned}$$

$\downarrow$   
general fields  
(spinor, vector, scalar...)

$$\tau_\xi \equiv \exp(-i\hbar \tau_\mu^\dagger \xi^\mu), \quad \tau_\mu^\dagger: \text{generator of space time translation}$$

expand  $\tau_\xi$  in terms of  $\xi^\mu$

$$\tau_\xi \phi \tau_\xi^\dagger(x) = (1 - i\hbar \tau_\mu^\dagger \xi^\mu) \phi (1 + i\hbar \tau_\mu^\dagger \xi^\mu) \quad (\star)$$

$$= \phi(x) - i[\tau_\mu^\dagger \xi^\mu, \phi]_{\text{order } \xi^\mu} + o(\xi^\mu)$$

and we have

$$i[\tau_\mu^\dagger \xi^\mu, \phi] = \partial_\mu \phi$$

$$\begin{aligned} \xi^\mu &\rightarrow -i\partial_\mu \\ \text{generator} &\quad \text{representation} \end{aligned}$$

#### 2) Lorentz transformation

$$\begin{aligned} &x \xrightarrow{\Lambda} x' = \Lambda x \\ &\text{coordinate} \quad \text{spinor representation} \end{aligned}$$

$$\text{field} \quad \hat{\phi}_{\mu}(x) = \hat{\phi}'_{\mu}(x) = (\tau_\mu \hat{\phi} \tau_\mu^\dagger)_{\mu\nu} = \Omega(\Lambda) \hat{\phi}_{\mu'}(\Lambda^{-1}x)$$

- $D(\alpha) = \exp(-i\hbar \omega_S \alpha \gamma_2)$   $\rightarrow$  spin angular momentum
- $\Omega(\alpha) = \exp(-i\hbar \omega_M \alpha \gamma_2)$   $\rightarrow$  angular momentum generator

- 6) Parity (mirror transf.)
- |  |  |
|--|--|
| $\vec{v} = \partial_x(x_1 - i\omega x_2)$<br>$= \partial_x(x_1 - \frac{\omega}{2}i\omega x_2) + i(\partial_{x_2} - \partial_{x_1})x_2 + O(\omega^2)$ | $t$ time component $\rightarrow$ fixed<br>$\vec{x}$ space component $\rightarrow$ $\vec{x}$ $\rightarrow$ $\vec{x}'$<br>$\vec{v} = \partial_x \vec{x} \rightarrow -\vec{v}'$ |
|--|--|
- easy to derive total angular momentum
- $$[\hat{M}_{\mu\nu}, \Phi] = (i(x_{\mu 2} - x_{\nu 1}) + S_{\mu\nu})\Phi$$
- ↓ orbital angular momentum  
generator of Lorentz representation group
- $\hat{M}_{\mu\nu}$  is operator, generator of Lorentz group.
- 7) Representation of group theory
- group  $g$ , element  $g \in G$ ,  $g: V \rightarrow V$
- $$V = \text{span}\{|i\rangle\}$$
- $$g|i\rangle = |f_i\rangle$$
- representation  $\rightarrow$  representation space  $V$
- $\downarrow$   
state basis:  $|i\rangle$  (e.g. Hilbert space)  
representation matrix  $g_{ji}$
- when talk about representation we talk about  $\{V_i\}$  together
- when talk about  $\{V_i\}$  together
- we talk about  $\{V_i\}$  together
- 8) Path Integral and action principle
- Path Integral and action principle
- Path Integral Quantization: particle  $\rightarrow$  field / wave
- Canonical Quantization: wave field  $\rightarrow$  particle
- easily describe particle picture  
(this is particle physics)
- every path has contribution
- 9) Question 9. History of Canonical Quantization
- From Replace
- |   |   |
|---|---|
| $\int d\vec{x} \psi(x)$ Heisenberg Form (from Heisenberg) | $\int d\vec{x} \psi(x)$ Poisson Bracket $\rightarrow$ Commutator (from Dirac) |
|---|---|
- History of Path Integral Quantization (Realized)  
Observed by Dirac in 1936
- defined and applied by Feynman in 1948
- e.g.  $\Phi'_a(x) = \Phi_a(x^{-1}x')$ , scalar under  $L_T$   $S_0^1 C_{1,3}$
- $\downarrow$  final
- 4) Representation of Lorentz group  $SO(1,3)$
- $SO(1,3) \approx SU(2) \otimes SU(2)$   
 $SO(1,3) \approx SU(2) \oplus SU(2)$ ?
- A good research topic for Homeland 3rd (to determine which one is right)
- which field it is?
- 5) Definition of  $\begin{cases} \text{Scalar} \\ \text{Vector} \end{cases}$ , they are just representation of a specific group
- Spinor (under transformation, its transformation Rule determine one field components)
- e.g. SO(3) invariant real scalar field theory
- $\Phi = \begin{pmatrix} \phi_1 \\ \phi_2 \\ \vdots \\ \phi_N \end{pmatrix}$  -  $d_\alpha$ ,  $\alpha = 1, 2, \dots, N$
- transformation  $\Phi'(x) = R \Phi(x)$ ,  $R$ :  $M \times N$  matrix
- $\vec{\Phi}$ : vector representation under  $SO(3)$ .  $R^T R = I$ ,  $R S_0^1 R^{-1} = S_0^1$

But in fact, in canonical quantization,  $\mu \rightarrow +\infty$ , or minus. (Bad!) we must do something additional to cancel the divergence.

\* Recall [lec XI]:

XII. Definition of Canonical Quantization

XIII. Review of Hamiltonian formulation of CFT (Classical field theory)

XIV. Canonical Quantization of Real Scalar field.

and of General field

(Vector, scalar, spinor)

1. From ordinary field functions  $\hat{\Phi}^\alpha(t, \vec{x})$ ,  $\hat{\Pi}^\alpha(t, \vec{x})$  to operator-valued field functions  $\hat{O}_a$  and

Their satisfying Equal-time Commutator / Anti-commutator:

$$[\hat{\Phi}^\alpha(t, \vec{x}), \hat{\Phi}^\beta(t, \vec{y})]_\pm = 0$$

$$[\hat{\Pi}^\alpha(t, \vec{x}), \hat{\Pi}^\beta(t, \vec{y})]_\pm = 0$$

$$[\hat{\Phi}^\alpha(t, \vec{x}), \hat{\Pi}^\beta(t, \vec{y})]_\pm = i\hbar \delta^3(\vec{x}-\vec{y}), \text{ also}$$

Theorem (spin-statistical theorem)

For Spin-Integer relativistic quantum field theory, Canonical Quantization chooses

Commutator  $[A_B B_A]_- = AB - BA$ ;

For Spin-Half Integer relativistic quantum field theory, Canonical Quantization chooses anti-commutator  $[A_B B_A]_+ = AB + BA$ ;

Note: Quantization of scalar, vector  $\rightarrow$  commutator

Quantization of Dirac-Spinors  $\rightarrow$  Anti-commutator

2.0 Hamiltonian Operator of General field

$$\mathcal{H}^\alpha = \int d^3x \hat{\Phi}^\alpha \rightarrow \hat{\Phi}^\alpha (\hat{\Phi}^\alpha, \hat{\Pi}^\beta)$$

Note: Operator-product ordering ambiguity.

Reason: Quantization is a guesswork from the special limit to general theory.  
(boundary theory)

It is impossible in principle, but it can be tested by experiment.

Remark 2: The Hamiltonian operator must be non-negative and finite.

$0 \leq \hat{H} < \infty$  (requirement for physical meaning which means, for any  $|\phi\rangle$ , the mean value  $\langle \phi | H | \phi \rangle \geq 0$

### 3. Heisenberg EOM

$$\dot{\hat{\Phi}}^\alpha = \{ \hat{\Phi}^\alpha, H \} = \frac{\partial H}{\partial \hat{\Pi}^\alpha} \rightarrow \dot{\hat{\Pi}}^\alpha = -i[\hat{\Phi}^\alpha, H] = \frac{\partial H}{\partial \hat{\Phi}^\alpha}$$

$$\dot{\Pi}^\alpha = \{ \Pi^\alpha, H \} = -\frac{\partial H}{\partial \Phi^\alpha} \quad \hat{\Lambda}^\alpha = -i[\hat{\Pi}^\alpha, H] = -\frac{\partial H}{\partial \Phi^\alpha}$$

↑ valid for Fermions both.

$$\dot{\hat{O}} = \frac{\partial}{\partial t} \hat{O} + \{ \hat{O}, H \} \rightarrow \dot{\hat{O}} = \frac{\partial}{\partial t} \hat{O} + [\hat{O}, \hat{H}]$$

4. Micro-causality and Quantum Measurement

Def: The micro-causality relation between two observables  $\hat{O}_1^\alpha(t, \vec{x})$  and  $\hat{O}_2^\beta(t, \vec{x})$ , for microscopic world

are specified by the commutator  $[\hat{O}_1^\alpha(t, \vec{x}), \hat{O}_2^\beta(t, \vec{y})] = 0$  for  $(x-y)^2 < 0$

Comments:  
 $[\Lambda^\alpha, \Omega^\beta]_- = A^\alpha B^\beta - B^\beta A^\alpha := [\Lambda^\alpha, \Omega^\beta]$   
 $[\Lambda^\alpha, \Omega^\beta]_+ = A^\alpha B^\beta + B^\beta A^\alpha$  preferred

Note:  $(x-y)^2 < 0$  two points are space-like separated.  
(no instantaneous interaction)

$[\hat{O}_1^\alpha(t, \vec{x}), \hat{O}_2^\beta(t, \vec{y})] = 0$  for  $(x-y)^2 < 0$ . The share the same eigenstates. Two quantum measurements by  $O_1$  and  $O_2$  do not interfere with each other.

Note: In canonical quantization, equal-time Commutator / Anti-commutator are specified.  
Micro-causality is NOT obvious, and must be verified.

### 5. Defining Symmetry and Anomaly

Def: Defining symmetry is observed in classical physics, and survives the quantization procedure.

such as Poincaré Symmetry (space-time and Lorentz Invariance)

Def. Anomaly is a symmetry observed in classical physics but violated in quantum physics. Such as Chiral Anomaly in Chiral Gauge Field Theory.

Note: Defining Symmetry is crucial  
Reason: They define Hamiltonian  $H = \int d^3x \hat{T}^{00} \rightarrow$   $\int d^3x \hat{T}^{00}$   $\rightarrow$  defining Symmetries.

Reason: They define Momentum  $P^\alpha = \int d^3x \hat{T}^{0i} \rightarrow$  conserved charge of space-time translation Angular momentum  $\hat{J}^{0S} = \int d^3x \hat{m}^{0S}$

Reason 2: Not obvious but must be verified  
 Physical Quantities / Expectation value  $\langle \cdot \rangle$  < 1. >  
 must be invariant under defining symmetries.  
 (transformation)

6<sup>o</sup> Finite physical quantities  $\leftarrow$  big problems in QFT  
 < 1. > expectation value  $\rightarrow$  must be finite in QFT  
 < 1. > amplitude

Note: After Canonical and Path Integral Quantization, we (unavoidably) get  $\mathcal{O}$  quantities.

What's the reason why we always get  $\mathcal{O}$  quantities?

QFT is a theory with a number of DFT  $\leftarrow$  How to overcome it?

Note: Artificial procedures Additional to Quantization are introduced/denied

- 1) Normal ordering of operator products
- 2) Regularization (Canonical)
- 3) renormalization (Path Integral)

Quantization is also artificial, but always favored by experiments (up to now).

Note: QFT is artificially constructed but always supported by High Energy Experiment up to today,

this means QFT { partly right... partly wrong }

Open question: What's well-defined QFT?

Does it exist? Does it become fundamental?

What about quantum gravity?

XI-XII: 4 Canonical Quantization of free real scalar field theory.  
 no interaction  
 Scalar on Lorentz transformation

4.1. Classical theory:

$$\text{Hamiltonian density: } \mathcal{H} = \frac{1}{2} \partial^\mu \phi \partial_\mu \phi - \frac{1}{2} m^2 \phi^2 - \frac{1}{2} \omega^2 \phi^2$$

$$\text{Conjugate momentum: } \pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \dot{\phi}$$

$$\text{K-G equation: } (\Box + m^2) \phi = 0$$

$$\text{momentum density: } \mathcal{P} = \pi_\phi \nabla \phi$$

$$\text{Hamiltonian density: } \mathcal{H} = \frac{1}{2} (\partial^\mu \phi)^2 + (m\phi)^2 + m^2 \phi^2$$

$$\text{Lagrange density: } \mathcal{L} = \frac{1}{2} (\partial^\mu \phi)^2 - \frac{1}{2} (m\phi)^2 - \frac{1}{2} \omega^2 \phi^2$$

4.2 Quantum theory QFT and Canonical Quantization

equal-time commutator  $[\hat{\phi}(t, \vec{x}), \hat{\pi}_\phi(t, \vec{x})] = i\hbar \delta^3(\vec{x} - \vec{y})$   
 (spin less,  $\zeta = 1$ )

Hamiltonian  $\hat{\mathcal{H}}(\hat{\phi}, \hat{\pi}_\phi) = \frac{1}{2} (\hat{\phi}^2 + (\partial_\mu \hat{\phi})^2 + m^2 \hat{\phi}^2)$   $\leftarrow$  no operator product ambiguity.

Hamiltonian operator  $\hat{H} = \int d^3x \hat{\mathcal{H}} = \int d^3x \frac{1}{2} (\hat{\phi}^2 + \nabla^\mu \phi \cdot \nabla_\mu \hat{\phi} + m^2 \hat{\phi}^2) \geq 0$

EOM:  $\begin{cases} \hat{\phi}(t, \vec{x}) = \hat{\pi}_\phi(t, \vec{x}) \\ \hat{\pi}_\phi(t, \vec{x}) = i [\hat{H}, \hat{\phi}(t, \vec{x})] \Rightarrow (\Box - m^2) \hat{\phi} = \hat{\pi}_\phi \end{cases}$  solved like ODE.

$$\Rightarrow (\partial_\mu \phi + m^2 \phi) \hat{\phi}(t, \vec{x}) = 0 \quad (* \text{ KG eq.})$$

Canonical equation are not explicitly Lorentz invariant.

But  $(* \text{ KG eq.})$  is Explicitly Lorentz invariant.

Next lec. Tue, Nov. 29, 2016 on shell condition  $\kappa^2 = m^2$ , to explain the meaning of  $m^2$   
 plain-wave solution Fock space, vacuum

Energy, normal ordering, micro-causality, Feynman propagator,  
 Canonical Quantization of Dirac field theory.

Lec XIII & XIV Canonical quantization

Nov 29, 2016

\* Homework 2nd, Due today

\* Homework 3rd, Due Dec. 3rd, 2016

Teaching plan  
 Lec XI-XIV: Canonical quantization of Free scalar field theory  
 4. 1. Free real scalar field  
 4. 2. Canonical quantization

Lec XVII Dec, 27, 2016  
 Feynman Rules, S-matrix and Cross-section

Lec XVIII, Jan. 3, 2017  
 Solution to Homework 3rd  
 Research reports

differentiated cross-section of  
 Bhabha Scattering  
 (Get answer in textbook)  
 of QFT

4. 3. Vacuum energy and normal ordering  
 4. 4. Fock space  
 Lec XVI  
 4. 5. Micro causality  
 4. 6. Feynman propagator  
 4. 7. Poincaré's Algebra

teaching idea, strategy  
 2012~2016  
 Calculations  
 online lec. notes  
 Solution to Homework 1st (Maxwell vector field theory)

Lec XVI Dec 13, 2016  
 Solution to Homework 1st (Maxwell vector field theory)

Lec XVI Dec 20, 2016  
 Solution to Homework 2nd (Dirac bispinor field theory)

basic concepts

Self-study  
 research experience

- What we have done on Canonical Quantization?

1) Prof. Canonical Quantization

2) Hamiltonian Formulation of General Classical field theory.

3) Canonical Quantization of General Classical Field.

find the proof of Lorentz invariant on rec. notes.

Space-time translation of Energy-momentum tensor

$$\chi' = \tau \chi = \chi + \xi$$

$$\phi'_\mu(x) = \tau_\mu \phi \tau_\nu^{-1}(x) = \phi(\chi - \xi)$$

4.1. Classical theory of real scalar field

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2, \quad m \geq 0$$

$$\bar{\pi}_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \dot{\phi}$$

$$H = \dot{\phi} \bar{\pi}_\phi - \mathcal{L} = \frac{1}{2} (\dot{\phi}^2 + (\nabla \phi)^2 + m^2 \phi^2) \geq 0$$

(energy must be non-negative number)

Box ( Klein-Gordon Equation ):

$$(\partial_\mu \partial^\mu + m^2) \phi = 0$$

Plain wave solution (of Klein-Gordon equation):

$$\phi(t, \vec{x}) = \int d^3 \vec{k} \left( \alpha(\vec{k}) e^{-ik \cdot x} + \alpha^*(\vec{k}) e^{ik \cdot x} \right)$$

just Fourier transform of

Klein-Gordon equation field

we need other constraints

$$(\square^2 + m^2) \phi = 0 \quad (\Rightarrow k^2 = m^2)$$

this is mass-shell relation

only after quantization, we will know meaning of  $m^2$

$\vec{k}$  is wave vector,  $\vec{R} = (k_x, \vec{k})$ ,  $w_{\vec{k}} = \sqrt{k^2 + m^2}$

Def.

$$\phi(\vec{k}) = (2\pi)^{3/2} \sqrt{2w_{\vec{k}}} \alpha(\vec{k}), \quad \text{define new coefficient, } \quad (\text{no particle here})$$

Def.

$$\phi(k_x, \vec{x}) = \int \frac{\partial \hat{\phi}^k}{(2\pi)^3 m^2} \frac{1}{\sqrt{2w_{\vec{k}}}} \left( \alpha(k_x) e^{-ik \cdot x} + \alpha^*(\vec{k}) e^{ik \cdot x} \right)$$

after quantization, this is 'Harmonic oscillator', and will simplify the solution

Different solution formulations exist.

Note:

$$\text{Def } \beta(\vec{k}) = (2\pi)^{3/2} \sqrt{2w_{\vec{k}}} \alpha(\vec{k})$$

$$\phi(k_x, \vec{x}) = \int \frac{\partial \hat{\phi}^k}{(2\pi)^3 m^2} \frac{1}{\sqrt{2w_{\vec{k}}}} \left( \beta(k_x) e^{-ik \cdot x} + \beta^*(\vec{k}) e^{ik \cdot x} \right)$$

thus is Lorentz invariant integral measure. ask for { ... }

3) Canonical Quantization

4.2. Canonical Quantization of real scalar field

$$\phi_\mu(x) = \tau_\mu \phi \tau_\nu^{-1}(x) = \phi(\chi - \xi)$$

( $\xi$  is particle mass)

we now can get Hamiltonian density and momentum density

$$\bar{\pi}_\phi = \dot{\phi}$$

$$T^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial \partial_\mu \phi} \partial^\mu \phi - \eta^{\mu\nu} \mathcal{L} = \partial^\mu \phi \partial^\nu \phi - \eta^{\mu\nu} \mathcal{L} \rightarrow \partial_\mu T^{\mu\nu} = 0$$

$$T^{00} = \dot{\phi}^2 - \mathcal{L} = \dot{\phi}^2 \geq 0 \quad (\text{energy})$$

$$T^{0i} = \dot{\phi} \partial^i \phi = p_i \quad (\text{momentum}), \quad \vec{p} = \nabla \phi$$

Space-time Lorentz invariance and angular-momentum tensor

$$\chi' = \tau \chi = \chi + \xi$$

$$\phi'_\mu(x) = \tau_\mu \phi \tau_\nu^{-1}(x) = \phi(\chi - \xi)$$

$$\text{the angular momentum tensor}$$

$$m_{\mu\nu\sigma} = x^\rho \partial_\mu u_\sigma - x^\sigma \partial_\mu u_\rho \rightarrow \partial_\mu m^{\mu\nu\sigma} = 0$$

$$\mathcal{J}^{\mu\nu} = \int d^3 x \mathcal{M}^{\mu\nu} \delta^3$$

$$(\text{conserved charge})$$

$$\text{Hamiltonian} \quad \hat{T}^{\mu\nu}(\hat{\phi}, \bar{\pi}_\phi) = \frac{1}{2} (\dot{\hat{\phi}}^2 + (\nabla \hat{\phi})^2 + m^2 \hat{\phi}^2) \geq 0$$

$$\text{from:} \quad \begin{cases} \hat{\phi}(t, \vec{x}) = \hat{\pi}_\phi(t, \vec{x}) & (\text{def of momentum}) \\ \dot{\hat{\phi}}(t, \vec{x}) = -i [\hat{\pi}_\phi, \hat{\phi}] = -i [\hat{\pi}_\phi, H] = -i z \hat{\phi} - m^2 \hat{\phi} & \Rightarrow (\partial_\mu \partial^\mu + m^2) \hat{\phi}(t, \vec{x}) = 0 \end{cases}$$

not explicit Lorentz invariant

$$\text{the form is explicit Lorentz invariant.}$$

$$(\text{physical result})$$

calculation of cross section

(28) • The plain wave solution (satisfying the on-shell condition  $k^2 = m^2$ )  
 $\hat{\phi}(t, \vec{x}) = \int \frac{d^3 k}{(2\pi)^3 2\omega_{\vec{k}}} \left( \hat{a}(\vec{k}) e^{-ik \cdot \vec{x}} + \hat{a}^*(\vec{k}) e^{ik \cdot \vec{x}} \right)$   
 $= \int \frac{d^3 k}{(2\pi)^3 2\omega_{\vec{k}}} \left( \beta(\vec{k}) e^{-ik \cdot \vec{x}} + \alpha^*(\vec{k}) e^{ik \cdot \vec{x}} \right)$   
 $= \int d^3 k \left( \hat{a}(\vec{k}) e^{-ik \cdot \vec{x}} + \hat{a}^*(\vec{k}) e^{ik \cdot \vec{x}} \right)$

Lemma 1: In canonical quantization, we talk about field operator  
 $\hat{\phi}(t, \vec{x}), \hat{n}_{\vec{k}}(t, \vec{x}) = i\delta^3(\vec{x}, \vec{k})$   
 $[\hat{a}(\vec{k}), \hat{a}^*(\vec{k}')] = [\hat{n}_{\vec{k}}(t, \vec{x}), \hat{n}_{\vec{k}'}(t, \vec{x}')] = 0$

Note:  $\hat{a}(\vec{k}), \hat{a}^*(\vec{k})$  are associated with Harmonic Oscillator (this is why we choose  $\beta(\vec{k}), \alpha(\vec{k})$ )

Check online lecture notes to get proof. (crucial)  
 $\hat{H} = \frac{1}{2} \int d^3 x \left( \hat{n}_{\vec{k}}^2 + (\vec{\phi})^2 + m^2 \hat{\phi}^2 \right)$   
↑ use field operator in momentum space

Note: we get Hamiltonian for simple Harmonic oscillator.

Theorem: free real scalar field theory is equivalent to infinity number of independent (quantum) simple harmonic oscillators (by  $\hat{a}(\vec{k})$  and  $\hat{a}^*(\vec{k})$ )

naive but crucial

$$\hat{H} = \frac{1}{2} \int d^3 k \left( \hat{a}(\vec{k}) \hat{a}^*(\vec{k}) + \hat{a}^*(\vec{k}) \hat{a}(\vec{k}) \right)$$

Note: we get Hamiltonian for simple Harmonic oscillator.

Theorem: free real scalar field theory is equivalent to infinity number of independent (quantum) simple harmonic oscillators (by  $a(\vec{k})$  and  $a^*(\vec{k})$ )

naive but crucial

Note: Harmonic oscillator	QFT (free real scalar field theory)
$a(\vec{k})$ lowering operator	Annihilation operator
$a^*(\vec{k})$ raising operator	Creation operator
$ 0\rangle$ ground state	$a(\vec{k}) 0\rangle = 0$
$ \text{Vacuum}\rangle$	$a(\vec{k}) \text{vac}\rangle = 0$

↓ maybe black hole  
entire theoretical physics (harmonic oscillator)

4.3 Vacuum energy and Normal ordering	$\rho_{\text{vac}}$	Vacuum energy	$E_{\text{vac}} = \int \frac{d^3 k}{(2\pi)^3} \frac{1}{2} \omega_{\vec{k}} \rho_{\text{vac}}$
• The vacuum $ \text{vac}\rangle$ denoted by no particle here	with talking about field space	problem	no DOF
• The vacuum $ \text{vac}\rangle$ denoted by harmonic oscillator	easy for Lorentz invariant decay	$\rho_{\text{vac}} \rightarrow 0$	$\rho_{\text{vac}} \rightarrow 0$
$a(\vec{k}) \text{vac}\rangle = 0$	$\forall \vec{k}$	$\forall \vec{k}$	$\forall \vec{k}$
$ \text{vac}\rangle$ is destroyed by entire domain of $a(\vec{k})$	very strong constraint normal ordering	$\forall \vec{k}$	$\forall \vec{k}$
def: $\text{NE 1}::=$	$:: \text{by boson bogoliubov when } t \rightarrow 0 \text{ both go to classical}$	$\text{NE 1}::=$	$\text{NE 1}::=$
The Hamiltonian (infinity num of SHOs)	$H = \int d^3 k \frac{1}{2} \text{tr}_{\text{WF}} \left( a(\vec{k}) a(\vec{k}) + a^*(\vec{k}) a^*(\vec{k}) \right) := \int d^3 k \omega_{\vec{k}} \rho_{\text{vac}}$	$\text{NE 1}::=$	$\text{NE 1}::=$
when use transform, computation could be easy.	$\left\{ \begin{array}{l} \left[ a(\vec{k}), a(\vec{k}') \right] = \delta^3(\vec{k} - \vec{k}') \\ \left[ \hat{a}^*(\vec{k}), \hat{a}(\vec{k}') \right] = \left[ \hat{n}_{\vec{k}}(t, \vec{x}), \hat{n}_{\vec{k}'}(t, \vec{x}') \right] = 0 \end{array} \right.$	$\downarrow$ zero point energy, infinity number of DOF or $\alpha$ , then no SHO	$\downarrow$ use choose $\beta$ or $\alpha$ , then no SHO is artificial
when use transform, computation could be easy.	$\left\{ \begin{array}{l} \left[ a(\vec{k}), a(\vec{k}') \right] = \delta^3(\vec{k} - \vec{k}') \\ \left[ \hat{a}^*(\vec{k}), \hat{a}(\vec{k}') \right] = \left[ \hat{n}_{\vec{k}}(t, \vec{x}), \hat{n}_{\vec{k}'}(t, \vec{x}') \right] = 0 \end{array} \right.$	$\downarrow$ zero point energy, infinity number of DOF or $\alpha$ , then no SHO	$\downarrow$ use choose $\beta$ or $\alpha$ , then no SHO is artificial
• The vacuum energy	$E_{\text{vac}} = \langle \text{vac}   H   \text{vac} \rangle$	$\downarrow$ Note: $\text{Dirac-delta function is defined by } \delta^3(\vec{r}) = \int \frac{d^3 k}{(2\pi)^3} e^{i\vec{k} \cdot \vec{r}}$	$\downarrow$ Note: $\text{Dirac-delta function is defined by } \delta^3(\vec{r}) = \int \frac{d^3 k}{(2\pi)^3} e^{i\vec{k} \cdot \vec{r}}$
• The vacuum energy	$E_{\text{vac}} = \int d^3 k \frac{1}{2} \text{tr}_{\text{WF}} \delta^3(\vec{r})$	$\downarrow$ so what vacuum energy $E_{\text{vac}} = V \int \frac{d^3 k}{(2\pi)^3} \frac{1}{2} \text{tr}_{\text{WF}}$	$\downarrow$ This is called Infrared divergence (associated with Volume)
• The vacuum energy	$E_{\text{vac}} = \langle \text{vac}   H   \text{vac} \rangle$	$\downarrow$ $\delta^3(\vec{r}) = \int \frac{d^3 k}{(2\pi)^3} \cdot 1 = \frac{V}{(2\pi)^3}$	$\downarrow$ This is still divergence. It is called Ultra-violet divergence.

The energy density $\rho_{\text{vac}} = \frac{E_{\text{vac}}}{V} = \int \frac{d^3 k}{(2\pi)^3} \frac{1}{2} \sqrt{\vec{k}^2 + m^2} \rightarrow +\infty$	$\vec{k} \rightarrow +\infty \quad \rho_{\text{vac}} \rightarrow +\infty$
This is still divergence.	
It is called Ultra-violet divergence.	
QFT is not well defined from the starting point.	
Remark 1: The origin of divergence in $E_{\text{vac}}$ , $\rho_{\text{vac}}$	
The box normalization condition (QFT in a box with periodic boundary condition)	

$$\text{The Hamiltonian } H = \frac{\hbar^2}{2} \frac{\hbar \omega^2}{\epsilon} \left( q_{\mu} q_{\mu}^{\dagger} + q_{\nu} q_{\nu}^{\dagger} + q_{\rho} q_{\rho}^{\dagger} + q_{\sigma} q_{\sigma}^{\dagger} \right), \quad [q_{\mu}, q_{\nu}^{\dagger}] = \delta_{\mu\nu}$$

$$= \frac{1}{2} \sum_{\mu} \hbar \omega_{\mu} q_{\mu} q_{\mu}^{\dagger} + \frac{1}{2} \sum_{\mu} \hbar \omega_{\mu}^2$$

The origin of  $\propto$  Euc. Vac is

$\propto$  number of SHO

The origin of  $\propto$  num of SHO is

Field has  $\propto$  number of DOF

Maybe there is no way for solving the problem, for creating correct QFT. Or QFT is not well defined. String theory, but also has infinity number.

Remark 2 : The Cosmological Constant problem

The observed vacuum energy of the Universe is almost zero

density

$$\rho_{\text{vac}}^{\text{experiment}} \approx 0$$

But the theoretical calculation/prediction

$$\rho_{\text{vac}}^{\text{theory}} \rightarrow \infty$$

$$\text{What's the way out? (Search google) no hope!}$$

Remark 3 : Casimir force (verified by experiment)

SFT in a box with periodic boundary conditions. When the vacuum is disturbed,

there is observed phenomenon.

Remark 4. With gravity, the vacuum energy in the current space-time can be observed. Quantum gravity must get rid of the divergence of vacuum energy.

4. 3. Normal ordering & Normal ordered product.

In flat space-time, the vacuum energy is the absolute energy, and only the relative energy difference is the observed quantity. This means the vacuum energy in the flat-space time can be removed by artificial procedures such as normal ordering.

Def. Normal ordered product  $\rightarrow$  in free scalar field theory is defined in the following way:

$$:\alpha(\mathbf{r})\alpha(\mathbf{r}') := \alpha(\mathbf{r}')\alpha(\mathbf{r}) := \alpha(\mathbf{r}')\alpha(\mathbf{r}')$$

In side normal ordering : ... , the bosonic operators for free spin-integer fields are always commutative.

As the result, the annihilation operators are always on LHS, and the creation operators are always on RHS.

$\hookrightarrow$  just artificial, in flat space-time

to get rid of infinity num of DOFs and infinity energy.

Example : The normal ordering of  $\hat{H}$

$$:\hat{H} := \int d^3\mathbf{r} \frac{1}{2} \hbar \omega_{\mathbf{k}} :q_{\mathbf{k}}^{\dagger} q_{\mathbf{k}} + q_{\mathbf{k}} q_{\mathbf{k}}^{\dagger} : = \int d^3\mathbf{r} \frac{1}{2} \hbar \omega_{\mathbf{k}} \left( q_{\mathbf{k}}^{\dagger} q_{\mathbf{k}} + q_{\mathbf{k}}^{\dagger} q_{\mathbf{k}} \right)$$

$\downarrow$  killed  $[a(\mathbf{k}), a^{\dagger}(\mathbf{k})]$

$$\tilde{E}_{\text{vac}} = \langle \text{vac} | \hat{H} | \text{vac} \rangle = 0$$

$$E_{\text{vac}} = \langle \text{vac} | \hat{H} | \text{vac} \rangle = 0$$

Now, everything becomes well-defined...  
but it can't be used to talk about gravity/cosmological problem.

Note: Normal ordering is artificial

Physics is not artificial } with normal-ordering and without normal-ordering  
the same classical theory must be obtained when  $\hbar \rightarrow 0$

(Physics can't be changed)

Check:

$$H = \frac{1}{2} (\hat{p}^2 + \hat{q}^2)$$

$$:H: = \hbar \omega \hat{q}^{\dagger} \hat{q} \quad \leftarrow \text{with normal ordering}$$

$$H = \hbar \omega (\hat{q}^{\dagger} \hat{q} + \frac{1}{2}) \quad \leftarrow \text{without normal ordering}$$

$$:H: = \hbar \omega \hat{q}^{\dagger} \hat{q} = \hbar \omega \frac{\hat{q}^{\dagger} - i\hat{p}}{\sqrt{2}} \frac{\hat{q} + i\hat{p}}{\sqrt{2}}$$

Physical Interpretation of Normal ordering

Quantum gravity must get rid of the divergence of vacuum energy.

Teaching plan

- ① Fock Space (particle is derived)
- ② micro-causality
- ③ Feynman propagator
- ④ Poincaré Algebra.

About renormalization

QFT  $\rightarrow$  divergence  $\left\{ \begin{array}{l} \text{Vacuum Energy} \Rightarrow \text{removed by normal-ordering} \\ \text{Probability Amplitude} \Rightarrow \text{renormalized by regularization} \end{array} \right.$   
(Feynman Rules, S-matrix)

III-defined  $\leftarrow$  Resultant HEP

about group

QFT (mathematics)  
Particle physics (experiments)  
III-defined supports S-matrix  
HEP (beautiful theory)

Homework 3rd , due Dec 22, 2016

Last Tuesday Canonical quantization of real free scalar field theory

#### 4.1 Classical Theory

#### 4.2 Canonical Quantization

#### 4.3 Vacuum Energy and Normal ordered procedure

Todays topic Fock space in momentum representation

#### 4.4 Fock space in coordinate representation

A.C micro causality and quantum measurement (micro  $\Rightarrow$  small particle causality  $\Rightarrow$  SR)

\* 4.1 Time-ordered product and Feynman propagator (motivation): ill divergences

\* 4.2 Realization of Poincaré Algebra. (special example of vacuum energy)

#### Subject for research report (Topic)

what about definitions of vector and spinors in  $SO(1,3)$

$$(\lambda_{\mu_2}) S(\lambda_2 \cdot b)$$

#### 4.4 Fock space in momentum representation.

Def. Fock space is the state space for quantum field theory (and particle physics), and is the state space for any number of particles, denoted by

$$\mathcal{F} = \bigoplus_{n=0}^{+\infty} \mathcal{F}_F^{(n)} = \mathcal{F}_L^{(0)} \oplus \mathcal{F}_L^{(1)} \oplus \dots \oplus \mathcal{F}_L^{(n)} \oplus \dots$$

$\mathcal{F}_L^{(n)}$ : n-particle space (vacuum)  $|vac\rangle \in \mathcal{H}^{(0)}$

$\mathcal{F}_L^{(n)}$ , usually NOT Hilbert Space in QM (particle number can be changed) and the completeness relation is strange, defined for the entire Fock space

2. Fock space defines the concept of particles for particle physics.

Def. Fock space in momentum representation

3. particles in Free Real Scalar Field Theory : freely moving spinless particles obeying the Bose-Einstein

statistics

2. Observables : The Hamiltonian  $H = \int d^3x \hat{f}^\dagger \hat{f}$  } from translation

The momentum  $\hat{P} = \int d^3x \nabla_\phi^\dagger \nabla_\phi \hat{f}$

The Angular momentum  $\hat{J} = \begin{cases} \text{spin} & \text{from L.T.} \\ \text{exist limit see P \& S's book} & \text{always zero} \end{cases}$

Normal ordering :  $\hat{H} = \int d^3k \hat{a}_k^\dagger a_k + \text{c.c.}$

Ordering :  $\hat{P} = \int d^3k \hat{k}^\dagger a_k^\dagger(k) a_k(k)$

Note: These assumption are NOT assumption : can be derived by commutator of  $a_k, a_k^\dagger$

Properties of the Entire Fock Space :

\* Feynman diagrams

portciles + propagators

Time ordered products

Time ordered

Space time

products

exist limit

see P \& S's book

always zero

Normal ordering :  $\hat{H} = \int d^3k \hat{a}_k^\dagger a_k + \text{c.c.}$

Ordering :  $\hat{P} = \int d^3k \hat{k}^\dagger a_k^\dagger(k) a_k(k)$

$\mathcal{H}^{(0)}$ : No particle space  $|vac\rangle$

$|\text{vac}\rangle \neq |0\rangle$  ground state

$\hat{P} = \int d^3k \hat{a}_k^\dagger a_k + \text{c.c.}$

$a(\vec{k})|vac\rangle = 0$ , for  $\vec{k} \neq 0 \Rightarrow$  QFT's statement

properties:  $\langle vac | vac \rangle = 1$

:  $\hat{P} |vac\rangle = 0$ , :  $\hat{H} |vac\rangle = 0$ , :  $\hat{J} |vac\rangle = 0$

:  $\hat{P} = \int d^3k \hat{a}_k^\dagger a_k + \text{c.c.}$

:  $\hat{H} = \int d^3k \hat{a}_k^\dagger \omega_k a_k + \text{c.c.}$

:  $\hat{J} = \int d^3k (\hat{a}_k^\dagger - \hat{a}_k) + \text{c.c.}$

properties:  $\langle \hat{k}' | \hat{k} \rangle = \delta^3(\vec{k} - \vec{k}') \Rightarrow$  Same as  $\alpha_{\vec{k}}$  ; equivalent to commutator below

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

$\hat{a}_k^\dagger |vac\rangle = \hat{a}_{k'}^\dagger |vac\rangle \Rightarrow \hat{a}_k^\dagger \hat{a}_{k'}^\dagger |vac\rangle = \hat{a}_{k'}^\dagger \hat{a}_k^\dagger |vac\rangle$

## 2. Thm 1

The Fock space created by  $a(\vec{k})$  and  $a^\dagger(\vec{k})$  is rotation-invariant under  $SO(3)$ , namely,  $\langle \vec{p} | \vec{r} \rangle = \delta^3(\vec{p} - \vec{r})$  is invariant under  $SO(3)$ .

- the probability amplitude is invariant under  $SO(3)$  automatically

Note:  $\int_{\mathbb{R}^3} \delta^3(\vec{p} - \vec{r}) | \vec{r} \rangle = 1 \rightarrow SO(3)$  invariant

Verifying can be got on online lec notes.

$$\delta^3(\vec{p}^B) \delta^3(\vec{p} - \vec{r}) = 1 \rightarrow SO(3)$$

$$\delta^3(\vec{p}^B) = \det(R) \det(\vec{r}), \text{ Let } (R) = 1 \quad SO(3) \text{ invariant}$$

$$\delta^3(\vec{p} - \vec{r}) = \langle \vec{p} | \vec{r} \rangle \text{ also Lorentz } SO(3) \text{ invariant}$$

## 3. Thm 2. (Construct Lorentz invariant Fock space)

The Fock space created by  $\beta(\vec{k}) = (2\pi)^{\frac{3}{2}} \sqrt{2\omega_k} a(\vec{k})$

$$\text{and } \beta^\dagger(\vec{k}) = (2\pi)^{\frac{3}{2}} \sqrt{2\omega_k} a^\dagger(\vec{k})$$

is invariant under Lorentz transformation.

namely  $\langle \vec{p} | \vec{R} \rangle = (2\pi)^{\frac{3}{2}} \omega_k \delta^3(\vec{p} - \vec{p}')$  is invariant under LT.

$$| \vec{k} \rangle = \beta^\dagger(\vec{k}) | \text{vac} \rangle, \quad | \vec{p} \rangle = \beta^\dagger(\vec{p}) | \text{vac} \rangle$$

Note: see online Lec. Notes.

$$\int_{\mathbb{R}^3} \frac{(2\pi)^3 k^3}{(2\pi)^3 2\omega_k} \cdot 2\omega_k \delta^3(\vec{p} - \vec{p}') = 1 \rightarrow \text{Lorentz invariant}$$

Feynman diagrams  
is created/constructed using  
this Lorentz invariant Fock space.  
it is what we want  
(probability amplitude)  
It is automatically LT invariant.

In particle physics calculations of probability amplitude,  
we choose Fock space by  $\beta(k)$  and  $\beta^\dagger(k)$ , for example, Feynman Rules.

## 4.5 Fock space in coordinate representation

$$\text{Statement 1: } | \vec{k} \rangle = \beta^\dagger(\vec{k}) | \text{vac} \rangle$$

means creating one-particle  
in momentum representation

Statement 2:  
 $\hat{\phi}(0, \vec{x}) | \text{vac} \rangle$  means creating a particle at  
the position  $\vec{x}$   
field operator  
in Schrödinger picture.

$$\hat{\phi}(0, \vec{x}) | \text{vac} \rangle = \int \frac{d^3 k}{(2\pi)^3 2\omega_k} (\beta(\vec{k}) e^{i\vec{k} \cdot \vec{x}} + \beta^\dagger(\vec{k}) e^{-i\vec{k} \cdot \vec{x}})$$

$$\text{Gauss: } \omega_k^2 = \sqrt{\vec{k}^2 + m^2}, \quad |\vec{k}| \ll m \Rightarrow 2\omega_k \approx 2m \quad (\text{low energy limit})$$

$$\hat{\phi}(0, \vec{x}) | \text{vac} \rangle \propto \int \frac{d^3 k}{(2\pi)^3} e^{-i\vec{k} \cdot \vec{x}} | \vec{k} \rangle \propto | \vec{x} \rangle \text{ in QM}$$

Note: In Non-relativistic QM, we define  $| \vec{x} \rangle \text{ by}$   
 $\hat{x} | \vec{x} \rangle = \vec{x} | \vec{x} \rangle$   
 $i\hat{p} | \vec{x} \rangle = \vec{p} | \vec{x} \rangle$

Note: In Non-relativistic QM, we define  $| \vec{x} \rangle \text{ by}$   
 $\hat{x} | \vec{x} \rangle = \vec{x} | \vec{x} \rangle$   
 $i\hat{p} | \vec{x} \rangle = \vec{p} | \vec{x} \rangle$   
it means  
 $|\vec{p}(\vec{0}, \vec{x}) | \text{vac} \rangle$   
is a particle at position  $\vec{x}$

$$\text{LT invariance} \Rightarrow \text{Lorentz invariance} \Rightarrow \text{Fock space} \Rightarrow \text{Lorentz invariance}$$

$$\langle \vec{p} | \vec{x} \rangle = \frac{1}{(2\pi)^3 n} e^{-i\vec{p} \cdot \vec{x}}$$

Statement 3: Normalization in coordinate rep.

$$D(x-y) = \langle \text{vac} | \hat{\phi}(x) \hat{\phi}^\dagger(y) | \text{vac} \rangle$$

(called D-function)

## 1) The probability amplitude

→ propagation of a free particle from  $y$  to  $x$   
creating a particle at  $y$  and destroyed at  $x$

Calculation:  $\hat{\phi}(t, \vec{x}) = \hat{\phi}^{(+)}(t, \vec{x}) + \hat{\phi}^{(-)}(t, \vec{x})$   
positive frequency  $\leftarrow a(\vec{k})$  associated with  $a^\dagger(\vec{k})$  → negative frequency  
frequency although we use field operator  $a(\vec{k}), a^\dagger(\vec{k})$  is important  
the commutator  $[a(\vec{k}), a^\dagger(\vec{k}')] \neq 0$

$$\begin{aligned} \phi^{(+)}(t, \vec{x}) &= \int \frac{d^3 k}{(2\pi)^3 2\omega_k} a(\vec{k}) e^{-i\omega_k t + i\vec{k} \cdot \vec{x}} \\ \phi^{(-)}(t, \vec{x}) &= \int \frac{d^3 k}{(2\pi)^3 2\omega_k} a^\dagger(\vec{k}) e^{i\omega_k t - i\vec{k} \cdot \vec{x}} \end{aligned}$$

In QM.  $i\frac{d}{dt} \leftrightarrow \hat{H}$ .

$$i\frac{d}{dt} e^{-i\omega_k t} = \omega_k \text{ positive energy}$$

$$i\frac{d}{dt} e^{i\omega_k t} = -\omega_k \text{ negative energy}$$

In relativistic QM, it suffers from negative-energy problem

In QFT, no such problem  
negative/positive energy  $\Rightarrow$  just what we call it  
just terminal

$$\begin{aligned} 2) \text{ compute } D(x-y) &= \langle \text{vac} | \hat{\phi}(x) \hat{\phi}^\dagger(y) | \text{vac} \rangle \\ &= \langle \text{vac} | [\phi^{(+)}(x), \phi^{(-)}(y)] | \text{vac} \rangle \end{aligned}$$

$$\begin{aligned}
 &= \langle \text{vac} | \int \frac{d^3 k}{(2\pi)^3} \frac{d^3 k'}{(2\pi)^3} \frac{w_k}{w_{k'}} [a(k), a^\dagger(k')] e^{-ik \cdot x + ik' \cdot y} | \text{vac} \rangle \\
 &= \langle \text{vac} | \int \frac{d^3 k^2}{(2\pi)^3 w_k^2} e^{-ik \cdot (x-y)} | \text{vac} \rangle \\
 &= [\phi^{(+)}(x), \phi^{(-)}(y)] = \int \frac{d^3 k^2}{(2\pi)^3 w_k^2} e^{-ik \cdot (x-y)} \neq 0
 \end{aligned}$$

- get exact result from text book

$D(x-y) \neq 0$  is the more important result  
 $\frac{\partial}{\partial y_\mu} D(x-y) = \langle \text{vac} | [\phi(x), \phi(y)] | \text{vac} \rangle = D(x-y)$

$$\begin{aligned}
 D(y-x) &= D(y-x) \\
 D(x-y) &|_{x=y} = D(y-x) |_{y=x} \\
 \text{Because } \phi(x, \phi(y)) &|_{x=y} = \phi(y) \phi(x) |_{x=y} \Rightarrow \text{equal time, two D-func is the same.}
 \end{aligned}$$

$$\begin{aligned}
 \text{Because } \phi(x, \phi(y)) &|_{x=y} = \phi(y) \phi(x) |_{x=y} \Rightarrow \text{equal time commutator} \\
 D(x-y) &\neq 0 \text{ for any } x, y
 \end{aligned}$$

$\hat{D}(x-y) \neq 0$  for SPACE-LIKE interval  $(x-y)^2 < 0$   
 This means causality for probability amplitude is broken (trick but interesting).

Note 4.  $\langle \text{vac} | \phi(x) \phi(y) | \text{vac} \rangle$  is NOT well defined in QFT.

$\hat{D}(x-y) = \Theta(x-y) \langle \text{vac} | \phi(x) \phi(y) | \text{vac} \rangle$  (Feynman propagator)  
 $\hat{D}_f(x-y) = D_f(x-y) + D_A(x-y)$  (Feynman propagator)

time ordering  
 $\hat{D}(x-y) = \Theta(x-y) \langle \text{vac} | \phi(x) \phi(y) | \text{vac} \rangle$  (Adiabatic)

$D_f(x-y) = D_f(x-y) + D_A(x-y)$  (Feynman propagator)

#### 4.6 Micro-causality and quantum measurement

probability violating causality, is OK!

Quantum measurement obey causality, is crucial! (violation doesn't appear in measurement)

Quantum measurement obey causality  
 the fact that two quantum measurement by  $\hat{D}_f$  and  $\hat{D}_A$  do not interfere with each other means  $[D_f, D_A] = 0$

Note: In Quantum mechanics, two quantum measurement by  $\hat{D}_f$  and  $\hat{D}_A$  do not interfere with each other means  $[D_f, D_A] = 0$   
 Bell states?

Then: Micro-causality in quantum real scalar field theory can be expressed as

$$\begin{aligned}
 &\langle \text{vac} | \delta^3 k \frac{d^3 k'}{(2\pi)^3 w_{k'}^2} [a(k), a^\dagger(k')] e^{-ik \cdot x + ik' \cdot y} | \text{vac} \rangle \\
 &= \langle \text{vac} | \int \delta^3 k^2 \frac{d^3 k'}{(2\pi)^3 w_k^2} e^{-ik \cdot (x-y)} | \text{vac} \rangle
 \end{aligned}$$

$\hat{D}(x-y) = 0$  for  $(x-y)^2 < 0$   
 It can be verified by using the supposed equal-time commutator in Canonical quantization  
 $\hat{D}(t, \vec{x}), \hat{D}(t, \vec{y}) = 0$   
 $\Rightarrow$  automatically result in micro-causality

proof: see online lecture notes

$$\begin{aligned}
 [\phi^{(+)}(x), \phi^{(+)}(y)]_{x-y, z=0} &= [\phi^{(+)}(x), \phi^{(+)}(y)] \\
 &+ [\phi^{(+)}(x), \phi^{(+)}(y)]_{xy}
 \end{aligned}$$

$$\begin{aligned}
 &= D(x-y) - D(y-x) \\
 &\stackrel{\text{D-func is Lorentz invariant}}{=} D(x-y) - D(y-x) \\
 &= [D(x-y) - D(y-x)]_{xy} \\
 &= [\phi^{(+)}(x) - \phi^{(+)}(y)]_{xy} = \phi^{(+)}(x-y)
 \end{aligned}$$

$$\begin{aligned}
 &= [\phi^{(+)}(x), \phi^{(+)}(y)]_{x-y, z=0} = \phi^{(+)}(x-y)^0
 \end{aligned}$$

Scalar particle:

particle = anti-particle  
 in QFT with charges  
 particle + anti-particle

the micro-causality predicts the existence of anti-particle

In real scalar field  $\tau$  with charge  
 $\hat{D}(x), \hat{D}(y), \hat{D}(x-y)_{x-y, z=0} = 0$   
 $\langle \text{vac} | \hat{D}(x) \hat{D}(y) | \text{vac} \rangle = \langle \text{vac} | \hat{D}(x-y) | \text{vac} \rangle$

$$\begin{aligned}
 &\stackrel{\text{In Complex scalar field}}{=} \langle \text{vac} | \hat{D}(x) \hat{D}(y) | \text{vac} \rangle \\
 &\stackrel{\text{In Complex scalar field}}{=} \langle \text{vac} | \hat{D}(x) \hat{D}(y) | \text{vac} \rangle
 \end{aligned}$$

$\hat{D}(x-y)_{x-y, z=0}$  anti-particle

Time ordering is fixed by the step function. (we call it time-ordered product).

- \* Homework 3rd, due Dec 20, 2016
- \* Final exam schedule
  - During: Jan 8-12, 2017
  - And suggestion: ASAP

Final exam  
open, classnotes & Homeworks  
Exam: Basic concepts

Recall last lecture:

- Fock space in Momentum representation.
- Rock space in Coordinate representation.

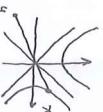
- Micro - Causality & Quantum measurement.

- Time - ordered and Feynman Propagator.

Comments on Micro - Causality

$$\text{Note 1: } D(x-y) = \langle \text{vac} | \hat{\phi}(x) \hat{\phi}^\dagger(y) | \text{vac} \rangle$$

$$\neq 0 \quad \text{for } (x-y)^2 < 0 \quad \text{space-like region}$$



Causality in SR is violated!

But not observable!

using out 1: Quantum measurement survive the micro - Causality relation:

$$[\hat{\phi}(x), \hat{\phi}^\dagger(y)]_{(x-y)^2 > 0} = 0$$

From measurement in space-like regions do not interact with each other.

Namely, particle's space-like travelling can not be observed!  
Remark: Gravity! What about micro-causality?

(Standard model has nothing to do with Gravity)

way out 2: Time ordered product

Causality (is time-ordering problem)

use impose constraint

$$0 (x^0 - y^0) \langle \text{vac} | \hat{\phi}(x) \hat{\phi}^\dagger(y) | \text{vac} \rangle$$

↑ ↗

← retarded Green function

Remarks: The micro-causality requires the existence of Anti-particles.

In a general QFT, the micro-causality relation  $[\hat{\phi}(x), \hat{\phi}^\dagger(y)]_{(x-y)^2 > 0} = 0$ .

↑ experiment  
free real scalar field theory,  
particle = anti-particle  
(anti-particle doesn't need  
to appear)

Plane wave:

$$\hat{\Phi}(x) = \int \frac{d^3 k}{(2\pi)^3 \omega_k} \left( b(\vec{k}) e^{-ikx} + c^\dagger(\vec{k}) e^{ikx} \right)$$

$$\hat{\Phi}^\dagger(y) = \int \frac{d^3 k}{(2\pi)^3 \omega_k} \left( b^\dagger(\vec{k}) e^{iky} + c(\vec{k}) e^{-iky} \right)$$

$b(\vec{k}) \rightarrow \text{annihilation of particle}$

$$[c, c^\dagger] = \delta^3$$

$$[b, b^\dagger] = \delta^3$$

$c(\vec{k}) \rightarrow \text{annihilation of anti-particle}$

$$[b, c] = 0$$

~ independent

Compute propagation of y to x:

$$\langle \text{vac} | \hat{\Phi}(x) \hat{\Phi}^\dagger(y) | \text{vac} \rangle = \int \frac{d^3 k' d^3 k}{(2\pi)^3 \omega_k \omega_{k'}} e^{-ikx} e^{iky}$$

$$\langle \text{vac} | b(\vec{k}') b^\dagger(\vec{k}) | \text{vac} \rangle \propto \delta(\vec{k} - \vec{k}')$$

$$= \int \frac{d^3 k}{(2\pi)^3 \omega_k} e^{-ik(x-y)} \neq 0$$

$$\langle \text{vac} | \hat{\phi}(x) \hat{\phi}^\dagger(y) | \text{vac} \rangle = \begin{array}{c} \vec{k} \\ \text{particle} \end{array}$$

Change the ordering:

$$\langle \text{vac} | \hat{\phi}^\dagger(y) \hat{\phi}(x) | \text{vac} \rangle \sim \langle \text{vac} | \begin{array}{c} \vec{k} \\ c(\vec{k}') c^\dagger(\vec{k}) \end{array} | \text{vac} \rangle$$

$$= \begin{array}{c} \vec{k} \\ \text{particle} \end{array} = \begin{array}{c} \vec{k} \\ \text{antiparticle} \end{array} = \int \frac{d^3 k}{(2\pi)^3 \omega_k} e^{ik(x-y)}$$

$$= \begin{array}{c} \vec{k} \\ \text{particle} \end{array} = \begin{array}{c} \vec{k} \\ \text{antiparticle} \end{array}$$

→ particle, anti-particle must creat in pair  
→ otherwise, micro-causality is killed.

Note 3: The micro-causality and CPT theorem  
Time reversal  
charge parity

$\beta_{34} - \beta_{38}$  Feynman propagator  $\begin{cases} \text{Green func.} \\ \text{Time ordered} \\ \text{Lorentz invariance} \end{cases}$

Conjugation  
(discrete symmetries)  
(particle & anti-particle)

$$\mathcal{T} [ \bar{\psi}(x) \psi(y) ] = \bar{\psi}(y) \psi(x) = \langle \text{vac} | \bar{\psi}(y) \psi(x) | \text{vac} \rangle$$

Note 4: Mathematical origin of micro-causality  
 $D(x-y) = D(y-x) \xrightarrow{\text{Lorentz transformation}} D(\gamma(x-y)) = D(\gamma(y-x))$   
 $(x-y)^2 \geq 0 \quad \text{CPT theorem}$   
 $\begin{cases} \text{proven by representation of} \\ \text{anti-particle micro-causality} \end{cases} \quad \begin{cases} \text{Lorentz group} \\ \text{relation} \end{cases}$

#### 4.6 Time-ordered product and Feynman propagator.

Feynman propagator plays a crucial role in perturbative QFT  
Time-ordered product is the way to solve the causality problem.

Def. Feynman propagator

$$\mathcal{T} \phi(x-y) = \Delta_F(x-y) = \langle \text{vac} | \mathcal{T} \phi(x) \phi(y) | \text{vac} \rangle$$

Dixon Feynman (maybe) where "T" stands for time-ordering of the operator product closely, (check Feynman's original paper to get natural understanding of  $\mathcal{T} \phi(x) \phi(y)$ )

$$\mathcal{T} \phi(x) \phi(y) = \begin{cases} \bar{\phi}(x) \phi(y) \text{ if } x < y \\ \phi(y) \bar{\phi}(x) \text{ if } y < x \end{cases}$$

Time-ordering of operator products always gives:  
operator at earliest time on the right most operator at next to earliest time on next to right most operator ... general def. of time-ordering product

operator of latest time on left most.

Lorentz invariance

$\mathcal{T} \phi(x) \phi(y) = \langle \text{vac} | \mathcal{T} \phi(x) \phi(y) | \text{vac} \rangle$

$$\begin{aligned} \mathcal{T} [ \bar{\psi}(x) \psi(y) ] &= \begin{cases} \bar{\psi}_1 \psi_2 \dots \bar{\psi}_n \psi_n & \text{from micro-causality, we} \\ \bar{\psi}_n \psi_{n-1} \dots \bar{\psi}_1 \psi_n & \text{knows time-ordering is unavoidable,} \\ \dots & \text{and it is the essential part of} \\ \dots & \text{modern QFT, we can naturally give} \\ S\text{-matrix} & \text{and compute the scattering amplitude.} \end{cases} \\ &\leftarrow \text{also, solution to divergence of } \mathcal{E}_{\text{vac}} \text{ normal-ordering product} \end{aligned}$$

2. Wick's theorem  $\mathcal{D}(x-y) = \Delta_F(x-y) = \bar{\psi}(x) \psi(y) - \bar{\psi}(y) \psi(x) \quad \mathcal{D}_F(x-y) = \Theta(x^2-y^2) \langle \text{vac} | \bar{\psi}(x) \psi(y) | \text{vac} \rangle$   
(Example form)

$$\begin{aligned} \mathcal{C} [ \bar{\psi}(x) \psi(y) ] &= \text{anti-particle / particle} \\ &\quad \text{anti-particle} \\ &= \int_{\text{vacuum}}^{\text{outgoing}} \int_{\text{vacuum}}^{\text{in}} \bar{\psi}(x) \psi(y) \quad \text{vacuum} \\ &= \int_{\text{vacuum}}^{\text{outgoing}} \int_{\text{vacuum}}^{\text{in}} \bar{\psi}(x) \psi(y) + \Theta(y^2-x^2) \mathcal{D}(y-x) \end{aligned}$$

CPT theorem  $\leftrightarrow$  micro-causality

CPT is right  $\Rightarrow$  micro-causality is right

$$\begin{aligned} \mathcal{D}(x-y) &= \int_y^x \int_x^y \int_x^y \int_y^x \cdots = \mathcal{D}(y-x) \\ &\quad \text{CPT theorem} \\ &\quad \text{CPT theorem} \leftrightarrow \text{micro-causality} \\ &\quad \text{CPT is right} \Rightarrow \text{micro-causality is right} \end{aligned}$$

$$\begin{aligned} \mathcal{T} \phi(x-y) &= \int_x^y \int_y^x \cdots \\ &= \int_x^y \int_y^x \int_x^y \int_y^x \cdots = \mathcal{D}(x-y) \end{aligned}$$

Contraction

time-ordering

normal-ordering

essential point to derive Feynman rules.

Note: This thm is special example of the Wick's thm in perturbative QFT.

Online lecture notes 2012-2013 for one semester = for two semester course.

$$\begin{aligned} \mathcal{T} \phi(x-y) &= \int_x^y \int_y^x \cdots \\ &= \int_x^y \int_y^x \int_x^y \int_y^x \cdots \quad \text{contraction between } \bar{\psi}(x) \text{ and } \psi(y) \\ &\quad \uparrow \quad \downarrow \\ &= \int_x^y \int_y^x \cdots \quad \text{time-ordering} \end{aligned}$$

Note: This thm is special example of the Wick's thm in perturbative QFT.

Note:  $\langle \text{vac} | \mathcal{T} \phi(x) \phi(y) | \text{vac} \rangle = \mathcal{D}_F(x-y)$

This means normal-ordering is killed in  $\langle \text{vac} | \dots | \text{vac} \rangle$

Normal-ordered product is automatically killed by EPV of vacuum.

Proof:  $\bar{\phi}(x) = \bar{\phi}^{(+)}(x) + \bar{\phi}^{(-)}(x)$

$$\begin{aligned} \bar{\phi}(x) \phi(y) &= : \bar{\phi}(x) \phi(y) : \\ &= \bar{\phi}^{(+)}(x) \phi^{(+)}(y) + \bar{\phi}^{(+)}(x) \phi^{(-)}(y) + \bar{\phi}^{(-)}(x) \phi^{(+)}(y) + : \bar{\phi}^{(-)}(x) \phi^{(-)}(y) : \\ &= \bar{\phi}^{(+)}(x) \phi^{(+)}(y) + \bar{\phi}^{(-)}(x) \phi^{(-)}(y) \end{aligned}$$

positive frequency negative frequency

$$N[\hat{\phi}^{(+)}_{(x)}, \hat{\phi}^{(+)}_{(y)}] = : \hat{\phi}^{(+)}_{(x)}, \hat{\phi}^{(+)}_{(y)} + \hat{\phi}^{(+)}_{(x)} \hat{\phi}^{(-)}_{(y)} + \hat{\phi}^{(-)}_{(x)} \hat{\phi}^{(+)}_{(y)} + \hat{\phi}^{(-)}_{(x)} \hat{\phi}^{(-)}_{(y)} :$$

/ annihilation  
on the right  
(kill vacuum energy)

Creation on the left

$$= \hat{\phi}^{(+)}_{(x)} \hat{\phi}^{(+)}_{(y)} + \hat{\phi}^{(+)}_{(y)} \hat{\phi}^{(+)}_{(x)} + \hat{\phi}^{(-)}_{(x)} \hat{\phi}^{(+)}_{(y)} + \hat{\phi}^{(-)}_{(y)} \hat{\phi}^{(+)}_{(x)}$$

$$+ \hat{\phi}^{(+)}_{(x)} \hat{\phi}^{(-)}_{(y)} - \hat{\phi}^{(-)}_{(x)} \hat{\phi}^{(+)}_{(y)}$$

$$= \hat{\phi}_{(x)} \hat{\phi}_{(y)} + (-1) [\hat{\phi}^{(+)}_{(x)}, \hat{\phi}^{(+)}_{(y)}]$$

Replace  $x \leftrightarrow y$

$$N[\hat{\phi}^{(+)}_{(x)}, \hat{\phi}^{(+)}_{(y)}] = \hat{\phi}^{(+)}_{(x)} \hat{\phi}^{(+)}_{(y)} - [\hat{\phi}^{(+)}_{(y)}, \hat{\phi}^{(+)}_{(x)}]$$

In normal-ordering, for bosonic-type operators,  $N[\hat{\phi}^{(+)}_{(x)}, \hat{\phi}^{(+)}_{(y)}] = N[\hat{\phi}_{(x)} \hat{\phi}_{(y)}]$

Inside time-ordering, for bosonic-type operators  $S$ ,  $T\hat{\phi}_{(x_1)} \hat{\phi}_{(x_2)} = T\hat{\phi}_{(x_2)} \hat{\phi}_{(x_1)}$

$$\hat{\phi}_{(x_1)} \hat{\phi}_{(x_2)} = N[\hat{\phi}_{(x_1)}, \hat{\phi}_{(x_2)}] + D_{(x_2-x_1)}$$

$$\left\{ \begin{array}{l} \hat{\phi}_{(x_1)} \hat{\phi}_{(x_2)} = N[\hat{\phi}_{(x_1)}, \hat{\phi}_{(x_2)}] + D_{(x_2-x_1)} \\ \hat{\phi}_{(x_1)}, \hat{\phi}_{(x_2)} = N[\hat{\phi}_{(x_1)}, \hat{\phi}_{(x_2)}] + D_{(x_2-x_1)} \end{array} \right.$$

$$T\hat{\phi}_{(x)} \hat{\phi}_{(y)} = N[\hat{\phi}_{(x)}, \hat{\phi}_{(y)}] + D_{(x-y)}$$

$$= \left\{ \begin{array}{l} \hat{\phi}_{(x>y)}, \quad \hat{\phi}_{(x<y)} = -\hat{\phi}_{(y>x)}, \hat{\phi}_{(y>x)} = D_{(x-y)} \\ \hat{\phi}_{(x < y)}, \quad \hat{\phi}_{(x>y)} = : \hat{\phi}_{(y>x)} : + D_{(y>x)} \end{array} \right.$$

$\uparrow$  see online lecture note for another type of proof.

from  $T$  to  $N$ .

Remark:

Time-ordering < for micro-causality

Simplified notation for many equations

↓ simplest form of Wick's theorem.

Wick's theorem: the time-ordered product is a linear combination of products of Feynman propagators and normal-ordered products.

$\uparrow$  key point to derive Feynman Rules

why it is useful?  
because normal-ordering kills vacuum. so the Wick's theorem greatly simplifies calculation.

Dyson's formula

{ scattering matrix  
Feynman Rules }

Excellent student self study  
Research calculation

2 semester majority (survive)

(Basic concept)

{ Semester excellent students survives  
Majority (survive)  
Majors left out }

annihilation  
on the right  
(kill vacuum energy)

$$T \text{ thm 2: Complex-variable Contour Integral}$$

$$P_F(x-y) = \int_C \frac{dk}{(2\pi)^4} \frac{i}{k^2 - m^2} e^{-ik \cdot (x-y)} = \int \frac{dk}{(2\pi)^4} \frac{i}{k^2 - m^2 + i\epsilon} e^{-ik \cdot (x-y)}$$

$\sim$  Contour

$\epsilon \rightarrow 0^+$

$x^0 y^0$ , clockwise contour  $C_1$  on the lower-half-plane of the Complex frequency plane.

frequency plane.

This is complex analysis

Complex energy!

(positive/negative energy)

Cauchy thm.

Only pole contributes to integral



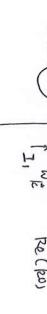
$\Rightarrow$  residue theorem

$x^0 y^0$  Anti-clockwise contour  $C_2$  on the upper-half-plane of Complex frequency plane.

Energy?



$\Rightarrow$  residue theorem



$\Rightarrow$  residue theorem

$$\text{Proof: } D_F(x-y) = \delta(x^0-y^0) D(x-y) + D(y^0-x^0) D(y-x)$$

$$D(x-y) = \int \frac{d^3 k}{(2\pi)^3 2\omega_k} e^{-ik \cdot (x-y)}$$

$$X^0 y^0, \quad D_F(x-y) = \int \frac{d^3 k}{(2\pi)^3 2\omega_k} e^{-ik \cdot (x-y)}$$

$$\text{and Complex-Variable Contour} \int_C \frac{dk}{(2\pi)^3} \frac{i}{k^2 - m^2} e^{-ik \cdot (x-y)} = \int \frac{dk}{(2\pi)^3} \int \frac{dk}{(2\pi)^3} \frac{i}{k^2 - m^2} e^{-ik \cdot (x-y)} e^{ik \cdot (y-pu)} e^{ik \cdot (y-qv)} e^{-ik \cdot (x-pu)} e^{-ik \cdot (x-qv)}$$

$$\int_C \frac{dk}{(2\pi)^3} \frac{i}{k^2 - m^2} e^{-ik \cdot (x-y)} = \int \frac{dk}{(2\pi)^3} \int \frac{dk}{(2\pi)^3} \frac{i}{k^2 - m^2} e^{-ik \cdot (x-y)} e^{ik \cdot (y-pu)} e^{ik \cdot (y-qv)} e^{-ik \cdot (x-pu)} e^{-ik \cdot (x-qv)}$$

$$= \int \frac{dk}{(2\pi)^3} \frac{i}{k^2 - m^2} e^{-ik \cdot (x-y)} = \int \frac{dk}{(2\pi)^3} \frac{i}{k^2 - m^2} e^{-ik \cdot (x-y)} = D(x-y)$$

when  $y^0 < x^0$ ,  $D_F(x-y) = \int \frac{dk}{(2\pi)^3 2\omega_k} e^{ik \cdot (x-y)} = D(y-x)$

Remark:

Teaching experiment online lecture notes (calculation)

Excellent student self study  
Research calculation





$$\left\{ \begin{array}{l} (\square + \mu_r^2) A \beta = \partial_\mu (\partial_\nu A) \\ \partial_\mu F \partial^\mu = -\mu_r^2 A \beta \end{array} \right.$$

Apply  $\partial \beta$  on both sides, we have  $\partial_\mu \partial^\mu F \partial^\nu \beta = -\mu_r^2 \partial_\mu A$   
 $\overline{\text{symmetric}} \quad \overline{\text{antisymmetric}}$

Remark 1:

when  $\mu_r^2 \neq 0$ , we have  $\partial_\mu A = 0$

when  $\mu_r^2 \neq 0$ , photon massive, Lorentz gauge fixing is intrinsic in theory.  
 (photon mass)

when  $\mu_r^2 = 0$ , photon massless,  $\partial_\mu A$  is imposed by hand

Photon is massive or massless?

Remark 2:  
 Research report: Lorentz gauge and photon mass

(Coulomb gauge must also relate to photon mass)

when  $\mu_r^2 \neq 0$ ,  $(\square + \mu_r^2) A_\mu = 0$  Klein-Gordon eq.

when  $\mu_r^2 = 0$ ,  $\square A_\mu = \partial_\mu (\partial^\nu A) + \mu_r^2 A_\mu$  Lorentz gauge fixing

Same equation  $\Rightarrow$  different physical meaning  
 (K-G)

Remark 3: Here,  $\mathcal{L} = -\frac{1}{2} (\partial_\mu A^\mu)^2 + \frac{1}{2} (\partial_\mu A)^2 + \frac{\mu_r^2}{2} A^2 \Rightarrow \partial_\mu \partial^\mu F = -\mu_r^2 A_\mu$

and other forms of  $\mathcal{L}' = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{\mu_r^2}{2} A^2 = -\frac{1}{4} F^2 + \frac{\mu_r^2}{2} A^2$

$$\text{Euler Lagrange: } \frac{\delta \mathcal{L}'}{\delta (\partial_\mu A^\mu)} = \mu_r^2 A^\mu$$

$$= \frac{\partial}{\partial (\partial_\mu A^\mu)} \left[ -\frac{1}{4} (\partial_\mu A_\nu \partial^\mu A^\nu) (\partial_\lambda A^\lambda \partial^\mu A^\mu) \right]$$

$$= -F^{\mu\nu}$$

$$\Rightarrow \mu_r^2 A^\mu = -\partial_\nu F^{\mu\nu}$$

Different Lagrangian density, but SAME Eqs. The secret is the action  $\mathcal{L} \neq \mathcal{L}'$ ,  $\mathcal{L}' = \mathcal{L} + \partial_\mu A^\mu$ ,  $\mu_r^2 A^\mu = -\partial_\nu F^{\mu\nu}$  is the same, and Eqs is derived from the action principle, the total derivative

$$S = \int d^4x \mathcal{L} = \int d^4x \mathcal{L}' = \int d^4x L + \int d^4x \partial_\mu A^\mu$$

Vanishes on the boundary after using the Gaussian theorem.

check:  $\mathcal{L} = -\frac{1}{2} (\partial_\mu A^\mu)^2 + \frac{1}{2} (\partial_\mu A)^2 + \frac{\mu_r^2}{2} A^2$

$$\mathcal{L}' = -\frac{1}{2} (\partial_\mu A^\mu)^2 + \frac{1}{2} (\partial_\mu A_\nu) (\partial^\nu A^\mu) + \frac{\mu_r^2}{2} A^2$$

$$(\partial_\mu A_\nu) (\partial^\nu A^\mu) = \partial_\mu (\partial_\nu A^\nu) - A^\mu \partial_\nu \partial^\nu A^\mu$$

$$= \partial_\mu (\partial_\nu A^\nu) - \partial^\mu (\partial_\nu A^\nu) + (\partial_\mu A)^2$$

$$= \partial_\mu (\partial_\nu A^\nu - \partial_\nu A^\mu) + (\partial_\mu A)^2$$

$\overline{\text{symmetric}} \quad \overline{\text{antisymmetric}}$   
 $\partial_\mu^\mu$ , vanishing on the boundary term in the action, not determining the form.

why  $\mathcal{L}'$  is popular than  $\mathcal{L}$ ?

The reason is that,  $F_{\mu\nu}$  is physical quantity (gauge invariant)  
 (in QFT, gauge invariant is the most important symmetry)

Hence,  $\mathcal{L} = -\frac{1}{4} F^2$  is also gauge invariant.

Get Maxwell equations:

$$\begin{cases} \text{first type: } \nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t} \\ \text{2nd type: } \nabla \cdot \vec{B} = 0 \end{cases}$$

$\Downarrow$  simple

Bianchi identity

$$\partial_\mu F_{\mu\nu} + \partial_\nu F_{\nu\mu} + \partial_\rho F_{\rho\mu} = 0$$

Lorentz invariance is more obviously gained.

Remark 1. Why only  $\partial_\mu F^{\mu\nu}$  are dynamics equations? what about Bianchi identity?  
 It is TOPOLOGY and GEOMETRY (Identity).

Prove the  $\nabla \cdot$ :

Bianchi identity is from the definition of the gauge potential.

$$\begin{cases} \vec{E} = -\nabla \phi - \frac{\partial \vec{A}}{\partial t} \\ \vec{B} = \nabla \times \vec{A} \end{cases} \quad (\vec{E}, \vec{B}) \leftrightarrow (\phi, \vec{A})$$

$$\nabla \times \vec{E} = \nabla \times (\nabla \phi) = \frac{\partial (\nabla \times \vec{A})}{\partial t} = -\partial_t^3$$

$$\nabla \cdot \vec{B} = 0$$

Because

$$\nabla \times \nabla \phi = \epsilon_{ijk} \partial_i \partial_j \phi = 0$$

$\nabla$  fully symmetric

anti-symmetric

$$\nabla \cdot \vec{B} = \nabla \cdot (\nabla \times \vec{A}) = 0. \quad (\sum_{i,j,k} \epsilon_{ijk} \partial_i A_k) = \sum_{i,j,k} \epsilon_{ijk} \partial_i \partial_j A_k = 0$$

Derive from Bianchi identity:

$$\partial = \delta, \quad \vec{r} = \vec{r}, \quad \vec{e} = \vec{e}$$

$$\partial_0 F_{ij} + \partial_i F_{0j} + \partial_j F_{0i} = 0$$

$$\Rightarrow \frac{\partial}{\partial t} (-\epsilon_{ijk} E^k) + \partial^i E^j - \partial^j E^i = 0$$

$$\Rightarrow \nabla \times \vec{E} = -\frac{\partial \vec{E}}{\partial t}$$

$$\alpha = \vec{v} \cdot \vec{E}, \quad r = k$$

$$\partial_0 F_{rs} + \partial_r F_{sj} + \partial_s F_{jr} = 0$$

$$\Rightarrow \partial_1 B_1 + \partial_2 B_2 + \partial_3 B_3 = 0$$

$$\Rightarrow \nabla \cdot \vec{B} = 0$$

Deriving first type has been done.

Derive the second type:  $\partial_k F^{0j} = 0 \Rightarrow \nabla \cdot \vec{E} = 0$

$$\text{write down } F^{0j} = \begin{pmatrix} 0 & -E_x & -E_y \\ 0 & 0 & -E_z \\ E_y & E_z & 0 \\ -E_x & -E_z & 0 \end{pmatrix} = \frac{\partial \vec{E}}{\partial x}$$

$$\alpha = \vec{v} \cdot \vec{E} = 0$$

$$\partial_2 F^{0j} = 0 \quad (\text{Gauss Law})$$

$$\begin{aligned} \partial_2 F^{0j} + \partial_j F^{02} &= 0 \\ \partial_0 F^{02} &= \partial_j (\epsilon_{ijk} B^k) = \epsilon_{ijk} \partial_j B^k = (0 \times \vec{B})^2 \end{aligned}$$

(c) The Energy-momentum tensor and angular momentum tensor

- The conserved current can be verified in the Lagrangian formulation (in the Action formulation) in different ways, all of which is called the proofs from Noether's theorem.
- Without Noether's theorem (before Noether), without Lagrangian formulation, with Newtonian dynamics (Lorentz force, work, energy). Derive conserved current.

Complex, but can be got from text book.

Reminder:  
Noether's thm (Lagrangian Formulation)  
without Noether's thm (Newtonian dynamics)

- The derivation from Lagrangian formulation:  
  
Noether's thm (Lagrangian Formulation)  
without Noether's thm (Newtonian dynamics)  
Not simple & excellent ✓

ideas:

maxwell eqs ( $\Rightarrow$  Bianchi  
first type)

first prove maxwell eq right  
 $\Rightarrow$  Bianchi right  
first prove Bianchi right  
 $\Rightarrow$  maxwell right

in first type of  $\mathcal{L}$ :

$$\frac{\partial \mathcal{L}}{\partial (A_\mu)} = -\partial^\mu A_\nu - \eta^{\mu\nu} \partial_\mu A_\nu$$

$$\frac{\partial \mathcal{L}'}{\partial (A_\mu)} = -F^{\mu\nu}$$

$$T^{\mu\nu} = -\mathcal{L} \gamma^{\mu\nu} + \frac{\partial \mathcal{L}}{\partial (A_\mu)} \partial^\mu A_\nu$$

$$\left. \begin{array}{l} \text{first type of } \mathcal{L} \\ \text{second type of } \mathcal{L}' \end{array} \right\}$$

$$\nabla^\mu = -\partial^\mu \theta + \frac{\partial \mathcal{L}}{\partial (A_\mu)} \partial^\mu A_\nu$$

In second type of  $\mathcal{L}'$ :

$$\frac{\partial \mathcal{L}'}{\partial (A_\mu)} = -F^{\mu\nu}$$

so that

$$T'{}^{\mu\nu} = \frac{1}{4} F^2 \gamma^{\mu\nu} - F^{\mu\rho} \partial_\rho A_\nu$$

(we want gauge invariant physical quantity, we define new energy-momentum tensor

$$\Theta^{\mu\nu} = T'{}^{\mu\nu} - \partial_\mu (F^{\nu\rho} A_\rho)$$

We can check  $\partial_\mu \Theta^{\mu\nu} = 0 = \partial_\mu T^{\mu\nu}$ ,  $\Theta^{\mu\nu}$  are both conserved current.

$$\uparrow \begin{array}{l} \text{symmetric} \\ \text{antisymmetric} \end{array}$$

$$= 0$$

$$\Theta^{\mu\nu} = \frac{1}{4} F^2 \gamma^{\mu\nu} + F^{\mu\rho} \partial_\rho A_\nu$$

now is gauge invariant energy-momentum tensor

5. Lorentz transformation and angular momentum tensor  $\partial_\mu M^{\mu\nu\rho\sigma} = 0$

$$\frac{\partial x^\nu}{\partial w^\sigma} = \frac{1}{2} (\eta^{\mu\rho} x_\sigma - \eta^{\mu\sigma} x_\rho), \quad \frac{\partial A^\nu}{\partial w^\sigma} = -\frac{i}{2} (\epsilon_{\mu\rho} \partial_\sigma A^\rho + \epsilon_{\mu\sigma} \partial_\rho A^\rho)$$

spin-1 tensor

$$(\Omega_{\mu\nu})_{\mu\nu} = i(\eta_{\mu\rho} \eta_{\nu}{}^{\rho} - \eta_{\mu\nu} \eta^{\rho\rho})$$

This means total angular momentum tensor

$$\frac{\partial \Omega^{\mu\nu}}{\partial w^\sigma} = \frac{1}{2} C^{\mu\rho\lambda\sigma} - \eta^{\mu\sigma} A^\rho$$

if photon is spin one, we get photon tensor in the following

$$M^{\mu\nu\rho\sigma} = L^{\mu\nu\rho\sigma} + 2 \frac{\partial \mathcal{L}}{\partial (A_\mu)} \frac{\partial A^\sigma}{\partial w^\rho}$$

$$\Rightarrow L^{\mu\nu\rho\sigma} = \epsilon \partial_T \gamma^{\mu\nu} - \sqrt{-g} \Gamma^{\mu\nu\rho\sigma}$$

Take  $L \Rightarrow L'$ ,  $\frac{\partial \mathcal{L}'}{\partial (A_\mu)} = -F^{\mu\nu}$ , result will <sup>be</sup> ready. ( $\mathcal{L}' = -\frac{1}{4} F^2$ )

$$M^{\mu\nu\rho\sigma} = L^{\mu\nu\rho\sigma} + i \frac{\partial \mathcal{L}'}{\partial F^{\mu\nu}} - F^{\mu\nu\rho\sigma} - \text{gauge invariant, but not obviously.}$$

$$\text{In Electrodynamics. } \vec{A} \rightarrow A_\mu, \quad j_\mu^\mu = -[C T^\mu - \frac{\partial \mathcal{L}}{\partial A_\mu}] \frac{\partial x^\mu}{\partial w^\mu} - \frac{\partial C}{\partial A_\mu} \frac{\partial A^\mu}{\partial w^\mu}$$

Research report : What about obviously caught invariant Angular-momentum tensor.  
( may occur in final exam )

Next lecture : Schedule for final exam  
Solution to the 2nd homework.

$$= \int d^4x \left\{ -\partial_\alpha A_\beta \partial^\alpha (\delta A^\beta) + \partial_\beta A^\beta (\partial_\alpha \delta A^\alpha) + \mu^2 A_\alpha \delta A^\alpha \right\}$$

Consider the Lagrangian for a real vector field  $A^\alpha$

$$\mathcal{L} = -\frac{1}{2} (\partial_\alpha A_\beta) (\partial^\alpha A^\beta) + \frac{1}{2} (\partial_\alpha A^\alpha) (\partial_\beta A^\beta) + \frac{1}{2} \mu^2 A_\alpha A^\alpha$$

- (a) Derive the equation of motion  $[\partial_\alpha(\square + \mu^2) - \partial_\alpha \partial_\beta] A^\beta_{(n)} = 0$ ; and verify the Lorentz condition  $\partial_\alpha A^\alpha = 0$ .
- (b) when  $\mu^2 = 0$ ,  $A^\mu = C_\mu \vec{A}$ , and  $\vec{E} = -\nabla \phi - \partial_\alpha \vec{A}$ ,  $\vec{B} = \nabla \times \vec{A}$ , Derive Maxwell equations.

- (c) With and without Noether's theorem, derive the energy-momentum tensor  $T_{\mu\nu}$  for  $\mu^2 = 0$ .

- (d) (Optional problem) With and without Noether's theorem, derive the angular-momentum tensor  $M_{\mu\nu\rho}$  for  $\mu^2 = 0$ .

Solutions:

(a)

Method 1.

From the Euler-Lagrange equation  $\frac{\partial \mathcal{L}}{\partial A^\beta} = \partial_\alpha \frac{\partial \mathcal{L}}{\partial (\partial_\alpha A^\beta)}$ ,

we compute both sides of it:

$$\begin{aligned} \frac{\partial \mathcal{L}}{\partial A^\beta} &= \frac{1}{2} \mu^2 \frac{\partial}{\partial A^\beta} (A_\alpha A^\alpha) = \frac{1}{2} \mu^2 \left( \eta_{\alpha\beta} A^\alpha + A_\alpha \delta_\beta^\alpha \right) = \mu^2 A_\beta \\ \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} &= -\frac{1}{2} \left( \delta_\mu^\alpha \eta_{\alpha\beta} \right) \partial^\alpha A^\beta - \frac{1}{2} \partial_\mu A_\beta \left( \eta^{\mu\rho} \delta_\rho^\sigma \right) + \frac{1}{2} \left( \delta_\mu^\alpha \delta_\rho^\sigma \right) \partial_\sigma A^\beta \\ &+ \frac{1}{2} \partial_\rho A^\rho \left( \delta_\mu^\alpha \delta_\beta^\sigma \right) = -\partial^\mu \partial_\beta A_\alpha + \partial_\alpha \partial_\beta A^\mu \\ \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} &= -\partial^\mu \partial_\mu \partial_\beta A_\alpha + \partial_\alpha \partial_\beta A^\mu = 0 \end{aligned}$$

Substituted these into Euler-Lagrange equation we obtain

$$(\mu^2 + \partial^\mu \partial_\mu) A_\beta - \partial_\alpha \partial_\beta A^\alpha + \mu^2 \partial_\alpha A_\beta = 0$$

$$\Leftrightarrow (\square + \mu^2) \eta_{\alpha\beta} A^\alpha - \partial_\alpha \partial_\beta A^\alpha = 0$$

The EOM has been derived above.

Method 2. From the action principle  $\delta S = \int d^4x \delta \mathcal{L} = 0$ .

We compute the variation of action  $S$ :

$$\delta S = \int d^4x \delta \mathcal{L} = \int d^4x \left\{ -\frac{1}{2} \partial_\alpha (\delta A_\beta) \cdot \partial^\alpha A_\beta - \frac{1}{2} \partial_\alpha A^\alpha \partial^\beta (\delta A_\beta) + \frac{1}{2} \partial_\alpha (\delta A^\alpha) \cdot \partial_\beta A^\beta + \frac{1}{2} \partial_\alpha A^\alpha \partial_\beta (\delta A^\beta) + \frac{1}{2} \partial_\alpha (\delta A^\alpha) \cdot \partial_\beta A^\beta + \frac{1}{2} \partial_\alpha A^\alpha \partial_\beta (\delta A^\beta) \right\}$$

$$\Rightarrow \nabla \times \vec{B} = \partial_t \vec{E} \quad \dots \textcircled{1}$$

$$\begin{aligned} \delta S &= \int d^4x \left\{ \partial^\alpha \partial_\alpha (\delta A^\beta) - \partial_\alpha \partial_\beta (\delta A^\alpha) + \partial_\alpha (\delta A^\beta) \partial_\beta \delta A^\alpha + \mu^2 \delta A_\alpha \delta A^\alpha \right\} \\ &\quad + \partial^\alpha \partial_\alpha \delta A_\beta \cdot \delta A^\beta - \partial_\alpha \partial_\beta \delta A_\beta \cdot \delta A^\alpha \end{aligned}$$

The totally derivative vanish because the fixed boundary, so that

$$\begin{aligned} \delta S &= \int d^4x \left\{ \partial^\alpha \partial_\alpha \delta A_\beta - \partial_\alpha \partial_\beta \delta A_\beta + \mu^2 \delta A_\alpha \right\} \delta A^\alpha \\ \text{Since } \delta A^\alpha \text{ are arbitrary, we have} \\ \delta S &= 0 \end{aligned}$$

$$\Leftrightarrow \partial^\mu \partial_\mu A_\alpha - \partial_\alpha \partial_\beta A^\beta + \mu^2 \partial_\alpha A_\alpha = 0$$

$$\Leftrightarrow \{(\square + \mu^2) \eta_{\alpha\beta} - \partial_\alpha \partial_\beta\} A^\beta = 0$$

That is EOM.

(b) Check the Lorentz Condition:

$$\text{use EOM } \partial^\mu \partial_\mu A_\alpha - \partial_\alpha \partial_\beta A^\beta + \mu^2 \partial_\alpha A_\alpha = 0 \quad \text{let } \partial^\alpha \text{ act both sides of it then}$$

we obtain

$$\partial^\alpha \partial^\mu \partial_\mu A_\alpha - \partial^\alpha \partial_\beta \partial_\beta A^\beta + \mu^2 \partial^\alpha A_\alpha = 0$$

Because we can change the order of derivative, which means  $\partial^\alpha \partial^\mu \partial_\mu A_\alpha = \partial^\alpha \partial_\beta \partial_\beta A^\beta$ .

So when  $\mu^2 \neq 0$ , that Lorentz condition can be proved  $\partial^\alpha A_\alpha = 0$ .

When  $\mu^2 = 0$ , the EOM becomes  $\partial^\mu \partial_\mu A_\alpha - \partial_\alpha \partial_\mu A^\mu = 0$ . Let  $\partial_\mu A_\alpha - \partial_\alpha A_\mu = F_{\mu\alpha}$

and the EOM becomes  $\partial^\mu F_{\mu\alpha} = 0$ .

Next we shows the relation between  $F_{\mu\alpha}$  and  $\vec{E}, \vec{B}$ .

$$\vec{E}_2 = -\partial_2 \phi - \frac{\partial A^2}{\partial t} = -\partial_2 A_0 + \partial_0 A_2 = F_{02}$$

$$\begin{aligned} \sum_{mni} B_i &= \sum_{mni} (\nabla \times \vec{A})_i = -\sum_{mni} \epsilon_{ijk} \partial_j A_k = -(\delta_{mj} \partial_{ik} - \delta_{mk} \partial_{ij}) \partial_j A_k \\ &= -\partial_m A_n + \partial_n A_m = -F_{mn} \end{aligned}$$

So that from the EOM  $\partial^\mu F_{\mu\alpha} = 0$ ,

$$\text{when } \alpha = 0 \Rightarrow \partial^\mu F_{\mu 0} = \partial^2 F_{10} = -\partial_2 F_{10} = \partial_2 E_2 = \nabla \cdot \vec{E} = 0 \quad \dots \textcircled{1}$$

$$\begin{aligned} \text{when } \alpha = k = 1, 2, 3 \Rightarrow \partial^\mu F_{\mu k} &= \partial_2 E_k + \partial^2 F_{2k} = \partial_2 E_k + \partial^2 \sum_{i=1}^2 F_{ik} = (\partial_t \vec{E} - \nabla \times \vec{B})_k = 0 \end{aligned}$$

And the Bianchi identity is automatically contained in the definition of  $F_{\mu\nu}$

$$\begin{aligned} \partial_\nu F_{\mu\nu} &= \partial_\nu F_{[\mu\nu]} = \partial_\nu (\partial_\mu A_\nu - \partial_\nu A_\mu) \\ &= \partial_\nu (\partial_\mu A_\nu - \partial_\nu A_\mu) = 0 \end{aligned}$$

Let's show  $\partial_\nu F_{\mu\nu} = 0$  are equivalent to another two Maxwell equations

$$\begin{aligned} \partial_\nu F_{ij} &= 0 \Rightarrow \partial_\nu \epsilon_{ijk} B_k + \partial_j B_i - \partial_i B_j = 0 \Rightarrow \partial_2 \tilde{E}_3 - \partial_3 \tilde{E}_2 = -\epsilon_{ijk} \partial_i B_k \\ &\Rightarrow \nabla \times \tilde{E}^i = -\partial_i \tilde{B}^j \quad \dots \textcircled{3} \\ \partial_\nu F_{23} &= 0 \Rightarrow \partial_1 B_2 - \partial_2 B_1 = 0 \Rightarrow \nabla \cdot \tilde{B} = 0 \quad \dots \textcircled{4} \end{aligned}$$

$\textcircled{3}\textcircled{4}$  are Maxwell equations.

(c) Derive the Energy-momentum tensor

Method 1. With Noether's theorem

When we make space-time translation  $x'^\mu = x^\mu + \epsilon^\mu$ , the charge of  $A^\beta c_x$ ,

$$\delta_\nu A^\beta c_x = A^\beta (x - \epsilon) - A^\beta c_x = \frac{\partial A^\beta}{\partial x^\nu} (-\epsilon^\rho)$$

But  $\delta_\nu A^\beta c_x = A^\beta c_{x'}, -A^\beta c_x = 0$ .

From Noether's theorem, the conserved current is

$$T^\mu_a = -\left( \mathcal{L} \eta^\mu - \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \partial_\mu A^\beta \right) \frac{-\delta x^\rho}{-\delta u^a} = \frac{-\delta x^\rho}{-\delta u^a} - \mathcal{L} \eta^\mu + \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \partial_\mu A^\beta$$

So the conserved current - energy-momentum tensor is

$$T_{\mu\nu} = -\mathcal{L} \eta_{\mu\nu} + \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \partial_\mu A^\beta \quad \dots \textcircled{5}$$

$$\begin{aligned} &= -\eta_{\mu\nu} \left( -\frac{1}{2} \partial_\alpha A_\beta \partial_\gamma A^\beta + \frac{1}{2} \partial_\alpha A^\alpha \partial_\beta A^\beta \right) + \partial_\beta A^\beta \partial_\mu A_\mu - \partial_\mu A_\beta \partial_\mu A^\beta \dots \textcircled{5} \\ &= \frac{1}{4} \eta_{\mu\nu} (F^{\alpha\beta} F_{\alpha\gamma} + 2\partial^\alpha A^\beta \partial_\beta A_\alpha) + (-F_{\mu\beta} - \partial_\beta A_\mu) (F_{\nu\beta} + \partial^\beta A_\nu) \dots \textcircled{6} \end{aligned}$$

The physical energy-momentum tensor should directly contain  $A_\mu$ , but should be expressed by observable quantities  $F_{\mu\nu}$ , so we use Lorentz condition to cancel the  $A_\mu$  term in  $\textcircled{6}$  and obtain

$$T'_{\mu\nu} = \frac{1}{4} \eta_{\mu\nu} F^{\alpha\beta} F_{\alpha\beta} - F_{\mu\beta} F_{\nu\beta}$$

Because  $(F_{\alpha\beta}) = \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & -B_3 & B_2 \\ 0 & B_3 & 0 & -B_1 \\ -E_2 & -B_2 & B_1 & 0 \end{pmatrix}$ , so the  $T'_{\mu\nu}$  can be written as

$$\begin{aligned} [T'_{\mu\nu}] &= \frac{1}{4} (\eta_{\mu\nu}) \left( -2\tilde{E}^2 + 2\tilde{B}^2 \right) + \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & -B_3 & B_2 \\ 0 & B_3 & 0 & -B_1 \\ -E_2 & -B_2 & B_1 & 0 \end{pmatrix} \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ E_1 & 0 & +B_3 & -B_2 \\ -B_3 & 0 & 0 & +B_1 \\ E_3 & +B_2 & -B_1 & 0 \end{pmatrix} = \begin{pmatrix} \frac{1}{2} (\tilde{E}^4 + \tilde{E}^2) & -E \times \tilde{E}^2 \\ -(\tilde{E} \times \tilde{B})^T & \frac{1}{2} (\tilde{E}^2 + \tilde{E}^2) \times \tilde{B} - \tilde{E} \tilde{E} - \tilde{B} \tilde{B} \end{pmatrix} \dots \textcircled{7} \end{aligned}$$

Method 2. Without Noether's theorem, but start directly from the space-time translation

Coordinate transformation :  $x'^\mu = x^\mu + \epsilon^\mu$

the charge of  $A^\beta c_x$  :  $\delta_\nu A^\beta = A^\beta (x'^\mu - \epsilon^\mu) - A^\beta c_{x'} = \partial_\mu A^\beta \cdot (-\epsilon^\mu)$

So the charge of  $\mathcal{L}$  can be written as

$$\begin{aligned} \delta_\nu \mathcal{L} &= \mathcal{L} (x'^\mu - \epsilon^\mu) - \mathcal{L} (x'^\mu) = \partial_\mu \mathcal{L} \cdot (-\epsilon^\mu) \\ &= \frac{\partial \mathcal{L}}{\partial A^\beta} \delta A^\beta + \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \delta (\partial_\mu A^\beta) \\ &\stackrel{\text{use varying equation}}{=} \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \right) \delta A^\beta + \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \delta_\mu \delta A^\beta = \partial_\mu \left[ \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \delta A^\beta \right] \end{aligned}$$

From the equation above, we obtain

$$\partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \partial_\mu A^\beta - \mathcal{L} \eta^\mu \right) (-\epsilon^\rho) = 0$$

Because  $-\epsilon^\rho$  is arbitrary, we have the conserved current

$$T^\mu = \frac{\partial \mathcal{L}}{\partial (\partial_\mu A^\beta)} \partial_\mu A^\beta - \mathcal{L} \eta^\mu$$

It is the same as equation  $\textcircled{6}$ , so we have already derived the correct energy-momentum tensor.

$$\begin{aligned} \text{calculate in detail} & \begin{pmatrix} \tilde{E}^2 - E_2 \tilde{B}_3 + E_3 \tilde{B}_2 & -E_3 \tilde{B}_1 + E_1 \tilde{B}_3 & -E_1 \tilde{B}_2 + E_2 \tilde{B}_1 \\ \frac{1}{2} (\tilde{E}^2 - \tilde{B}^2) & -E_2 \tilde{B}_3 + \tilde{B}^2 & -E_1 \tilde{B}_3 - B_3 \tilde{B}_2 \\ \frac{1}{2} (\tilde{E}^2 - \tilde{B}^2) & -E_2 \tilde{B}_1 + E_1 \tilde{B}_2 & -E_2 \tilde{B}_1 - B_2 \tilde{B}_1 \end{pmatrix} + \begin{pmatrix} E^2 - E_2 \tilde{B}_3 + E_3 \tilde{B}_2 & -E_2 \tilde{B}_1 + E_1 \tilde{B}_3 & -E_1 \tilde{B}_2 + E_2 \tilde{B}_1 \\ -E_2 \tilde{B}_3 + \tilde{B}^2 & -E_1 \tilde{B}_3 - B_3 \tilde{B}_2 & -E_1 \tilde{B}_3 + B_3 \tilde{B}_2 \\ -E_2 \tilde{B}_1 + E_1 \tilde{B}_2 & -E_2 \tilde{B}_1 - B_2 \tilde{B}_1 & -E_2 \tilde{B}_1 + B_2 \tilde{B}_1 \end{pmatrix} \\ &= \frac{1}{2} (\tilde{B}^2 + \tilde{E}^2) T_{\mu\nu} + \begin{pmatrix} 0 & -\tilde{E} \times \tilde{B} \\ -(\tilde{E} \times \tilde{B})^T & -\tilde{E} \tilde{E} - \tilde{B} \tilde{B} \end{pmatrix} \end{aligned}$$

Method 3. Derive Energy-momentum tensor directly from Maxwell equations - (assume  $c=\mu_0=\epsilon_0=1$ )

- We begin with Lorentz force formula  $\vec{f} = \rho \vec{E} + \vec{\rho} \vec{v} \times \vec{B}$ .

The energy change of source particle is

$$\begin{aligned} \frac{dU}{dt} &= \int d\vec{r} \cdot \vec{v} = \int d\vec{r} \int^{\vec{v}} \vec{v} \cdot \vec{E} \\ &= \int d\vec{r} \left( \nabla \times \vec{B} - \frac{\partial \vec{E}}{\partial t} \right) \cdot \vec{E} = \int d\vec{r} \left\{ \nabla \cdot (\vec{B} \times \vec{E}) + \vec{B} \cdot (\nabla \times \vec{E}) - \vec{E} \cdot \frac{\partial \vec{E}}{\partial t} \right\} \\ &= \int d\vec{r} \left\{ \nabla \cdot (\vec{B} \times \vec{E}) - \vec{B} \cdot \frac{\partial \vec{E}}{\partial t} - \vec{E} \cdot \frac{\partial \vec{E}}{\partial t} \right\} \\ \text{use Gaussian theorem, } \int_V \nabla \cdot (\vec{B} \times \vec{E}) d\vec{r} &= \int_V \vec{B} \cdot (\nabla \times \vec{E}) \text{ vanish, and } \vec{E} \cdot \frac{\partial \vec{E}}{\partial t} = \frac{\partial}{\partial t} \frac{1}{2} \vec{E}^2, \\ \vec{B} \cdot \frac{\partial \vec{E}}{\partial t} &= \frac{\partial}{\partial t} \frac{1}{2} \vec{B}^2 \end{aligned}$$

So we obtain

$$\frac{dU}{dt} + \int d\vec{r} \frac{\partial}{\partial t} \frac{1}{2} (\vec{E}^2 + \vec{B}^2) = 0$$

If we believe the energy is conserved, then  $\mathcal{L} = \frac{1}{2} (\vec{E}^2 + \vec{B}^2)$  must be the energy density of electric-magnetic field.

We then derive the energy current  $\vec{j}$ :

$$\begin{aligned} \frac{\partial \mathcal{L}}{\partial t} &= \frac{\partial}{\partial t} \left( \frac{1}{2} \vec{E}^2 + \frac{1}{2} \vec{B}^2 \right) = \vec{E} \cdot \frac{\partial \vec{E}}{\partial t} + \vec{B} \cdot \frac{\partial \vec{B}}{\partial t} = \vec{E} \cdot (\nabla \times \vec{B}) + \vec{B} \cdot (-\nabla \times \vec{E}) \\ \text{So we can obtain } \frac{\partial \mathcal{L}}{\partial t} + \nabla \cdot (\vec{E} \times \vec{B}) &= 0 \end{aligned}$$

This is a continuous equation,  $\vec{E} \times \vec{B} = \vec{j}$  must be energy current.

we then derive the momentum current  $\frac{\partial \mathcal{L}}{\partial \vec{v}}$ :

$$\begin{aligned} \frac{\partial \vec{j}}{\partial t} &= \frac{\partial}{\partial t} (\vec{E} \times \vec{B}) = (\nabla \times \vec{B}) \times \vec{B} = -\vec{J} \times \vec{E} \\ &= \vec{B} \cdot \vec{B} - (\vec{B} \cdot \vec{B}) \times \vec{B} = \vec{E} \times (\nabla \times \vec{E}) \\ \nabla \cdot \vec{E} = 0 &\Rightarrow \vec{B} \cdot \vec{B} = 0 \\ \vec{B} \cdot \vec{B} = 0 &\Rightarrow \vec{J} = \nabla \cdot (\vec{E} \times \vec{B}) + \nabla \cdot (\vec{E} \times \vec{E}) - \frac{1}{2} \nabla \cdot (\vec{B} \cdot \vec{B}) - \frac{1}{2} \nabla \cdot (\vec{E} \cdot \vec{E}) \\ &= \nabla \cdot (\vec{B} \vec{B} + \vec{E} \vec{E} - \frac{1}{2} \vec{B}^2 \mathbb{I} - \frac{1}{2} \vec{E}^2 \mathbb{I}) \end{aligned}$$

So we can get

$$\frac{\partial \vec{j}}{\partial t} + \nabla \cdot \left( \frac{1}{2} \vec{B}^2 \mathbb{I} + \frac{1}{2} \vec{E}^2 \mathbb{I} - \vec{B} \vec{B} - \vec{E} \vec{E} \right) = 0$$

This is a continuous equation,  $\vec{J} = \frac{1}{2} \vec{B}^2 \mathbb{I} + \frac{1}{2} \vec{E}^2 \mathbb{I} - \vec{B} \vec{B} - \vec{E} \vec{E}$  must be momentum current.

Finally, we join  $\vec{u}, \vec{v}, \vec{F}$  to form a  $4 \times 4$  matrix

$$[T'_{\mu\nu}] = \begin{pmatrix} u & -\vec{j} \\ -\vec{j}^T & \vec{J} \end{pmatrix} = \begin{pmatrix} \frac{1}{2} (\vec{B}^2 + \vec{E}^2) & -\vec{E} \times \vec{B} \\ -(\vec{E} \times \vec{B})^T & \frac{1}{2} (\vec{B}^2 + \vec{E}^2) \mathbb{I} - \vec{B} \vec{B} - \vec{E} \vec{E} \end{pmatrix}$$

It looks the same as equation (8), so it must be a real tensor, that is, it's energy-momentum tensor.

(d) Derive the angular momentum tensor.

Method 1. Using Noether's theorem

With Lorentz transformation  $\pi' = \lambda \pi$ , the charge of  $A^{\beta(x)}$  is

$$\delta_0 A^{\beta(x)} = \lambda^{\beta(x)} \lambda^{-1} \pi - \lambda^{\beta(x)} \pi_0 + \delta A^{\beta(x)}$$

$$\frac{\delta A^{\beta}}{\delta u^{\mu 0}} = -\frac{1}{2} (\delta_0 \sigma) \pi_\mu \pi_\nu \gamma^{\beta \mu} = \frac{1}{2} (\gamma_{\mu \nu} \gamma_0^\beta - \gamma_{\nu \mu} \gamma_0^\beta) \pi_\mu \gamma^{\beta \nu}$$

$$\lambda^{\beta(x)} \lambda^{-1} \pi - \lambda^{\beta(x)} \pi_0 = \frac{\partial A^{\beta}}{\partial \pi} \left\{ \frac{1}{2} (\pi_0 \pi_0 - \pi_0 \pi_0) \pi^\lambda \right\} (\delta u^{\mu \nu})$$

So, use Noether's thm and we obtain the conserved current

$$M^{\mu \nu} = \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} \frac{\partial A^{\beta}}{\partial x^\lambda} \left\{ -\frac{1}{2} (\pi_0 \pi_0 - \pi_0 \pi_0) \pi^\lambda \right\} - \mathcal{L} \frac{\partial}{\partial x^\lambda} \left\{ -\frac{1}{2} (\pi_0 \pi_0 - \pi_0 \pi_0) \pi^\lambda \right\}$$

$$= -\frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} \frac{1}{2} (\pi^{\mu \nu} \pi_0 - \pi_0 \pi_0) \pi^\lambda \frac{\partial}{\partial x^\lambda} \quad \dots \quad (6)$$

$$\begin{aligned} &= T^{\mu \nu} \left\{ -\frac{1}{2} (\pi_0 \pi_0 - \pi_0 \pi_0) \right\} - \frac{1}{2} \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} A_0 + \frac{1}{2} \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} A_0 \\ &= -\frac{1}{2} \left( \pi_0 T^{\mu}_0 - \pi_0 T^{\mu}_0 \right) - \frac{1}{2} \left( \frac{\partial \mathcal{L}}{\partial (g^{\mu \nu} A^{\beta})} A_0 + \frac{1}{2} \frac{\partial \mathcal{L}}{\partial (g^{\mu \nu} A^{\beta})} A_0 \right) \end{aligned}$$

Then we obtain the angular momentum tensor

$$M^{\mu \nu \sigma} = -\frac{1}{2} (\pi_0 T^{\mu}_0 - \pi_0 T^{\mu}_0) - \frac{1}{2} \left( \frac{\partial \mathcal{L}}{\partial (g^{\mu \nu} A^{\beta})} A_0 + \frac{1}{2} \frac{\partial \mathcal{L}}{\partial (g^{\mu \nu} A^{\beta})} A_0 \right) \quad \dots \quad (7)$$

Method 2. Using the Lorentz transformation, but without Noether's thm

With Lorentz transformation  $\pi' = \lambda \pi$ , the charge of  $A^{\beta}$  is

$$\delta_0 A^{\beta(x)} = \lambda^{\beta(x)} \lambda^{-1} \pi - \lambda^{\beta(x)} \pi_0 + \delta A^{\beta(x)}$$

$\lambda^{\beta(x)} \lambda^{-1} \pi - \lambda^{\beta(x)} \pi_0 = \frac{\partial A^{\beta}}{\partial \pi} \left\{ \frac{1}{2} (\pi_0 \pi_0 - \pi_0 \pi_0) \pi^\lambda \right\}$  see lecture notes

$$\begin{aligned} &= \frac{\partial A^{\beta}}{\partial \pi} \frac{1}{2} \omega^{\mu \nu} \left( \pi_0 \pi_0 - \pi_0 \pi_0 \right) \pi^\lambda + \left( -\frac{1}{2} \right) \omega^{\mu \nu} \left( \pi_0 \pi_0 \right)_\mu^\lambda A_0 \gamma^{\mu \beta} \\ &= \nabla_\mu A^\beta \frac{1}{2} \omega^{\mu \nu} \left( \pi_0 \pi_0 - \pi_0 \pi_0 \right) \pi^\lambda + \left( -\frac{1}{2} \right) \omega^{\mu \nu} \left( \pi_0 \pi_0 \right)_\mu^\lambda A_0 \gamma^{\mu \beta} \end{aligned} \quad \dots \quad (8)$$

So the change of  $\mathcal{L}$  is

$$\begin{aligned} \delta_0 \mathcal{L} &= \mathcal{L} (\lambda^{-1} \pi) - \mathcal{L} (\pi) = \partial_\mu \mathcal{L} \cdot \frac{1}{2} \omega^{\mu \nu} (\pi_0 \pi_0 - \pi_0 \pi_0) \pi^\lambda \gamma^{\mu \beta} \\ &= \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} \delta_0 A^{\beta} + \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} \delta_0 A^{\beta} + \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} \pi_0 (\delta_0 A^{\beta}) \\ &= \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} \delta_0 A^{\beta} \right) \dots \quad (9) \end{aligned}$$

Substituting we have

$$\begin{aligned} (1) = (2) \Rightarrow \partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (g_{\mu \nu} A^{\beta})} \left\{ \frac{1}{2} \omega^{\mu \nu} (\pi_0 \pi_0 - \pi_0 \pi_0) \pi^\lambda \right\} \right) &= 0 \\ -\mathcal{L}_{,\frac{1}{2}} \omega^{\mu \nu} \left( \pi_0 \pi_0 - \pi_0 \pi_0 \right) \pi^\lambda \gamma^{\mu \beta} &= 0 \end{aligned}$$

Cancel the arbitrary parameter ( $\omega \rho_0$ ) from both sides of above equation we have the conserved current  $\partial_\mu \nu \rho_0 = 0$ , and

$$\begin{aligned} M_{\mu\nu} &= \frac{\partial L}{\partial(\partial_\mu A_\nu)} \left( -\frac{1}{2} \cdot \partial_\mu \partial_\nu (\partial_\lambda A^\lambda) + (\partial_\mu \partial_\nu - \partial_\nu \partial_\mu) A^\lambda \right) \\ &\quad - \frac{1}{2} \frac{\partial L}{\partial(\partial_\mu A^\lambda)} \left( \eta_{\mu\lambda} \partial_\nu^2 - \eta_{\nu\lambda} \partial_\mu^2 \right) A_\lambda \eta^{\mu\nu} \\ &\quad + \frac{1}{2} \mathcal{L} T_{\mu\nu} \left( \nu_0 \delta_{\mu\nu} - \nu_0 \delta_{\mu\nu} \right) \nu_0 \end{aligned}$$

which is the same as equation (10), so use the same steps in method 1, we obtain

$$M_{\mu\nu\rho} = -\frac{1}{2} \left( \nu_0 T_{\mu\nu} - \nu_0 T_{\nu\mu} \right) - \frac{1}{2} \left( \frac{\partial L}{\partial(\partial_\mu A_\nu)} A_\nu - \frac{\partial L}{\partial(\partial_\nu A_\mu)} A_\mu \right) \dots \quad (12)$$

where  $\frac{\partial L}{\partial(\partial_\mu A_\nu)}$  problem has computed  $\boxed{\boxed{\boxed{\quad}}} = \partial_\mu A_\nu + \partial_\nu A_\mu$ .

Method 3. Directly use Maxwell equations, we has already gained  $\vec{J} = \vec{E} \times \vec{B}$  is the momentum density so  $\vec{r} \times \vec{J}$  must be angular-momentum density, and

$$\frac{\partial L}{\partial t} (\vec{r} \times \vec{J}) = \vec{r} \times \frac{\partial L}{\partial t} \vec{J} = -\vec{r} \times (\nabla \times \vec{F}) \stackrel{\nabla \times \vec{F} = 0}{=} -\nabla \cdot (\vec{r} \times \vec{F})$$

where  $\vec{F} = \omega \vec{U} - \vec{E} \vec{E} - \vec{B} \vec{B}$  has been derived in problem (C), so we have

$$\frac{\partial}{\partial t} (\vec{r} \times \vec{J}) + \nabla \cdot (\vec{r} \times \vec{F}) = 0.$$

It is a continuous equation, so that we know  $\vec{J} = \vec{r} \times \vec{F}$  must be angular momentum current. This formula of angular momentum tensor is look like the first term of equation (12), but a factor  $-\frac{1}{2}$  and a spin-term is missing.

#### Reference:

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Problem 1. Peskin and Schroeder's book, page 39-40

- 1.1 Verify (3.18) to satisfy the Lorentz Algebra (3.17)
- 1.2 Verify (3.20) and (3.21) as examples of (3.19)

Solution:

1.1 The generator here for representation of four vector is

$$(J^{\mu\nu})_{\alpha\beta} = i(\delta^\mu{}_\alpha \delta^\nu{}_\beta - \delta^\nu{}_\alpha \delta^\mu{}_\beta)$$

namely, we lift the script  $\alpha$  and obtain

$$(J^{\mu\nu})^\alpha{}_\beta = i(\delta^\mu{}_\alpha \delta^\nu{}_\beta - \delta^\nu{}_\alpha \delta^\mu{}_\beta)$$

We then directly verify it to satisfy the Lorentz Algebra (3.17):

$$[J^{\mu\nu}, J^{\rho\sigma}]^\alpha{}_\gamma = (J^{\mu\nu})^\alpha{}_\beta (J^{\rho\sigma})^\beta{}_\gamma - (J^{\rho\sigma})^\alpha{}_\beta (J^{\mu\nu})^\beta{}_\gamma$$

$$= i(\gamma^{\mu\alpha}\delta^\nu{}_\beta - \gamma^{\nu\alpha}\delta^\mu{}_\beta) \cdot i(\gamma^{\rho\beta}\delta^\sigma{}_\gamma - \gamma^{\sigma\beta}\delta^\rho{}_\gamma)$$

$$= i(\gamma^{\rho\alpha}\delta^\sigma{}_\beta - \gamma^{\sigma\alpha}\delta^\rho{}_\beta) \cdot i(\gamma^{\mu\beta}\delta^\nu{}_\gamma - \gamma^{\nu\beta}\delta^\mu{}_\gamma)$$

$$\stackrel{\text{expand}}{=} i \cdot i \left( \gamma^{\mu\alpha}\gamma^{\rho\beta}\delta^\nu{}_\gamma - \gamma^{\nu\alpha}\gamma^{\rho\beta}\delta^\mu{}_\gamma - \gamma^{\mu\alpha}\gamma^{\sigma\beta}\delta^\rho{}_\gamma + \gamma^{\nu\alpha}\gamma^{\sigma\beta}\delta^\rho{}_\gamma - \gamma^{\rho\alpha}\gamma^{\mu\beta}\delta^\nu{}_\gamma + \gamma^{\rho\alpha}\gamma^{\nu\beta}\delta^\mu{}_\gamma + \gamma^{\rho\alpha}\gamma^{\sigma\beta}\delta^\nu{}_\gamma - \gamma^{\sigma\alpha}\gamma^{\nu\beta}\delta^\mu{}_\gamma \right)$$

$$= i \left( \gamma^{\mu\alpha}i(\gamma^{\rho\beta}\delta^\nu{}_\gamma - \gamma^{\nu\beta}\delta^\rho{}_\gamma) + \gamma^{\rho\alpha}i(-\gamma^{\nu\beta}\delta^\gamma{}_\gamma + \gamma^{\sigma\beta}\delta^\mu{}_\gamma) + \gamma^{\sigma\alpha}i(-\gamma^{\nu\beta}\delta^\rho{}_\gamma + \gamma^{\rho\beta}\delta^\nu{}_\gamma) \right)$$

must satisfy the Lorentz Algebra.

(Step 1) Let's firstly prove the commutator  $[S^{\mu\nu}, V^\rho] = i(\gamma^{\mu\rho} - \gamma^{\nu\rho})$

$$[S^{\mu\nu}, V^\rho] = \frac{i}{4} [\gamma^{\mu\nu} - \gamma^{\nu\mu}, V^\rho]$$

$$= \frac{i}{4} \left( \gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho - \gamma^{\mu\nu}V^\rho + \gamma^{\nu\mu}V^\rho \right)$$

This is just Lorentz Algebra. Here we use  $\gamma^\mu$  to denote to metric rather than  $g^{\mu\nu}$ .

1.2. We first verify (3.20) as a example of  $(V')^\alpha = (\mathbb{I} - \frac{i}{2} J^{\mu\nu} \omega_{\mu\nu})^\alpha_\beta V^\beta$

$\omega_{12} = -\omega_{21} = 0$  ( $i\omega$  represents a rotation along  $x$ -axis), we have

$$\begin{aligned} (V')^\alpha &= \left\{ \delta^\alpha{}_\beta - \frac{i}{2} \left( J^{12} \cdot \mathbb{I} - J^{21} \cdot \mathbb{I} \right)^\alpha{}_\beta \right\} V^\beta \\ &= V^\alpha - \frac{i}{2} \cdot 2\omega \cdot (J^{12})^\alpha{}_\beta V^\beta \\ &= V^\alpha + \omega \left( \eta^{1\alpha} \delta^2{}_\beta - \eta^{2\alpha} \delta^1{}_\beta \right) V^\beta \end{aligned}$$

$$\begin{aligned} \mathbb{I} \omega J^{12} \cdot \mathbb{I} &= (1 - \frac{i}{2} J^{\mu\nu} \omega_{\mu\nu})^\alpha{}_\beta V^\beta \\ V^{1\alpha} &= (1 - \frac{i}{2} J^{\mu\nu} \omega_{\mu\nu})^\alpha{}_\beta V^\beta \end{aligned}$$

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$$\begin{aligned} (V')^\alpha &= \begin{pmatrix} V^0 & \\ V^1 & -\theta V^2 \\ V^2 & +\theta V^1 \end{pmatrix} = \begin{pmatrix} 1 & -\theta & \\ 0 & 1 & \\ \theta & 0 & 1 \end{pmatrix} \begin{pmatrix} V^0 \\ V^1 \\ V^2 \end{pmatrix} = \begin{pmatrix} 1 & -\theta & \\ 0 & 1 & \\ \theta & 0 & 1 \end{pmatrix} (\mathbb{I} V) \\ &= V^\alpha + \theta \left( \eta^{0\alpha} \delta^1{}_\beta - \eta^{1\alpha} \delta^0{}_\beta \right) V^\beta = V^\alpha + \theta \left( \eta^{0\alpha} V^1 - \eta^{1\alpha} V^0 \right) \end{aligned}$$

We next verify (3.21). Set  $\omega_{\alpha\beta} = -\omega_{\beta\alpha} = \beta$  ( $\omega$  represents a boost along  $x$ -axis), and we obtain

$$(V')^\alpha = \left\{ \delta^\alpha{}_\beta - \frac{i}{2} (J^{01})^\alpha{}_\beta \right\} V^\beta = V^\alpha - \frac{i}{2} \cdot 2\beta (J^{01})^\alpha{}_\beta V^\beta$$

$$\begin{aligned} (V')^\alpha &= \left( \frac{V^0 + \beta V^1}{\sqrt{3}} \right) = \begin{pmatrix} 1 & \beta & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} V^0 \\ V^1 \\ V^2 \end{pmatrix} = \begin{pmatrix} 1 & \beta & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} (V) \end{aligned}$$

Problem 2. Peskin and Schroeder's book Page 40-42

2.1 Verify (3.23) to satisfy the Lorentz Algebra (3.17).

2.2 Verify (3.26) and (3.27), (3.29).

2.3 Verify Dirac equation (3.31) and Lagrangian (3.34) Lorentz invariant.

Solution:

2.1. We have assumed that  $\{v^\mu, v^\nu\} = v^\mu v^\nu + v^\nu v^\mu = 2g^{\mu\nu} \mathbb{I}_{4\times 4}$ . Let's prove  $S^{\mu\nu} = \frac{i}{2} [v^\mu, v^\nu]$

must satisfy the Lorentz Algebra.

$$[S^{\mu\nu}, V^\rho] = \frac{i}{4} [\gamma^{\mu\nu} - \gamma^{\nu\mu}, V^\rho]$$

$$= \frac{i}{4} \left( \gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho - \gamma^{\mu\nu}V^\rho + \gamma^{\nu\mu}V^\rho \right)$$

$$= \frac{i}{4} \left( \gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho - \gamma^{\mu\nu}V^\rho + \gamma^{\nu\mu}V^\rho \right)$$

$$\downarrow \text{odd two term} + 2\gamma^{\mu\nu}V^\rho - 2\gamma^{\nu\mu}V^\rho$$

$$= \frac{i}{4} \left( 2\gamma^{\mu\nu}V^\rho - 2\gamma^{\nu\mu}V^\rho \right)$$

$$= \frac{i}{4} \left( 4\gamma^{\mu\nu}V^\rho - 4\gamma^{\nu\mu}V^\rho \right) = i(\gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho)$$

$$= \frac{i}{4} \left( 2\gamma^{\mu\nu}V^\rho - 2\gamma^{\nu\mu}V^\rho \right)$$

$$= \frac{i}{4} \left( 4\gamma^{\mu\nu}V^\rho - 4\gamma^{\nu\mu}V^\rho \right) = i(\gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho)$$

$$= \frac{i}{4} \left( 2\gamma^{\mu\nu}V^\rho - 2\gamma^{\nu\mu}V^\rho \right)$$

$$= \frac{i}{4} \left( 4\gamma^{\mu\nu}V^\rho - 4\gamma^{\nu\mu}V^\rho \right) = i(\gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho)$$

$$= \frac{i}{4} \left( 2\gamma^{\mu\nu}V^\rho - 2\gamma^{\nu\mu}V^\rho \right)$$

$$= \frac{i}{4} \left( 4\gamma^{\mu\nu}V^\rho - 4\gamma^{\nu\mu}V^\rho \right) = i(\gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho)$$

$$= \frac{i}{4} \left( 2\gamma^{\mu\nu}V^\rho - 2\gamma^{\nu\mu}V^\rho \right)$$

$$= \frac{i}{4} \left( 4\gamma^{\mu\nu}V^\rho - 4\gamma^{\nu\mu}V^\rho \right) = i(\gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho)$$

$$= \frac{i}{4} \left( 2\gamma^{\mu\nu}V^\rho - 2\gamma^{\nu\mu}V^\rho \right)$$

$$= \frac{i}{4} \left( 4\gamma^{\mu\nu}V^\rho - 4\gamma^{\nu\mu}V^\rho \right) = i(\gamma^{\mu\nu}V^\rho - \gamma^{\nu\mu}V^\rho)$$

(Step 2) use the conclusion of step 1 to prove the original problem:

$$\begin{aligned}
 [S^{\mu\nu}, S^{\rho\sigma}] &= i \left[ S^{\mu\nu}, S^{\rho\sigma} \right] = i \left[ S^{\mu\nu}, \gamma^\rho \gamma^\sigma + \gamma^\rho [S^{\mu\nu}, \gamma^\sigma] - \gamma^\sigma [S^{\mu\nu}, \gamma^\rho] \right] \\
 &= i \frac{i}{4} \left\{ S^{\mu\nu}, \gamma^\rho (\gamma^\sigma - \gamma^\nu g^{\nu\sigma}) \gamma^\tau + \gamma^\rho (S^{\mu\nu}, \gamma^\sigma) \gamma^\tau - [S^{\mu\nu}, \gamma^\rho] \gamma^\tau \right\} \\
 &= -i \frac{i}{4} \left\{ i(\gamma^\mu g^{\nu\rho} - \gamma^\nu g^{\mu\rho}) \gamma^\tau + \gamma^\rho (S^{\mu\nu}, \gamma^\sigma) \gamma^\tau \right\} \\
 &\quad - \gamma^\sigma \cdot i \left( \gamma^\mu g^{\nu\rho} - \gamma^\nu g^{\mu\rho} \right) - i \left( \gamma^\mu g^{\nu\sigma} - \gamma^\nu g^{\mu\sigma} \right) \gamma^\tau \\
 &= i \frac{i}{4} \cdot i \left\{ (\gamma^\mu g^{\nu\sigma} - \gamma^\nu g^{\mu\sigma}) \gamma^\rho \gamma^\tau + (\gamma^\rho \gamma^\tau - \gamma^\sigma \gamma^\nu) g^{\mu\rho} \right. \\
 &\quad \left. + \left( \gamma^\mu \gamma^\nu - \gamma^\nu \gamma^\mu \right) g^{\rho\sigma} + \left( \gamma^\rho \gamma^\sigma - \gamma^\sigma \gamma^\rho \right) g^{\mu\nu} \right\}
 \end{aligned}$$

use definition of  $S^{\mu\nu}$ , and  $S^{\mu\nu} = -S^{\nu\mu}$

$$= i \left( g^{\mu\rho} S^{\nu\sigma} - g^{\mu\sigma} S^{\nu\rho} - g^{\nu\sigma} S^{\mu\rho} + g^{\mu\nu} S^{\rho\sigma} \right)$$

use definition of  $S^{\mu\nu}$ , and  $S^{\mu\nu} = -S^{\nu\mu}$

2.2 To verify (3.26) and (3.27), use the 4-matrices here

$$\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & 2x_2 \end{pmatrix} \quad \gamma^i = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix}$$

(step 2)

We can directly compute

$$\begin{aligned}
 \gamma^0 \gamma^i &= \begin{pmatrix} 1 & 0 \\ 0 & \sigma^i \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & \sigma^i \end{pmatrix} = \begin{pmatrix} -\sigma^i & 0 \\ 0 & \sigma^i \end{pmatrix} \quad \gamma^0 \gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} = \text{id}_{4\times 4} \\
 \gamma^i \gamma^0 &= \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} = \begin{pmatrix} \sigma^i & 0 \\ 0 & \sigma^i \end{pmatrix} \\
 \gamma^i \gamma^j &= \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix} \begin{pmatrix} 0 & \sigma^j \\ -\sigma^j & 0 \end{pmatrix} = \begin{pmatrix} -i\varepsilon^{ijk}\gamma^k - \delta^{ij} & 0 \\ 0 & -i\varepsilon^{ijk}\gamma^k - \delta^{ij} \end{pmatrix} \\
 \gamma^i \gamma^j &= \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix} \begin{pmatrix} 0 & \sigma^j \\ -\sigma^j & 0 \end{pmatrix} = \begin{pmatrix} i\varepsilon^{ijk}\gamma^k - \delta^{ij} & 0 \\ 0 & i\varepsilon^{ijk}\gamma^k - \delta^{ij} \end{pmatrix}
 \end{aligned}$$

(step 2)

To get the generator  $S^{\mu\nu}$ , we must first verify  $\gamma^\mu$  satisfy  $\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}$ .

From Step 1, it is easy to show  $\gamma^0 \gamma^i + \gamma^i \gamma^0 = 0$ ,

so we have

$$\{\gamma^0, \gamma^i\} = 2g^{\mu\nu} \text{id}_{4\times 4}$$

(step 3)

Next we compute  $S^{\mu\nu}$ :

$$\begin{aligned}
 S^{\mu\nu} &= \frac{i}{4} [\gamma^0, \gamma^i] = \frac{i}{4} \left( \gamma^0 (\gamma^i - \gamma^i g^{\mu\nu}) - \left( \begin{array}{cc} -\sigma^i & 0 \\ 0 & -5^i \end{array} \right) \right) \\
 &= \frac{i}{4} \left( -2\sigma^i \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} - \begin{pmatrix} \sigma^i & 0 \\ 0 & -\sigma^i \end{pmatrix} \right) \\
 &= -\frac{i}{2} \begin{pmatrix} \sigma^i & 0 \\ 0 & -\sigma^i \end{pmatrix} = \begin{pmatrix} \sigma^i & 0 \\ 0 & -\sigma^i \end{pmatrix}
 \end{aligned}$$

And because  $S^{\mu\nu} = -S^{\nu\mu}$ , so  $S^{\mu\nu}$  must equals to 0.

(step 4) We now prove (3.29)  $\Lambda_{\mu_1}^{-1} \gamma^\mu \Lambda_{\mu_2} = \Lambda_{\mu_2}^{-1} \gamma^\mu \dots$  ①, where  $\Lambda_\mu = e^{-i\omega_\mu S^{\mu\nu}}$

We have already prove  $[S^{\mu\nu}, \gamma^\mu] = i(\gamma^\mu g_{\alpha\mu} - \gamma^\nu g_{\mu\nu})$  in 2.1 (step 1), we rewrite it into a more compact form

$$\begin{aligned}
 [\gamma^\mu, S^{\mu\nu}] &= +i(-i\gamma^\mu g_{\alpha\mu} + i\gamma^\nu g_{\mu\nu}) = i(g_{\mu\alpha} \delta^\mu_\alpha - g_{\mu\nu} \delta^\mu_\nu) \gamma^\nu \\
 &= (\gamma^\mu \delta^\nu_\nu - \gamma^\nu \delta^\mu_\nu) g_{\mu\nu} + (\delta^\mu_\nu - \delta^\nu_\mu) g_{\mu\nu} \\
 &\stackrel{\text{definition of } S^{\mu\nu}}{=} (\gamma^\mu \delta^\nu_\nu - \gamma^\nu \delta^\mu_\nu) g_{\mu\nu} + (\delta^\mu_\nu - \delta^\nu_\mu) g_{\mu\nu}
 \end{aligned}$$

where  $(\gamma^\mu \delta^\nu_\nu - \gamma^\nu \delta^\mu_\nu)$  is generator of 4-vector representation of LG.

Now we can verify the infinitesimal form of ①:

$$\Lambda_{\mu_1}^{-1} \gamma^\mu \Lambda_{\mu_2} = \Lambda_{\mu_2}^{-1} \gamma^\mu$$

$$\begin{aligned}
 \Leftrightarrow & \left( 1 + i \frac{i}{2} (W^{\mu\nu} S^{\nu\sigma}) \gamma^\mu \right) \gamma^\mu \left( 1 - i \frac{i}{2} (W^{\mu\nu} S^{\nu\sigma}) \gamma^\mu \right) = \left[ \delta^\mu_\mu - i \frac{i}{2} W^{\mu\nu} S^{\nu\sigma} \right] \gamma^\mu \\
 \Leftrightarrow & \gamma^\mu - i \frac{i}{2} W^{\mu\nu} [f_\nu, S^{\nu\sigma}] + o(w^2) = \left[ \delta^\mu_\mu - i \frac{i}{2} W^{\mu\nu} S^{\nu\sigma} \right] \gamma^\mu
 \end{aligned}$$

which is always true.

We can also verify the finite form of ①. Use identity  $\partial^A B \bar{C}^{-A} = \sum_{k=0}^n \frac{1}{k!} [A^{(k)}, B]$

where  $[A^{(R)}, B] = [A, [A^{(R-1)}, B]]$  and  $[A^{(0)}, B] = B$ . We directly compute

$$\begin{aligned}
 \Lambda_{\mu_1}^{-1} \gamma^\mu \Lambda_{\mu_2} &= e^{\frac{i}{2} \mu_\mu S^{\mu\nu} \gamma^\nu} e^{\frac{i}{2} \nu_\nu S^{\nu\sigma} \gamma^\sigma} e^{-\frac{i}{2} \nu_\sigma S^{\sigma\rho} \gamma^\rho} \\
 &= \sum_{k=0}^n \frac{1}{k!} \left[ \left( \frac{i}{2} W^{\mu\nu} S^{\nu\sigma} \right)^{(k)} \gamma^\mu \right] \gamma^\nu \gamma^\sigma \gamma^\rho \\
 &= \sum_{k=0}^n \frac{1}{k!} \left[ \left( -\frac{i}{2} W^{\mu\nu} S^{\nu\sigma} \right)^{(k)} \right] \gamma^\mu \gamma^\nu \gamma^\sigma \gamma^\rho \\
 &= \left( -\frac{i}{2} W^{\mu\nu} S^{\nu\sigma} \right) \gamma^\mu \gamma^\nu \gamma^\sigma \gamma^\rho \\
 &= \left( -\frac{i}{2} W^{\mu\nu} S^{\nu\sigma} \right)^{(2)} \gamma_\mu \gamma_\nu \gamma_\sigma \gamma_\rho \\
 &= \left[ \left( -\frac{i}{2} W^{\mu\nu} S^{\nu\sigma} \right)^{(2)} \right] \gamma_\mu \gamma_\nu \gamma_\sigma \gamma_\rho
 \end{aligned}$$

and so on ...

2.3. To verify the Dirac eq. and Lagrangian Lorentz invariant, we do a transformation of coordinate  $\pi \rightarrow \lambda \pi$ , and the transformation of field  $\psi \rightarrow \lambda^{-1} \psi(\lambda^{-1}\pi) = \psi'(\pi)$ .

We ought to prove that: if and only if  $\psi'(\pi)$  satisfy the Dirac equation, then the field  $\psi(\pi)$  must also satisfy Dirac equation.

(Step 1) Verify the Lorentz invariance of Dirac equation:

$$(i \gamma^\mu \partial_\mu - m) \psi'(\pi) = 0$$

$$\begin{aligned} & \left\langle \begin{aligned} & (i \gamma^\mu \partial_\mu - m) \wedge'_2 \psi(\lambda^{-1}\pi) = 0 \\ & \psi'(\pi) = [\wedge'_2 \psi(\lambda^{-1}\pi)]^+ \gamma^0 = \psi^+(\lambda^{-1}\pi) \wedge'_2 \gamma^0 \end{aligned} \right\rangle \\ & \left\langle \begin{aligned} & (i \gamma^\mu \wedge'_2 \cdot (\lambda^{-1})^\sigma \partial_\sigma \psi - \wedge'_2 m \psi)_{(\lambda^{-1}\pi)}^+ = 0 \\ & = \partial_\mu [\psi(\lambda^{-1}\pi)] \end{aligned} \right\rangle \\ & = \partial_\mu [\psi(\lambda^{-1}\pi)] (\partial_\sigma \psi) (\lambda^{-1}\pi) \end{aligned}$$

$$\begin{aligned} & \left\langle \begin{aligned} & \wedge'_2 \left( i \gamma_2^{-1} \gamma^\mu \wedge'_2 (\lambda^{-1})^\sigma \partial_\sigma \psi - m \psi \right)_{(\lambda^{-1}\pi)}^+ = 0 \\ & = (\lambda^{-1})^\sigma \delta_\mu^\nu (\partial_\nu \psi) (\lambda^{-1}\pi) \end{aligned} \right\rangle \\ & = \psi^+(\lambda^{-1}\pi) \gamma^0 \left( 1 + \frac{i}{2} \omega_{\mu\nu} \omega^{\mu\nu} \right) \gamma^0 \end{aligned}$$

$$\begin{aligned} & \left\langle \begin{aligned} & \wedge'_2 \left( i \gamma_2^{-1} \gamma^\mu \wedge'_2 (\lambda^{-1})^\sigma \partial_\sigma \psi - m \psi \right)_{(\lambda^{-1}\pi)}^+ = 0 \\ & = \psi^+(\lambda^{-1}\pi) \gamma^0 \wedge'_2 \gamma^0 \end{aligned} \right\rangle \\ & = \psi^+(\lambda^{-1}\pi) \gamma^0 \wedge'_2 \gamma^0 \end{aligned}$$

$$\begin{aligned} & \left\langle \begin{aligned} & \wedge'_2 \left( i \gamma_2^{-1} \gamma^\mu \wedge'_2 (\lambda^{-1})^\sigma \partial_\sigma \psi - m \psi \right)_{(\lambda^{-1}\pi)}^+ = 0 \\ & = \psi^+(\lambda^{-1}\pi) \gamma^0 \wedge'_2 \gamma^0 \end{aligned} \right\rangle \\ & = \psi^+(\lambda^{-1}\pi) \gamma^0 \wedge'_2 \gamma^0 \end{aligned}$$

$$\begin{aligned} & \left\langle \begin{aligned} & [(i \gamma^\mu \partial_\mu - m) \psi]_{(\lambda^{-1}\pi)} = 0 \\ & = \psi(\lambda^{-1}\pi) \end{aligned} \right\rangle \end{aligned}$$

Which means if  $\psi'(\pi)$  satisfy Dirac eq. in new coordinate system, then

$\psi(\pi)$  must also satisfy Dirac eq. in old coordinate system.

(Step 2) To verify the Lorentz invariance of Lagrangian, we first prove

$$\gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma = \gamma^\sigma \gamma^\mu \gamma^\nu \gamma^\rho$$

In solution 2.2 (Step 1) we have verified that  $\gamma^\mu \gamma^\nu = -\gamma^\nu \gamma^\mu$ , so we have

$$\begin{aligned} & [\gamma^\mu, \gamma^\nu]^+ = \gamma^\mu \gamma^\nu \gamma^\rho - \gamma^\nu \gamma^\mu \gamma^\rho = \gamma^\mu \gamma^\rho \gamma^\nu - \gamma^\nu \gamma^\mu \gamma^\rho = 0 \\ & \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma = \frac{i}{4} [\gamma^\mu, \gamma^\nu \gamma^\rho \gamma^\sigma] = \frac{i}{4} \left( \begin{array}{c} \gamma^\mu \gamma^\nu \gamma^\rho - \gamma^\nu \gamma^\mu \gamma^\rho \\ + \gamma^\mu \gamma^\rho \gamma^\nu - \gamma^\rho \gamma^\mu \gamma^\nu \end{array} \right) \xrightarrow{\text{commute twice}} \\ & = \frac{i}{4} \left( \begin{array}{c} \gamma^\mu \gamma^\nu \gamma^\rho - \gamma^\nu \gamma^\mu \gamma^\rho \\ + \gamma^\mu \gamma^\rho \gamma^\nu - \gamma^\rho \gamma^\mu \gamma^\nu \end{array} \right) \end{aligned}$$

act dagger to commutator we have

$$= 0$$

$$\begin{aligned} & [\gamma^\mu, \gamma^\nu]^+ = \gamma^\mu \gamma^\nu \gamma^\rho - \gamma^\nu \gamma^\mu \gamma^\rho = \gamma^\mu \gamma^\rho \gamma^\nu - \gamma^\nu \gamma^\mu \gamma^\rho = 0 \\ & \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma = \frac{i}{4} [\gamma^\mu, \gamma^\nu \gamma^\rho \gamma^\sigma] \xrightarrow{\text{commute once}} \\ & = \frac{i}{4} \left( \begin{array}{c} \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma - \gamma^\nu \gamma^\mu \gamma^\rho \gamma^\sigma \\ + \gamma^\mu \gamma^\rho \gamma^\nu \gamma^\sigma - \gamma^\rho \gamma^\mu \gamma^\nu \gamma^\sigma \end{array} \right) = 0. \end{aligned}$$

And for  $\gamma_0 \gamma_i$  we have

$$\begin{aligned} & [\gamma_0, \gamma_i]^+ = \gamma_0 \gamma_i + \gamma_i \gamma_0 = \gamma_0 \gamma_i + \gamma_i \gamma_0 = 0 \\ & \gamma_0 \gamma_i \gamma_j \gamma_k = \frac{i}{4} [\gamma_0, \gamma_i \gamma_j \gamma_k] = \frac{i}{4} \left( \begin{array}{c} \gamma_0 \gamma_i \gamma_j \gamma_k - \gamma_i \gamma_0 \gamma_j \gamma_k \\ + \gamma_0 \gamma_j \gamma_i \gamma_k - \gamma_j \gamma_0 \gamma_i \gamma_k \end{array} \right) = 0. \end{aligned}$$

Act dagger to anticommutator we have

$$\begin{aligned} & \gamma_0 \gamma_i \gamma_j \gamma_k = \gamma_0 \gamma_i \gamma_j \gamma_k \\ & \gamma_0 \gamma_i \gamma_j \gamma_k = \frac{i}{4} [\gamma_0, \gamma_i \gamma_j \gamma_k] \xrightarrow{\text{commute once}} \\ & = \frac{i}{4} \left( \begin{array}{c} \gamma_0 \gamma_i \gamma_j \gamma_k - \gamma_i \gamma_0 \gamma_j \gamma_k \\ + \gamma_0 \gamma_j \gamma_i \gamma_k - \gamma_j \gamma_0 \gamma_i \gamma_k \end{array} \right) = 0. \end{aligned}$$

So that we always have

$$\gamma^0 \gamma^\mu \gamma^\nu = \gamma^\nu \gamma^\mu \gamma^0$$

(Step 3) Now we verify the Lorentz invariance of  $\mathcal{L}_{\text{Dirac}}$ , first, under coordinate transformation

$$\begin{aligned} & \psi'(\pi) = \wedge'_2 \psi(\lambda^{-1}\pi) \\ & \bar{\psi}'(\pi) = [\wedge'_2 \psi(\lambda^{-1}\pi)]^+ \gamma^0 = \psi^+(\lambda^{-1}\pi) \wedge'_2 \gamma^0 \\ & = \psi^+(\lambda^{-1}\pi) \left( 1 + \frac{i}{2} \omega_{\mu\nu} \omega^{\mu\nu} \right) \gamma^0 \\ & = \psi^+(\lambda^{-1}\pi) \gamma^0 \left( 1 + \frac{i}{2} \omega_{\mu\nu} \omega^{\mu\nu} \right) \\ & = \psi^+(\lambda^{-1}\pi) \gamma^0 \wedge'_2 \gamma^0 \\ & = \bar{\psi}(\lambda^{-1}\pi) \gamma^0 \wedge'_2 \gamma^0 \\ & = \bar{\psi}(\lambda^{-1}\pi) \wedge'_2 \gamma^0 \end{aligned}$$

Now, the  $\mathcal{L}_{\text{Dirac}}$  in new coordinate system is

$$\begin{aligned} & \mathcal{L}'_{\text{Dirac}}(\pi) = \bar{\psi}'(\pi) \left( i \gamma^\mu \partial_\mu - m \right) \psi'(\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \wedge'_2 \left( i \gamma^\mu (\lambda^{-1})^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \wedge'_2 \left( i \gamma^\mu \wedge'_2 (\lambda^{-1})^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \left( i \gamma_2^{-1} \gamma^\mu \wedge'_2 \gamma_2^{-1} \gamma^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \left( i \gamma_2^{-1} \gamma^\mu \wedge'_2 \gamma_2^{-1} \gamma^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \left( i \gamma^\mu \wedge'_2 \gamma^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \left( i \gamma^\mu \wedge'_2 \gamma^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \left( i \gamma^\mu \wedge'_2 \gamma^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \end{aligned}$$

it has been proved in step 1.

$$\begin{aligned} & \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma = \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma \\ & \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma = \frac{i}{4} [\gamma^\mu, \gamma^\nu \gamma^\rho \gamma^\sigma] \xrightarrow{\text{commute twice}} \\ & = \frac{i}{4} \left( \begin{array}{c} \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma - \gamma^\nu \gamma^\mu \gamma^\rho \gamma^\sigma \\ + \gamma^\mu \gamma^\rho \gamma^\nu \gamma^\sigma - \gamma^\rho \gamma^\mu \gamma^\nu \gamma^\sigma \end{array} \right) \\ & = \bar{\psi}(\lambda^{-1}\pi) \left( i \gamma^\mu \wedge'_2 \gamma^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) \\ & = \bar{\psi}(\lambda^{-1}\pi) \left( i \gamma^\mu \wedge'_2 \gamma^\sigma \partial_\sigma - m \right) \psi(\lambda^{-1}\pi) = \mathcal{L}_{\text{Dirac}}(\lambda^{-1}\pi) \end{aligned}$$

So  $\mathcal{L}_{\text{Dirac}}$  is invariant under Lorentz transformation.

The final equations are always true, so that  $\psi_{\alpha}$  must be solution of Dirac eq.

### Problem 3. Peskin and Schroeder's book , Page 48-49

- 3.1. Verify (3.59) as a positive energy solution of Dirac eq. (3.31) in momentum space  
in momentum space
- 3.2 Verify (3.62) as a negative solution of (3.31) in momentum space.
- 3.3 Verify (3.63), (3.64) (3.66) and (3.67).

Solution :

$$\text{3.1. Dirac equation (3.31) is } (\bar{i} \gamma^\mu \partial_\mu - m) \psi(x) = 0$$

$$\text{It's positive energy solution } \psi(x) = u(p)e^{-ip \cdot x}, u(p) = \begin{pmatrix} \sqrt{p \cdot \sigma} & s \\ p \cdot \bar{\sigma} & s \end{pmatrix}$$

(step 1) We first reformulate Dirac equation into Weyl equations.

$$(\bar{i} \gamma^\mu \partial_\mu - m) \psi(x) = \left( \bar{i} \begin{pmatrix} \sigma^\mu & 0 \\ 0 & \bar{\sigma}^\mu \end{pmatrix} \partial_\mu - m \right) \psi(x)$$

$$= \left[ \bar{i} \begin{pmatrix} 0 & \sigma^\mu \partial_\mu \\ \bar{\sigma}^\mu \partial_\mu & 0 \end{pmatrix} - \begin{pmatrix} m \gamma_{2+2} & \\ & m \gamma_{2+2} \end{pmatrix} \right] \psi(x)$$

$$= \begin{pmatrix} -m \gamma_{2+2} & i \sigma^\mu \partial_\mu \\ i \bar{\sigma}^\mu \partial_\mu & -m \gamma_{2+2} \end{pmatrix} \psi(x) = 0$$

This is just Weyl equations.

(step2) We verify  $\psi(x) = \begin{pmatrix} \sqrt{p \cdot \sigma} & s \\ p \cdot \bar{\sigma} & s \end{pmatrix} e^{-ip \cdot x}$  is positive energy ( $p_0 > 0$ ) solution of Weyl equations.

$$\begin{pmatrix} -m & i \sigma^\mu \partial_\mu \\ i \bar{\sigma}^\mu \partial_\mu & -m \end{pmatrix} \begin{pmatrix} \sqrt{p \cdot \sigma} & e^{-ip \cdot x} \\ p \cdot \bar{\sigma} & e^{-ip \cdot x} \end{pmatrix} = 0$$

$$\Leftrightarrow \begin{cases} -m \int \sqrt{p \cdot \sigma} \gamma^\mu \partial_\mu \sqrt{p \cdot \sigma} e^{-ip \cdot x} + i \sigma^\mu \partial_\mu \int \sqrt{p \cdot \sigma} s e^{-ip \cdot x} = 0 \\ i \bar{\sigma}^\mu \partial_\mu \int p \cdot \bar{\sigma} s e^{-ip \cdot x} - m \int p \cdot \bar{\sigma} \gamma^\mu \partial_\mu s e^{-ip \cdot x} = 0 \end{cases}$$

$$\begin{aligned} &\stackrel{\text{①} \times \sqrt{p \cdot \sigma}}{\Leftrightarrow} \begin{cases} -m \int \sqrt{p \cdot \sigma} \gamma^\mu \partial_\mu \sqrt{p \cdot \sigma} s e^{-ip \cdot x} = 0 \\ i \bar{\sigma}^\mu \partial_\mu \int p \cdot \bar{\sigma} s e^{-ip \cdot x} = 0 \end{cases} \quad \begin{aligned} &\stackrel{\text{②} \times \sqrt{p \cdot \bar{\sigma}}}{\Leftrightarrow} \begin{cases} -m \int p \cdot \bar{\sigma} \gamma^\mu \partial_\mu m \eta = 0 \\ -\bar{\sigma}^\mu \partial_\mu m \eta + m p \cdot \bar{\sigma} \eta = 0 \end{cases} \end{aligned} \\ &\stackrel{\text{①} \times \sqrt{p \cdot \sigma}}{\Leftrightarrow} \begin{cases} -m p \cdot \sigma \xi + \sigma^\mu \partial_\mu \sqrt{p \cdot \sigma} p \cdot \bar{\sigma} \xi = 0 \\ \bar{\sigma}^\mu \partial_\mu \sqrt{p \cdot \sigma} p \cdot \bar{\sigma} \xi = 0 \end{cases} \quad \begin{aligned} &\stackrel{\text{②} \times \sqrt{p \cdot \bar{\sigma}}}{\Leftrightarrow} \begin{cases} -m p \cdot \bar{\sigma} \eta + \sigma^\mu \partial_\mu m \eta = 0 \\ -\bar{\sigma}^\mu \partial_\mu m \eta + m p \cdot \bar{\sigma} \eta = 0 \end{cases} \end{aligned} \end{aligned}$$

$$\begin{aligned} &\Leftrightarrow \begin{cases} -m p \cdot \sigma \xi + \sigma^\mu \partial_\mu m \xi = 0 \\ \bar{\sigma}^\mu \partial_\mu m \xi = 0 \end{cases} \quad \begin{aligned} &\stackrel{\text{①} \times \sqrt{p \cdot \sigma} - \text{②} \times \sqrt{p \cdot \bar{\sigma}}}{=} \begin{cases} (p \cdot \sigma)(p \cdot \bar{\sigma}) = (E \gamma_{11} - \vec{p} \cdot \vec{\sigma})(E \gamma_{11} + \vec{p} \cdot \vec{\sigma}) = m^2 \\ = E^2 - (\vec{p} \cdot \vec{\sigma})^2 = E^2 - (p \cdot \bar{\sigma})^2 = m^2 \end{cases} \\ &\text{also } (p \cdot \bar{\sigma})(p \cdot \sigma) = -2m \delta^{rs}, \quad 2^{\text{st}} \int p_\mu \delta^{rs} = 2 \delta^{rs}, \quad \text{we will} \end{aligned} \end{aligned}$$

### (step3) Verify (3.62)

$$\begin{aligned} \bar{u}^r(p) u^s(p) &= \begin{pmatrix} \xi_r^\dagger \sqrt{p \cdot \sigma} & s \\ p \cdot \bar{\sigma} & s \end{pmatrix} \begin{pmatrix} \xi_s^\dagger \sqrt{p \cdot \sigma} & s \\ p \cdot \bar{\sigma} & s \end{pmatrix} = \begin{pmatrix} \xi_r^\dagger \sqrt{p \cdot \sigma} \xi_s & s \\ p \cdot \bar{\sigma} \xi_s & s \end{pmatrix} \\ &\stackrel{\text{①} \times \bar{\sigma}^\mu \partial_\mu + \text{②} \times \sigma^\mu \partial_\mu}{=} \begin{pmatrix} \xi_r^\dagger \sqrt{p \cdot \sigma} \xi_s & s \\ p \cdot \bar{\sigma} \xi_s & s \end{pmatrix} \\ &= \xi_r^\dagger \sqrt{p \cdot \sigma} \xi_s + \xi_r^\dagger \sqrt{p \cdot \bar{\sigma}} \sqrt{p \cdot \sigma} \xi_s \\ &= \xi_r^\dagger \sqrt{p \cdot \sigma} \xi_s \end{aligned}$$

$$\begin{aligned} &\stackrel{(p \cdot \sigma)(p \cdot \bar{\sigma}) = m^2}{=} \begin{pmatrix} \xi_r^\dagger \sqrt{p \cdot \sigma} & s \\ p \cdot \bar{\sigma} & s \end{pmatrix} \\ &\stackrel{\text{① has proved in last step}}{=} 2m \xi_r^\dagger \xi_s = 2m \delta^{rs} \end{aligned}$$

$$\begin{aligned} u^r(p) u^s(p) &= \begin{pmatrix} \xi_r^\dagger \sqrt{p \cdot \sigma} & s \\ p \cdot \bar{\sigma} & s \end{pmatrix} \begin{pmatrix} \xi_s^\dagger \sqrt{p \cdot \sigma} & s \\ p \cdot \bar{\sigma} & s \end{pmatrix} \\ &= \xi_r^\dagger \xi_s (\sqrt{p \cdot \sigma})^2 \xi_s + \xi_r^\dagger (\sqrt{p \cdot \sigma})^2 \xi_s \\ &= \xi_r^\dagger (\xi_{\frac{11}{1}} - \vec{p} \cdot \vec{\sigma}) \xi_s + \xi_r^\dagger (\vec{p} \cdot \vec{\sigma} + \vec{p} \cdot \vec{\sigma}) \xi_s \\ &= 2E \vec{p} \cdot \xi_s \xi_s = 2E \vec{p} \delta^{rs} \end{aligned}$$

3.2 Verify (3.62)  $\psi(x) = v(p) e^{+ip \cdot x} = \begin{pmatrix} \sqrt{p \cdot \sigma} & ? \\ -\bar{p} \cdot \bar{\sigma} & ? \end{pmatrix} e^{+ip \cdot x}$  as negative energy solution of Dirac equation. By directly calculating we obtain

$$\begin{pmatrix} i \gamma^\mu \partial_\mu - m & \phi(x) \\ i \bar{\sigma}^\mu \partial_\mu - m & \phi(x) \end{pmatrix} \begin{pmatrix} \sqrt{p \cdot \sigma} \gamma \cdot e^{+ip \cdot x} \\ -\bar{p} \cdot \bar{\sigma} \gamma \cdot e^{+ip \cdot x} \end{pmatrix} = 0$$

$$\Rightarrow \begin{cases} -m \sqrt{p \cdot \sigma} \gamma \cdot e^{+ip \cdot x} + i \sigma^\mu \partial_\mu \sqrt{p \cdot \sigma} \gamma \cdot e^{+ip \cdot x} = 0 \\ i \bar{\sigma}^\mu \partial_\mu \sqrt{p \cdot \sigma} \gamma \cdot e^{+ip \cdot x} + m \int p \cdot \bar{\sigma} \gamma \cdot e^{+ip \cdot x} = 0 \end{cases}$$

$$3.2 \text{ Verify (3.62)} \quad \psi(x) = v(p) e^{+ip \cdot x} = \begin{pmatrix} \sqrt{p \cdot \sigma} & ? \\ -\bar{p} \cdot \bar{\sigma} & ? \end{pmatrix} e^{+ip \cdot x}$$

$$\begin{aligned} &\stackrel{i \partial_\mu e^{+ip \cdot x}}{=} \begin{cases} -\bar{p}_\mu \partial_\mu e^{+ip \cdot x} & \\ = \bar{p}_\mu \partial_\mu e^{+ip \cdot x} & \end{cases} \quad \begin{cases} -m \int \sqrt{p \cdot \sigma} \gamma^\mu \partial_\mu \sqrt{p \cdot \sigma} \gamma \cdot e^{+ip \cdot x} = 0 \\ \bar{p}^\mu \partial_\mu \sqrt{p \cdot \sigma} \gamma \cdot e^{+ip \cdot x} + m \int p \cdot \bar{\sigma} \gamma \cdot e^{+ip \cdot x} = 0 \end{cases} \quad \text{①} \\ &\stackrel{-\bar{p}_\mu \partial_\mu \sqrt{p \cdot \sigma} \gamma \cdot e^{+ip \cdot x}}{=} \begin{cases} -m \int p \cdot \sigma \gamma^\mu \partial_\mu m \eta = 0 \\ -\bar{\sigma}^\mu \partial_\mu m \eta + m p \cdot \bar{\sigma} \eta = 0 \end{cases} \quad \text{②} \end{aligned}$$

$$\begin{aligned} &\stackrel{\text{①} \times \sqrt{p \cdot \sigma}}{\Leftrightarrow} \begin{cases} -m p \cdot \sigma \xi + \sigma^\mu \partial_\mu \sqrt{p \cdot \sigma} p \cdot \bar{\sigma} \xi = 0 \\ \bar{\sigma}^\mu \partial_\mu \sqrt{p \cdot \sigma} p \cdot \bar{\sigma} \xi = 0 \end{cases} \quad \begin{aligned} &\stackrel{\text{②} \times \sqrt{p \cdot \bar{\sigma}}}{\Leftrightarrow} \begin{cases} -m p \cdot \bar{\sigma} \eta + \sigma^\mu \partial_\mu m \eta = 0 \\ -\bar{\sigma}^\mu \partial_\mu m \eta + m p \cdot \bar{\sigma} \eta = 0 \end{cases} \end{aligned} \\ &\stackrel{\text{①} \times \sqrt{p \cdot \sigma} - \text{②} \times \sqrt{p \cdot \bar{\sigma}}}{\Leftrightarrow} \begin{cases} (p \cdot \sigma)(p \cdot \bar{\sigma}) = (E \gamma_{11} - \vec{p} \cdot \vec{\sigma})(E \gamma_{11} + \vec{p} \cdot \vec{\sigma}) = m^2 \\ = E^2 - (\vec{p} \cdot \vec{\sigma})^2 = E^2 - (p \cdot \bar{\sigma})^2 = m^2 \end{cases} \end{aligned}$$

which has always true.

And because  $v(p)$  and  $v(p)$  just differ by a minus sign, which will (not) appear in  $\bar{u}^r(p)v(p)$  and  $u^r(p)v(p)$ . So the normalization is automatically satisfied:  $\bar{u}^r(p)v(p) = -2m \delta^{rs}$ ,  $2^{\text{st}} \int p_\mu \delta^{rs} = 2 \delta^{rs}$ .





$$[L_{\mu\nu}, L_{\rho\sigma}] = i [\eta_{\mu\rho} [L_{\nu\sigma} - (L_{\mu\nu})] - (L_{\nu\sigma}) + (L_{\mu\nu})]$$

\* Final exam schedule

Time : Jan 17, 2017, 2:30 PM - 4:30 PM (Tuesday, afternoon)

Place : NOT fixed yet  
Big classroom (or in old building up school of Physics)

May be elsewhere.  
Final Exam : Easy ↗ Basic Concepts  
↳ Basic Calculation

Open test - open to classmates & homeworks.

\* Schedule on last class, Jan 3, 2017

Student's talk on their research report! PDF file or PPT, 20 minutes or 30 minutes.

Solution to the second homework

Problem 1. Lorentz group & Lorentz algebra

$$[\eta^{\mu\nu}, \eta_{\rho\sigma}] = i [\eta^{\rho\mu}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta_{\rho\sigma}) + (\eta^{\rho\sigma})] \quad (1.1)$$

Total angular momentum (generator) -  $\eta^{\mu\nu} = -\eta_{\mu\nu}$

Sum of orbital and spin

$$D(\omega) = \exp(-i \frac{\omega}{2} \eta^{\mu\nu})$$

$$\omega^{\mu\nu} = -\omega_{\mu\nu}$$

$$\eta^{\mu\nu} = L^{\mu\nu} + S^{\mu\nu}$$

Orbital spin

$$\begin{aligned} \omega^{\mu\nu} &= \omega^{\mu\nu} \text{ Angular momentum} \\ &= i(\eta^{\mu\nu} - \eta_{\nu\mu}) \quad \text{if } \eta^{\mu\nu} = \eta_{\nu\mu} \\ &\quad \text{if } \eta^{\mu\nu} \neq \eta_{\nu\mu} \end{aligned}$$

Spin =  $\frac{1}{2}$  angular momentum =  $\eta^{\mu\nu}$  + ...

Try to prove  $L^{\mu\nu}$

$$S^{\mu\nu} (S^{\rho\sigma})^{-1}, S^{\mu\nu} - \frac{1}{2}$$

Satisfying Lorentz algebra (1.1)

$$\begin{aligned} \omega^{\mu\nu} &= \frac{1}{2}\eta^{\mu\nu} = \frac{i}{4}[\eta^{\mu\nu}, S^{\rho\sigma}] \\ &= \{ \eta^{\mu\nu}, \eta^{\rho\sigma} \} = 2\eta^{\mu\nu} \end{aligned}$$

$$\begin{aligned} L^{\mu\nu} &= i(\eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma})) \\ \Leftrightarrow [L^{\mu\nu}, L^{\rho\sigma}] &= i^2 [x_{\mu\nu}x_{\rho\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma})] \\ &= i \left\{ i \left[ \eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma}) \right] \right. \\ &\quad \left. - i \left[ \eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma}) \right] \right\} \\ &= i \left[ \eta^{\mu\rho}\eta_{\nu\sigma} - i\eta_{\sigma\nu}\eta^{\rho\mu} \right. \\ &\quad \left. - i\eta^{\mu\rho}\eta_{\nu\sigma} + i\eta_{\nu\mu}\eta^{\rho\sigma} \right] \end{aligned}$$

$$\begin{aligned} &= i \left[ (\eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma})) \right. \\ &\quad \left. - (\eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma})) \right] \\ &= 0 \end{aligned}$$

$$\begin{aligned} &= i \left( (\eta^{\mu\nu}\eta_{\rho\sigma}) - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma}) \right) \\ &\quad - (\eta^{\mu\nu}\eta_{\rho\sigma}) + (\eta^{\mu\nu}) \\ &= i \left( (\eta^{\mu\nu}\eta_{\rho\sigma}) - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma}) \right) \\ &\quad + (\eta^{\mu\nu}\eta_{\rho\sigma}) - (\eta^{\mu\nu}) \\ &= 0 \end{aligned}$$

$\Rightarrow$   $\omega^{\mu\nu} = 0$ :

Index change

$$\begin{aligned} [\bar{J}_{\mu\nu}, \bar{J}_{\rho\sigma}] &= \left[ (\omega_{\mu}, \omega_{\sigma}) - (\omega_{\nu}, \omega_{\rho}) - (\omega_{\mu}, \omega_{\rho}) + (\omega_{\nu}, \omega_{\sigma}) \right] \\ &= \left[ (\bar{J}_{\mu\nu})^{\beta} (\bar{J}_{\rho\sigma})^{\alpha} - (\bar{J}_{\mu\nu})^{\alpha} (\bar{J}_{\rho\sigma})^{\beta} \right] \\ &= \bar{J}_{\mu\nu} \cdot \bar{J}_{\rho\sigma} \quad \text{in General relativity} \\ &= \int (\bar{J}_{\mu\nu})^{\beta} (\bar{J}_{\rho\sigma})^{\alpha} = i (\eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma})) \end{aligned}$$

$$\begin{aligned} &= i^2 (\eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma})) \\ &\quad + (\eta^{\mu\rho}\eta_{\nu\sigma} - (\eta^{\mu\nu}) - (\eta^{\rho\sigma}) + (\eta^{\mu\sigma})) \\ &= i^2 \eta^{\mu\rho}\eta_{\nu\sigma} - i^2 \eta^{\mu\nu} - i^2 \eta^{\rho\sigma} + i^2 \eta^{\mu\sigma} \\ &= i^2 \eta_{\nu\mu}\eta^{\rho\sigma} = i^2 \eta_{\nu\mu}\eta^{\rho\sigma} = i^2 \eta_{\nu\mu}\eta^{\rho\sigma} = i^2 \eta_{\nu\mu}\eta^{\rho\sigma} \end{aligned}$$

$$= i^2 \eta_{\nu\mu}\eta^{\rho\sigma} = i^2 \eta_{\nu\mu}\eta^{\rho\sigma}$$

$$so \quad (J_{\mu\nu})^{\rho} (J_{\rho\sigma})^{\tau} = i \left( (\nu_{\rho}, \mu, \sigma) - (\mu, \nu, \sigma) + (\rho, \nu, \sigma) - (\rho, \sigma, \nu) \right)$$

$$now \quad [J_{\mu\nu}, J_{\rho\sigma}]^{\alpha} = i \left[ (\nu_{\rho}, \mu, \sigma) - (\mu, \nu, \sigma) - (\rho, \nu, \sigma) + (\rho, \sigma, \nu) \right] - (\mu, \nu, \sigma, \rho)$$

$$= i \left[ (\nu_{\rho}, \mu, \sigma) - (\mu, \nu, \sigma) - (\nu_{\sigma}, \mu, \rho) + (\mu, \nu, \rho) \right]$$

$$\stackrel{(1)}{=} - (\nu_{\rho}, \nu, \sigma) + (\rho, \mu, \sigma, \nu) + (\nu_{\sigma}, \rho, \mu) + (\rho, \nu, \sigma, \rho) \\ \stackrel{(2)}{=} - (\mu, \nu, \sigma, \nu) + (\rho, \mu, \sigma, \nu) + (\nu_{\sigma}, \rho, \mu) + (\rho, \nu, \sigma, \rho) \\ \stackrel{(3)}{=} - (\mu, \nu, \sigma, \nu) + (\mu, \nu, \sigma, \rho) + (\mu, \nu, \sigma, \rho)$$

$$+ (\mu, \nu, \sigma, \rho) - (\mu, \nu, \sigma, \rho) - (\mu, \nu, \sigma, \rho)$$

$$\nu_{\mu\nu} = \frac{i}{4} \left[ \nu_{\rho\nu} - \nu_{\mu\rho} \right] = \frac{1}{2} \nu_{\mu\nu}$$

$$= \frac{i}{4} \left( \nu_{\rho\nu} - \nu_{\mu\rho} \right), \quad \{ \nu_{\mu\nu}, \nu_{\rho\sigma} \} = 2\nu_{\mu\nu} - three approach \\ is totally equivalent.$$

$$Approach 1. \quad [J_{\mu\nu}, J_{\rho\sigma}]^{\alpha} = \left( \frac{i}{4} \right)^2 \left[ \nu_{\rho\nu} \nu_{\sigma\mu} - \nu_{\rho\mu} \nu_{\sigma\nu} \right]$$

$$= \left( \frac{i}{4} \right)^2 \left( \nu_{\rho\nu} \nu_{\sigma\mu} - \nu_{\rho\mu} \nu_{\sigma\nu} - (\mu, \nu, \sigma, \rho) + (\mu, \nu, \sigma, \rho) \right)$$

$$where \quad \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \\ \stackrel{\leftrightarrow}{=} \nu_{\mu} \left( \nu_{\rho\nu} - \nu_{\mu\rho} \right) \nu_{\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma}$$

$$= 2\nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \\ = 2\nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} + \nu_{\mu} \nu_{\rho\sigma} \nu_{\nu\mu} - \nu_{\mu\rho} \nu_{\sigma\nu}$$

$$+ 2\nu_{\mu\nu} \nu_{\rho\sigma} - 2\nu_{\mu\rho} \nu_{\nu\sigma} + 2\nu_{\mu} \nu_{\rho\sigma} \nu_{\nu\mu} - 2\nu_{\mu\rho} \nu_{\sigma\nu}$$

$$+ 2\nu_{\mu\nu} \nu_{\rho\sigma} - 2\nu_{\mu\rho} \nu_{\nu\sigma} - 2\nu_{\mu\rho} \nu_{\sigma\nu} + 2\nu_{\mu\nu} \nu_{\rho\sigma} + 2\nu_{\mu} \nu_{\rho\sigma} \nu_{\nu\mu} - 2\nu_{\mu\rho} \nu_{\sigma\nu}$$

How to go on?

$$Approach 2. \quad \nu_{\mu\nu} = \frac{i}{4} \left( \nu_{\rho\nu} - \nu_{\mu\rho} \right) \leftarrow anti-symmetric, \nu_{\mu\nu}$$

$$= \begin{cases} 0 & \mu = \nu \\ \frac{i}{2} \nu_{\mu\nu} & \mu \neq \nu \end{cases} \quad \nu_{\mu\nu} = -\nu_{\nu\mu}$$

$$[S_{\mu\nu}, S_{\rho\sigma}] = \frac{i}{4} \left( \nu_{\rho\nu} - \nu_{\mu\rho} \right) \left( \nu_{\sigma\mu} - \nu_{\nu\sigma} \right) - \left( \nu_{\rho\nu} \nu_{\sigma\mu} - \nu_{\mu\rho} \nu_{\nu\sigma} \right)$$

$$= \left( \frac{i}{2} \right)^2 \left[ \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \right] = \left( \frac{i}{2} \right)^2 \left[ 2\nu_{\mu\nu} \nu_{\rho\sigma} - 2\nu_{\mu\rho} \nu_{\nu\sigma} \right]$$

How to go on?

$$Approach 3. \quad [S_{\mu\nu}, S_{\rho\sigma}] = \frac{i}{4} \left[ \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \right] - C(\mu, \nu, \sigma, \rho)$$

$$= \frac{i}{4} \left( \nu_{\mu\nu} \nu_{\rho\sigma} + \nu_{\mu\rho} \nu_{\nu\sigma} + \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \right) - C(\mu, \nu, \sigma, \rho)$$

$$[S_{\mu\nu}, S_{\rho\sigma}] = \frac{i}{4} \left( \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \right) \leftarrow (\mu, \nu, \sigma, \rho)$$

$$= \frac{i}{4} \left( \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \right) - (\mu, \nu, \sigma, \rho)$$

$$= \frac{i}{4} \left( \nu_{\mu\nu} \nu_{\rho\sigma} + \nu_{\mu\rho} \nu_{\nu\sigma} - \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \right) - (\mu, \nu, \sigma, \rho)$$

$$= \frac{i}{4} \left( -\nu_{\mu} \nu_{\rho\nu} + 2\nu_{\mu\rho} \nu_{\nu\nu} - \nu_{\mu} \nu_{\rho\nu} \right) - (\mu, \nu, \sigma, \rho)$$

$$= \frac{i}{4} \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right) - (\mu, \nu, \sigma, \rho)$$

$$= \frac{i}{4} \times \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right) = i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right)$$

$$+ i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right) = i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right)$$

$$+ i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right) = i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right)$$

$$+ i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right) = i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right)$$

$$+ i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right) = i \left( \nu_{\mu} \nu_{\rho\nu} - 2\nu_{\mu\rho} \nu_{\nu\nu} \right)$$

How to go on.

Final relationship between  $\nu_{\mu\nu}$  and  $\nu_{\mu\nu}$

$$\nu_{\mu\nu} = \frac{i}{4} \left( \nu_{\mu} \nu_{\nu} - \nu_{\nu} \nu_{\mu} \right)$$

$$= \frac{i}{4} \left( \nu_{\mu} \nu_{\nu} - (2\nu_{\mu\nu} - \nu_{\mu} \nu_{\nu}) \right) = \frac{i}{2} \left( \nu_{\mu} \nu_{\nu} - \nu_{\mu\nu} \right)$$

$$\Rightarrow \nu_{\mu\nu} = \nu_{\mu\nu} = 2\nu_{\mu\nu}$$

$$and \quad \nu_{\mu\nu} \nu_{\rho\sigma} = \left( \nu_{\mu} \nu_{\rho} - 2\nu_{\mu\rho} \nu_{\nu\sigma} \right) = (0, \nu, \sigma)$$

$$+ 2\nu_{\mu\nu} \nu_{\rho\sigma} \left( \nu_{\mu} \nu_{\rho} - 2\nu_{\mu\rho} \nu_{\nu\sigma} \right) - i \nu_{\mu\nu} \nu_{\rho\sigma} - i \nu_{\mu\nu} \nu_{\rho\sigma}$$

$$= \frac{i}{2} \left( \nu_{\mu} \nu_{\rho} - \nu_{\mu\rho} \nu_{\nu\sigma} + (\mu, \nu, \sigma) + (\mu, \nu, \sigma) \right) - (\mu, \nu, \sigma)$$

$$now \quad [S_{\mu\nu}, S_{\rho\sigma}] = i \left( \nu_{\mu} \nu_{\rho} \nu_{\nu\sigma} - (\mu, \nu, \sigma) - (\mu, \nu, \sigma) \right) + \left( \mu, \nu, \sigma \right)$$

$$[S_{\mu\nu}, S_{\rho\sigma}] = \frac{i}{4} \left( \nu_{\rho\nu} - \nu_{\mu\rho} \right) \left( \nu_{\sigma\mu} - \nu_{\nu\sigma} \right) - \left( \mu, \nu, \sigma \right)$$

$$= \left( \frac{i}{2} \right)^2 \left[ \nu_{\mu\nu} \nu_{\rho\sigma} - \nu_{\mu\rho} \nu_{\nu\sigma} \right] = \left( \frac{i}{2} \right)^2 \left[ 2\nu_{\mu\nu} \nu_{\rho\sigma} - 2\nu_{\mu\rho} \nu_{\nu\sigma} \right]$$

$$+ 2\nu_{\mu\nu} \nu_{\rho\sigma} - 2\nu_{\mu\rho} \nu_{\nu\sigma} - 2\nu_{\mu\nu} \nu_{\rho\sigma} + 2\nu_{\mu\rho} \nu_{\nu\sigma}$$

How to go on?

Note: Result  $[S_{\mu\nu}, \gamma_\rho] = i(\gamma_{\rho\nu}\gamma_\mu - \gamma_{\rho\mu}\gamma_\nu)$  is important relation

$$\begin{aligned} &= i[\gamma_{\rho\nu}\gamma_\mu\gamma_\rho - \gamma_{\rho\mu}\gamma_\nu] \\ &= i[\gamma_{\rho\nu}\gamma_\mu\gamma_\rho - \gamma_{\rho\mu}\gamma_\nu] \gamma_\sigma \\ &= -\gamma_\nu [\gamma_\mu\gamma_\rho - \gamma_{\nu\rho}\gamma_\mu] \gamma_\sigma \\ &= -(\gamma_{\mu\nu})_\rho^\sigma \gamma_\sigma \end{aligned}$$

$$[\gamma_\rho, S_{\mu\nu}] = (\gamma_{\mu\nu})_\rho^\sigma \gamma_\sigma$$

very important! we can construct spin-1 from spin-1/2 algebra.

For vector representation

$$\gamma^\alpha = \gamma^\alpha_\beta \gamma^\beta, \quad \gamma^\alpha_\beta = \exp(-i\omega_\mu \gamma^\mu \gamma^\alpha)^\beta_\mu$$

$$= (1 - i\omega_2 \gamma_{\mu\nu} (\gamma^{\mu\nu})^\alpha_\beta + O(\omega^2))$$

$$\gamma^{\mu\nu} = -\bar{\gamma}^{\nu\mu}, \quad \omega_{\mu\nu} = -\omega_{\nu\mu}$$

$$(\gamma^{\mu\nu})^\alpha_\beta = i(\gamma^{\mu\alpha}\gamma^\nu_\beta - \gamma^{\nu\alpha}\gamma^\mu_\beta)$$

Photon's spin is very strange!

such that:

$$-\gamma_2^\nu \omega_{\mu\nu} \gamma^\mu_\beta$$

$$= -\gamma_2^\nu \omega_{\mu\nu} (\gamma^{\mu\alpha}\gamma^\nu_\beta - \gamma^{\nu\alpha}\gamma^\mu_\beta)$$

$$= \omega_{\mu\nu} \gamma^\mu\gamma^\nu_\beta = \omega_\beta$$

very easy, first infinitesimal expansion

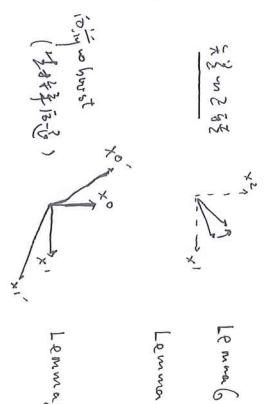
$$\gamma^\alpha_\beta = \delta^\alpha_\beta + \omega^\alpha_\beta + O(\omega^2)$$

$$\text{so: } \omega_{12} = 0, \quad \omega_{21} = -i, \quad \omega_1^1 = -i, \quad \omega_2^2 = i$$

$$(\gamma^\alpha_\beta) = \left( \begin{array}{cc} 1 & -i \\ 0 & 0 \end{array} \right) + \left( \begin{array}{cc} 0 & -i \\ 0 & 0 \end{array} \right)$$

$$\text{first: } \omega_{01} = i, \quad \omega_{00} = -i, \quad \omega_0^1 = i, \quad \omega_0^2 = i$$

$$(\gamma^\alpha_\beta) = \left( \begin{array}{cc} 1 & i \\ 0 & 0 \end{array} \right) + \left( \begin{array}{cc} 0 & i \\ 0 & 0 \end{array} \right)$$



Again, talk about  $\gamma^{\mu\nu}$

$$\begin{aligned} \gamma^\alpha_\beta &= \delta^\alpha_\beta - \gamma_2^\nu \omega_{\mu\nu} \gamma^\mu_\beta + O(\omega^2) \\ \gamma^\alpha_\beta &\stackrel{\text{rotate}}{=} \delta^\alpha_\beta - i\omega_0 \gamma^\mu \gamma^\nu \gamma^\alpha_\beta + O(\omega^2), \\ &= \left( \begin{array}{cc} 1 & -i \\ 0 & 0 \end{array} \right) + \left( \begin{array}{cc} 0 & -i \\ 0 & 0 \end{array} \right) \\ &= \left( \begin{array}{cc} 1 & 0 \\ 0 & 1 \end{array} \right) + \left( \begin{array}{cc} 0 & -i \\ 0 & 0 \end{array} \right) \\ &= \left\{ \begin{array}{l} d=1 \\ \mu=2 \\ \nu=1 \end{array} \right\} \rightarrow \text{anti-sym.} \end{aligned}$$

$$\gamma^{12} = \left( \begin{array}{cc} 0 & -1 \\ 1 & 0 \end{array} \right).$$

Lemma:  $S_{\mu\nu} = \frac{i}{4} [\epsilon_{\mu\nu}, \gamma^\nu]$ , calculate  $S_{\mu\nu} \cdot \gamma_3 = 2\gamma_{\mu\nu}$  (most complete)

(key point: Lorentz invariance of)

Lemma 1:  $\gamma^\mu \omega_\mu, \quad S_{\mu\nu} = \gamma_2^\nu \omega_{\mu\nu}$  ( $\mu = \nu, \quad S_{\mu\nu} = 0$ )

Lemma 2:  $[\gamma_\rho, S_{\mu\nu}] = (\gamma_{\mu\nu})_\rho^\sigma \gamma_\sigma$  (Lorentz Algebra)

Lemma 3:  $[\gamma_\rho, S_{\mu\nu}] = i(\gamma_{\rho\mu} S_{\nu\rho} - (\mu\nu\rho) - (\rho\mu\nu)) + (\rho\nu\mu)$

Lemma 4:  $S_{\mu\nu} = \frac{i}{2} \gamma^{\mu\nu}$

$$(S_{\mu\nu})^\dagger = -\gamma_2^\nu (S_{\mu\nu})^\dagger = -\gamma_2^\nu (\delta^{\mu\nu} (\gamma^{\mu\nu}))^\dagger = \gamma_2^\nu (\delta^{\mu\nu})^\dagger = -S_{\mu\nu}$$

$\gamma^\nu$  is not Hermitian

Lemma 5:  $\{S_{\mu\nu}, \gamma_\rho\} = 0$

$$S_{\mu\nu} \cdot \gamma^\nu = -i \gamma^\nu S_{\mu\nu} \quad S_{\mu\nu} \cdot \gamma^\nu = -S_{\mu\nu}$$

$$(S_{\mu\nu})^\dagger \gamma^\nu = \gamma^\nu S_{\mu\nu} \quad \text{(why?)}.$$

Problem 2.3 Direct Representation

$$\text{Lemma 6: } \gamma^\alpha = \left( \begin{array}{cc} 1 & 0 \\ 0 & 1 \end{array} \right) + i \left( \begin{array}{cc} 0 & 1 \\ -1 & 0 \end{array} \right) \quad S_{\mu\nu} = \frac{i}{2} \left( \begin{array}{cc} 1 & 0 \\ 0 & -1 \end{array} \right) \left( \begin{array}{cc} \gamma_1 & \gamma_2 \\ -\gamma_2 & \gamma_1 \end{array} \right) = -\gamma_2 \left( \begin{array}{cc} \gamma_1 & 0 \\ 0 & -\gamma_1 \end{array} \right)$$

Lemma 7:  $S_{\mu\nu} = \frac{i}{2} \gamma_\mu \gamma_\nu$

$$(\gamma_\mu)^+ = -\gamma_2^\nu (\gamma_\mu^\nu)^+ = -\gamma_2^\nu (-i)(-i) = -\gamma_2^\nu \gamma_2^\nu \Rightarrow (\gamma_\mu)^+ = S_{\mu\nu}$$

Lemma 8:  $[S_{\mu\nu}, \gamma_\rho] = 0$

$$\text{since } \delta^{\mu\nu} \delta_{\rho\sigma} = \delta^{\mu\rho} \delta_{\nu\sigma} \quad (S_{\mu\nu})^\dagger \gamma_\rho = \gamma_\rho (S_{\mu\nu})^\dagger$$

we have  $(S_{\mu\nu})^\dagger \gamma_\rho = \gamma_\rho S_{\mu\nu}$

$$S_{\mu\nu} = \gamma_2^\nu \delta_{\mu\nu} = \gamma_2^\nu \left( \begin{array}{cc} 0 & \gamma_1 \\ \gamma_1 & 0 \end{array} \right) = \gamma_1^{\mu\nu} \left( \begin{array}{cc} 0 & \gamma_1 \\ \gamma_1 & 0 \end{array} \right) = \gamma_1^{\mu\nu} \sum_k$$

Lemma 10:

$$\text{Spin-1/2 representation}$$

$$D^{\mu} = \exp(-i\gamma_2 \omega_{\mu} S^{\mu})$$

(key point to talk about Dirac invariance)

$$(D^{\mu})^{-1} = \exp(-i\gamma_2 \omega_{\mu} S^{\mu})$$

$$\text{but } (D^{\mu})^{-1} = \exp\left(-i\left[\omega_i(S^{\mu})^i + \omega_j(S^{\mu})^j\right]\right)$$

however

$$(D^{\mu})^{-1} = \exp\left(-i\left(\omega_{\mu} S^{\mu} + \omega_j S^j\right)\right)$$

$$(D^{\mu})^{-1} = \exp\left(-i\left(\omega_{\mu} S^{\mu} + \omega_j S^j\right) + \dots\right)$$

$$\begin{aligned} &= \left[ 1 + i\left(\omega_{\mu} S^{\mu} + \omega_j S^j\right) + \dots \right]^{1/2} \\ &= \left[ 1 + i\left(\omega_{\mu} S^{\mu} + \omega_j S^j\right) \right]^{1/2} \end{aligned}$$

Dirac eq. invariance under Lorentz transform:

$$\begin{aligned} \psi(x) &\rightarrow D^{\mu} \psi(x) \\ \bar{\psi}(x) &\rightarrow \bar{\psi}(x) (D^{\mu})^{-1} \end{aligned}$$

$D^{\mu}$   $\Rightarrow$  NOT unitary transformation

$$\bar{\psi}(x) = \psi^{\dagger}(x) \rightarrow \psi^{\dagger}(x) (D^{\mu})^{-1} = \bar{\psi}(x) (D^{\mu})^{-1}$$

$\Rightarrow$  that is why we use  $\bar{\psi}$

$$\bar{\psi}(x) \bar{\psi}(x) \leftrightarrow \bar{\psi}(x) (D^{\mu})^{-1} (D^{\nu})^{-1} \bar{\psi}(x) = \bar{\psi}(x) \bar{\psi}(x) \Rightarrow \text{Lagrangian Invariance}$$

Lemma 11.  $(D_{\mu})^{-1} \gamma^{\mu} D_{\mu} = \gamma^{\mu} \gamma^{\mu}$  both involve relation of spin-1 and spin-1/2

$$\begin{aligned} \gamma^{\mu} &= \exp\left(-i\gamma_2 \omega_{\mu} S^{\mu}\right) \\ \gamma^{\mu} &= \exp\left(-i\gamma_2 \omega_{\mu} S^{\mu}\right) \end{aligned}$$

$$\exp\left(+i\gamma_2 \omega_{\mu} S^{\mu}\right) \gamma^{\mu} \exp\left(-i\gamma_2 \omega_{\mu} S^{\mu}\right) = \exp\left(-i\gamma_2 \omega_{\mu} S^{\mu}\right) \gamma^{\mu}$$

$$A = i\gamma_2 \omega_{\mu} S^{\mu}$$

$$\beta = \gamma^{\mu}$$

$$\text{Use Baker-Hausdorff formula: } \exp A \beta \exp -A = \sum_{n=0}^{\infty} \frac{C_n}{n!} \beta^n$$

$$\begin{cases} A = i\gamma_2 \omega_{\mu} S^{\mu} \\ \beta = \gamma^{\mu} \end{cases}$$

$$\text{Use Baker-Hausdorff formula: } \exp A \beta \exp -A = \sum_{n=0}^{\infty} \frac{C_n}{n!} \beta^n$$

$$C_0 = \beta = \gamma^{\mu}$$

$$C_1 = [A, \beta] = \gamma_2 \omega_{\mu} [S^{\mu}, \beta]$$

Computation:

$$C_0 = \beta = \gamma^{\mu}$$

$$C_1 = [A, \beta] = \gamma_2 \omega_{\mu} [S^{\mu}, \beta] = -i\gamma_2 \omega_{\mu} (\gamma^{\mu})^{\nu} \gamma_{\nu} = (-i\gamma_2 \omega_{\mu} \gamma^{\mu})^{\nu} \gamma_{\nu}$$

$$C_2 = [A, C_1] = \frac{i}{2} \omega_{\mu} [S^{\mu}, (-i\gamma_2 \omega_{\mu} \gamma^{\mu})^{\nu} \gamma_{\nu}]$$

$$= (-i\gamma_2 \omega_{\mu} \gamma^{\mu})^{\nu} \gamma_{\nu} (-i\gamma_2 \omega_{\mu} \gamma^{\mu})^{\nu} \gamma_{\nu}$$

$$\begin{aligned} &= -i\gamma_2 \omega_{\mu} (\gamma^{\mu})^{\nu} \gamma_{\nu} (-i\gamma_2 \omega_{\mu} \gamma^{\mu})^{\nu} \gamma_{\nu} \\ &= [(-i\gamma_2 \omega_{\mu} \gamma^{\mu})^2]^{\nu} \gamma_{\nu} \end{aligned}$$

$$\begin{aligned} &\text{These means:} \\ &e^{A \beta e^{-A}} = \left( \mathbb{1} + (-i\gamma_2 \omega_{\mu} \gamma^{\mu})^{\nu} \gamma_{\nu} \right)^{\mu} \gamma_{\mu} \\ &= \exp(-i\gamma_2 \omega_{\mu} \gamma^{\mu})^{\mu} \gamma_{\mu} \end{aligned}$$